

1 **Highly siderophile and chalcophile element behaviour in abyssal-type**
2 **and supra-subduction zone mantle: new insights from the New**
3 **Caledonia ophiolite**

4

5 Arianna Secchiari ^{1,2}, Philipp Gleissner¹, Chunhui Li³, Alexey Goncharov⁴, Harry
6 Becker¹, Delphine Bosch⁵, Alessandra Montanini²

7

8 ¹ Institut für Geologische Wissenschaften, Freie Universität Berlin, Germany,

9 ² Department of Chemistry, Life Sciences and Environmental Sustainability, University
10 of Parma, Italy,

11 ³ Department of Geophysics and Planetary Sciences, University of Science and
12 Technology of China, Hefei, China,

13 ⁴ Department of Geophysics, Saint Petersburg State University

14 ⁵ Géosciences Montpellier, Université de Montpellier, France

15

16 **Abstract**

17

18 The New Caledonia Ophiolite hosts one of the largest obducted mantle sections
19 worldwide, offering a unique opportunity to investigate key mantle processes. The
20 ophiolite comprises refractory harzburgites, locally overlain by mafic-ultramafic
21 cumulates, and minor lherzolites. Previous geochemical studies indicated that the
22 lherzolites are akin to abyssal-type peridotites, while the harzburgites underwent

23 multiple melting episodes in MOR and supra-subduction zone environments, followed
24 by late stage metasomatism.

25 In this work, Os isotopes, highly siderophile (HSE) and chalcophile element data are
26 reported for the New Caledonia peridotites, in order to constrain the behaviour of
27 these elements in abyssal-type and fore-arc mantle.

28 The variably serpentinised lherzolites (LOI = 6.4 - 10.7 %) yield slightly subchondritic
29 to suprachondritic initial Os isotopic compositions ($^{187}\text{Os}/^{188}\text{Os}_i = 0.1273\text{-}0.1329$) and
30 subchondritic to chondritic Re/Os ratios (0.04-0.11). The gently sloping HSE patterns
31 with increasing depletion towards Au show concentrations in the range of other
32 lherzolites from MOR or continental setting. Sulphur contents are high and variable
33 (202-1268 ppm), and were likely increased during serpentinisation. By contrast,
34 Se/Te ratios and concentrations are within the range of primitive mantle (PM) values,
35 meaning that these elements were not significantly mobilised during serpentinisation.

36 Although displaying homogenous petrographic and geochemical features, the
37 harzburgites are characterised by extremely heterogeneous Re-Os and HSE
38 compositions.

39 Type-A harzburgites exhibit subchondritic $^{187}\text{Os}/^{188}\text{Os}_i$ (0.1203-0.1266) and low Re/Os
40 ratios (0.01-0.04). The strong IPGE-PPGE fractionations ($\text{Pd}_N/\text{Ir}_N = 0.21\text{-}0.56$),
41 coupled with positive Pt anomalies and S-Se-Te abundances often below the detection
42 limit, suggest high melt extraction rates, resulting in sulphide consumption and Os-Ru
43 metal alloy stabilisation.

44 Type-B harzburgites possess strongly fractionated, Os-Ir-Pt poor (Os = 0.003-0.072
45 ng/g, Ir = 0.0015-0.079 ng/g) and Pd-Re enriched patterns, associated with
46 chondritic to suprachondritic measured $^{187}\text{Os}/^{188}\text{Os}$ (0.127-0.153). These characters
47 are uncommon for highly depleted mantle residues. Interaction with an oxidised
48 component does not appear as a viable mechanism to account for the IPGE-depleted

49 patterns of type-B harzburgites, as calculated oxygen fugacities are close to the FMQ
50 buffer ($\text{Log } \Delta\text{FMQ} = 0.35 \text{ to } 0.65$).

51 The strikingly uniform mineralogical and geochemical features displayed by both
52 harzburgite sub-types suggest that the different HSE patterns are not linked to their
53 recent evolution, implying that subduction-related processes were superimposed on
54 geochemical heterogeneous mantle domains, which exerted an important control on
55 HSE behaviour during melt extraction and post melting metasomatism.

56 We propose that the HSE characters of the studied peridotites reflect the presence of
57 a highly heterogeneous mantle source with a long term ($> 1 \text{ Ga}$) evolution, possibly
58 linked to the Zealandia formation.

59

60 **1. Introduction**

61

62 Highly siderophile elements (HSE: PGE+ Au-Re) are powerful geochemical tracers that
63 can provide useful information for a variety of mantle processes, such as mantle
64 melting, metasomatism and melt-fluid/mantle interaction (e.g. Luguet et al., 2001,
65 2003, 2007; Lorand et al., 2008; Ackerman et al., 2009). However, our knowledge
66 concerning the behaviour of HSE in mantle source rocks of primitive arc magmas and
67 the role of the subduction zone environment on HSE partitioning (i.e. hydrous melting,
68 melt/fluid-mantle interaction) still remains quite fragmentary. Furthermore, although
69 abundant HSE data are now available for different types of mantle peridotites, HSE
70 data on fore-arc peridotites are remarkably scarce (e.g. Becker and Dale, 2016).

71 The New Caledonia ophiolite (Peridotite Nappe) hosts one of the largest and best
72 preserved mantle sections worldwide, providing an excellent opportunity to
73 investigate upper mantle processes. The rock exposures are dominated by harzburgite
74 tectonites bearing a supra-subduction zone affinity (Marchesi et al., 2009; Ulrich et

75 al., 2010; Pirard et al., 2013; Secchiari et al., 2020). The main geochemical and
76 isotopic features of these rock-types reflect a complex polyphase evolution, including
77 several melting episodes in different geodynamic settings and subduction zone
78 metasomatism (Marchesi et al., 2009; Ulrich et al., 2010; Secchiari et al., 2020).
79 Minor spinel and plagioclase lherzolites, with compositions similar to abyssal
80 peridotites, occur as discrete bodies in the north-western part of the island. The
81 lherzolites record a different history compared to the extremely refractory
82 harzburgites, as highlighted by their different geochemical signature (Ulrich et al.,
83 2010; Secchiari et al., 2016).

84 In this work, a set of fully characterised peridotites (i.e. whole rock and in situ major
85 and trace element contents, Sr-Nd-Pb isotopes) from New Caledonia (Secchiari et al.,
86 2016, 2019, 2020) has been used to investigate Re-Os, HSE and chalcophile element
87 (S-Se-Te) systematics. The main aims of this work are: 1) to examine the behaviour
88 of these elements in the lherzolites (i.e. presumed abyssal peridotites) and in the
89 ultra-depleted harzburgites, which may represent rocks from a former supra-
90 subduction zone mantle wedge; 2) to constrain the behaviour of HSE and chalcophile
91 elements during subduction zone processes.

92

93 **2. Geological setting and petrological background**

94

95 New Caledonia is a NW–SE elongated island located in the SW Pacific region, between
96 the eastern margin of Australia and the Vanuatu archipelago (Fig. 1a). The island
97 represents the emerged portion of the submarine Norfolk Ridge and it is composed by
98 a mosaic of volcanic, sedimentary and metamorphic terranes, ranging in age from
99 Permian to Miocene (Aitchison et al., 1995; Cluzel et al., 2001, 2012; Lagabrielle et
100 al., 2013). These terranes were amalgamated during two major tectonic events: 1) an

101 Early Cretaceous tectonic convergence phase (Paris, 1981) and 2) a Paleocene to Late
102 Eocene subduction culminated in the obduction of the ophiolite. Both events were
103 characterised by high-pressure low-temperature (HP-LT) metamorphism in connection
104 with plate convergence. New Caledonia can be sub-divided into four main geological
105 domains (Cluzel et al., 2001; see Fig. 1): (i) the Basement units (pre-Late Cretaceous
106 basement and Late Coniacian-to-Late Eocene sedimentary cover), (ii) the Cenozoic
107 HP-LT metamorphic belt, (iii) the basaltic Poya Terrane and (iv) a large slab of
108 peridotites, named the Peridotite Nappe.

109 The Peridotite Nappe represents an allochthonous sheet of oceanic lithosphere
110 belonging to the former South Loyalty basin thrust on the continental basement of the
111 Norfolk Ridge at the end of the Eocene subduction. The emplacement of the ophiolitic
112 nappe resulted from the failed subduction of the Norfolk Ridge tip in a NE-dipping
113 subduction zone, which culminated in the obduction of the Loyalty subarc lithosphere
114 ~ 34 Ma ago (Cluzel et al. 2012).

115 The Peridotite Nappe has an extension of about 8000 km² and is mostly exposed in
116 the Massif du Sud, where a thick harzburgite–dunite unit, locally overlain by
117 kilometre-scale lenses of mafic and ultramafic intrusives, crops out. The sequence is
118 believed to represent a crust-mantle boundary that records the onset of the Eocene
119 subduction in a nascent arc setting (Marchesi et al., 2009; Pirard et al., 2013;
120 Secchiari et al., 2018). Recent geochemical studies have shown that the ultramafic
121 intrusives (i.e. dunites and wehrlites) crystallised from variably depleted melts with
122 island arc basalt affinity, after massive interactions with the underlying harzburgite
123 (Marchesi et al., 2009; Pirard et al., 2013). In contrast, the mafic rocks (i.e.
124 gabbro-norites) have a cumulate origin (Marchesi et al., 2009; Pirard et al., 2013;
125 Secchiari et al., 2018) and derive from crystallisation of primitive, non-aggregated,
126 ultra-depleted melts showing involvement of a subduction-related component in their
127 source (Secchiari et al., 2018).

128 The harzburgites are also exposed in the northern Tiébaghi massif (Ulrich et al., 2010)
129 or as sparse tectonic klippen in the central part of the island (e.g. Kopeto, Poya,
130 Koniambo), where exceptionally fresh peridotites display primary mineral assemblages
131 similar to the more serpentinitised rocks of the Massif du Sud.

132 The New Caledonia harzburgites bear an overall ultra-depleted composition, inherited
133 from a complex multistage evolution linked to the development of the Eocene
134 subduction system (Marchesi et al., 2009; Ulrich et al., 2010; Secchiari et al., 2020).

135 Geochemical studies have proposed that the harzburgites formed by high degrees of
136 fluid-assisted melting (up to 20-25 % in a supra-subduction zone environment, see
137 Marchesi et al., 2009; Ulrich et al., 2010). More recently, the work of Secchiari et al.
138 (2020) provided further constraints on the evolution of these rock-types, tracking
139 their history from melting to late stage metasomatism. Accordingly, the harzburgites
140 underwent two partial melting episodes in the spinel stability field: a first melting
141 phase in a MOR setting (15% melting degrees), followed by hydrous melting in a
142 supra-subduction zone setting (up to 18% fluid-assisted melting). Post-melting
143 cooling and re-equilibration at lithospheric conditions was accompanied by interaction
144 with ultra-depleted slab-derived hydrous melts (Secchiari et al., 2019, 2020). These
145 metasomatic processes are testified in the harzburgites by the widespread occurrence
146 of secondary metasomatic phases (i.e. thin films of Al_2O_3 -poor orthopyroxene,
147 and low Al_2O_3 and Na_2O clinopyroxene), L-MREE and Zr-Hf bulk rock enrichments, and
148 possibly by the poorly radiogenic Nd isotopic ratios shown by some samples (Secchiari
149 et al., 2020).

150 Compared to the harzburgites from the central and the southern massifs, Tiébaghi
151 samples display a more fertile nature, as revealed by the higher bulk trace element
152 concentrations as well as by the occurrence of a small fraction (up to 4 vol.%) of
153 clinopyroxene (see Ulrich et al., 2010; Secchiari, PhD thesis).

154 The main geochemical and petrological features of the spinel and plagioclase
155 lherzolites are thought to reflect moderate melt extraction degrees (8-9%) in a MOR
156 environment, followed by refertilisation by depleted MORB-type melts, yielding
157 plagioclase lherzolites. The main petrological and geochemical features of the
158 lherzolites have been reported in detail by Secchiari et al. (2016).

159

160 *2.1 Sample description*

161

162 In this contribution, sixteen samples of peridotites fully characterised for lithophile
163 element geochemistry (i.e. major, trace element and Sr-Nd-Pb isotope compositions)
164 were analysed for mass fractions of all PGE, Re, Au, S, Se, Te and $^{187}\text{Os}/^{188}\text{Os}$.
165 Detailed descriptions of the lherzolites and harzburgites, including trace element
166 chemistry and Sr-Nd-Pb isotopes, are provided in the works of Secchiari et al. (2016)
167 and Secchiari et al. (2019, 2020), respectively.

168 Lherzolite samples are from the Poum and Babouillat areas, while the harzburgites
169 were collected from several outcrops and mine zones along the island: Yaté, Kopeto,
170 Poya, Poro and Tiébaghi (Fig. 1b and Table 1). The lherzolites include slightly
171 serpentinised (LOI= 6.9 - 10.7 %) spinel and plagioclase lherzolites, while the
172 harzburgites are typically not or only weakly serpentinised (LOI = 0 - 3 %), except
173 for samples YA1, TI1 and TI2 (LOI = 6.0 - 9.0 %).

174 Both lherzolites and harzburgites are low strain mantle tectonites, showing dominant
175 porphyroclastic textures (Fig. S1a-b) and local protomylonite development. Spinel
176 lherzolites have 7-8 vol. % clinopyroxene and display a typical abyssal-type REE
177 signature. The plagioclase lherzolites show melt impregnation microstructures (Fig.
178 S1b) and are slightly enriched in incompatible trace element (REE, Ti, Y, and Zr) with
179 respect to the spinel lherzolites. Harzburgites are extremely depleted rocks, as

180 highlighted by the general absence of clinopyroxene (with the exception of sample
181 TI2, where clinopyroxene is ~ 4 vol. %, Fig. S1c) and remarkably low incompatible
182 trace element contents (Secchiari et al., 2020). The primary mantle assemblage is
183 composed of olivine, orthopyroxene and spinel. The occurrence of thin films of
184 metasomatic ortho- and clinopyroxene (Fig. S1d) was interpreted as resulting from
185 percolation of small fractions of subduction-related magmas (Secchiari et al., 2019,
186 2020).

187 All the lherzolites studied here contain trace amounts of sulphides with variable size,
188 shape and position. Sulphides in five selected samples (i.e. BA1, POU2, POU3, POU1A,
189 BAB1B, see Table S1) were analysed for their major element chemistry composition.

190 Frequent interstitial sulphide grains occur as polyhedral blebs, rounded or irregular-
191 shaped grains (Fig. S2a-d-e), generally located at olivine-orthopyroxene or at olivine-
192 olivine grain boundaries. They range from tiny crystals ($\sim 60 \times 30 \mu\text{m}$) up to $300 \times$
193 $100/150 \mu\text{m}$ in maximum dimensions. Enclosed sulphides hosted within large olivine
194 porphyroclasts have been also observed. Sulphide inclusions (Fig. S2a-e) have
195 variable size (30 to $80 \mu\text{m} \times 60$ to $150 \mu\text{m}$) and shape, from polyhedral to spherical.
196 Very tiny enclosed sulphide grains ($\sim 5\text{-}6 \mu\text{m} \times 10 \mu\text{m}$) have been recognised in
197 sample POU1A.

198 Major element composition of the sulphide phases is relatively homogeneous (Table
199 S1 and Fig. S2, S3), showing no significant difference among enclosed and interstitial
200 sulphides. Chemical composition of the studied sulphides mainly fall in the field of Ni-
201 poor ($\text{Fe}/\text{Ni}=1.3\text{-}2.1$) monosulphide solid solution + Liquid (Fig. S3), but Ni-rich
202 ($\text{Fe}/\text{Ni} = 0.7\text{-}0.8$) pentlandite, resulting from cooling and re-equilibration (Guo et al.,
203 1999), has been also identified in four of the investigated samples (POU2, POU3,
204 POU1A, BAB1B).

205 In the harzburgites, sulphide phases have not been identified, nor through
206 petrographic observation or microprobe analyses. This is consistent with the refractory
207 nature and the high melt extraction rates experienced by these rock-types, which
208 prevented sulphide retention in the residual mantle source.

209

210 **3. Analytical methods**

211

212 *3.1 HSE and chalcophile elements*

213

214 Seven lherzolites and nine harzburgites (including four duplicates) have been analysed
215 in the geochemistry laboratory at Freie Universität for HSE, S, Se, Te mass fractions
216 in whole rocks and $^{187}\text{Os}/^{188}\text{Os}$.

217 Detailed procedure descriptions have been given in previous works from this
218 laboratory (e.g. Fischer-Gödde et al., 2011; Wang et al., 2013; Wang and Becker,
219 2013). The methods will only be briefly summarised below.

220 About 2.5 g of sample powder was weighed into 90 ml quartz glass digestion vessels
221 and spiked with mixed ^{191}Ir - ^{99}Ru - ^{194}Pt - ^{105}Pd , ^{77}Se - ^{125}Te , ^{185}Re - ^{190}Os and ^{34}S solutions.
222 Then, 5 ml 14 mol/L, N_2 -bubbled HNO_3 and 2.5 ml 9 mol/L HCl were added. The
223 vessels were immediately sealed with Teflon tape and samples were digested for 16 h
224 at 320°C and 100 bar. After digestion, osmium was extracted from the reverse aqua
225 regia into chloroform, back extracted into HBr (Cohen and Waters, 1996), and further
226 purified by micro distillation from a H_2SO_4 -dichromate solution into 15 μl of HBr (Birck
227 et al., 1997)

228 Osmium isotopes were determined as OsO_3^- in negative mode using the Thermo
229 Finnigan Triton TIMS, using a secondary electron multiplier. Signal intensities of the

230 spike isotope ^{190}Os of samples were $\sim 150,000\text{--}500,000$ cps. Standard runs with
231 different amounts of Os on the filament (10 pg and 100 pg) were also run in between
232 the studied samples, yielding an average value of 0.1139 ± 0.0002 (2 s. d., $n = 24$)
233 for 100 pg loads. Two hundreds scans were collected in each measurement for high-
234 Os samples, while at least 120-140 scans were obtained for the low-Os samples. Raw
235 data were corrected for isobaric OsO_3^- interferences, mass fractionation using the
236 $^{192}\text{Os}/^{188}\text{Os}$ ratio of 3.08271, contributions from the Os spike solution and blank
237 contributions. $^{187}\text{Os}/^{188}\text{Os}$ were finally adjusted relative to the mean of the Os
238 standard. The oxygen isotope compositions used for the oxide correction of Os oxide
239 molecules were $^{18}\text{O}/^{16}\text{O}$ of 0.00204 and $^{17}\text{O}/^{16}\text{O}$ of 0.00037 (Nier, 1950).

240 About 50% of the digestion solution was used for separation of the HSE fraction and
241 about 30% for S–Se–Te separation. Chemical separation of the HSE fraction from the
242 matrix was performed on columns filled with 10 ml of pre-cleaned Eichrom 50W-X8
243 (100–200 mesh) cation exchange resin (Fischer-Gödde et al., 2011). During
244 separation, the HSE fraction was collected in 14 ml 0.5 mol/L HCl-40 vol.% acetone
245 mixture. After the volume of the solution has been reduced to about 2 ml it was
246 analysed for Au, Re, Ir and Pt. In order to remove interfering Cd, the remaining
247 solution was further purified in 0.2 mol/L HCl on 3 ml Eichrom 50W-X8 (100–200
248 mesh) resin. The collected solution was evaporated to near dryness and the residue
249 was taken up in 0.28 M HNO_3 for ICP-MS analysis. The analyses were carried out
250 using a single collector Element XR instrument. We used either a Scott-type spray
251 chamber (Re, Ir, Pt, Au) or an Aridus-I desolvator (Ir, Ru, Pt, Rh, Pd) at an oxide
252 formation rate of $\text{CeO}^+/\text{Ce}^+ < 0.004$.

253 A two-step ion exchange chromatography method was used for separation of S, Se
254 and Te (see Wang et al., 2013). Sulphur measurements were performed on the S–Se
255 fraction at medium mass resolution mode on the Element XR. Selenium and Tellurium
256 were measured using a double pass Scott type glass spray chamber at low mass

257 resolution mode on the Element XR, combined with a hydride generation sample
258 introduction system by reacting the sample solution with 1 g/100 g NaBH₄ in 0.05
259 mol/L NaOH (see Wang et al., 2013 for details).

260 For each batch of analysis, one procedural blank has been used. Procedural blanks
261 yielded the following mean values (\pm 1 s.d., n = 4-5): Re = 2.5 \pm 2.0 pg; Os = 0.5 \pm
262 0.3 pg with ¹⁸⁷Os/¹⁸⁸Os ratios of 0.14 \pm 0.03; Ir = 15 \pm 5 pg; Ru = 45 \pm 14 pg; Rh =
263 24 \pm 22 pg; Pt = 23 \pm 29 pg; Pd = 640 \pm 330 pg; Au = 4 \pm 2 pg; Te = 1.1 \pm 0.8 ng;
264 Se = 2.3 \pm 0.8 ng; S = 2.8 \pm 0.7 μ g. Samples were corrected for total procedural
265 blanks using the mean values. Blank corrections for Re are negligible for most of the
266 analysed samples (\leq 0.3 - 0.8 %), but more significant for the harzburgites (\sim 4 - 8
267 %). Blank corrections for Pt and Pd are again negligible for the lherzolites (\sim 0.2 - 0.3
268 %), a few percents for the harzburgites (\sim 0.4 - 4 %, with the exception of KPT2,
269 KPT5 and PO3 for which the correction for Pt is \sim 11 - 36 %). Blanks of Os, Ir, Ru and
270 Rh are insignificant for most of the samples (\leq 0.4 %) but higher for the most
271 depleted harzburgites, i.e. KPT2, KPT5 (\sim 2 - 7 % for Os, Ir, Rh) and PO3 (\sim 9 % for
272 Os and Ir, 13 % for Rh). Blank corrections for S, Se and Te in lherzolites range
273 between 1 - 1.7 % (S - Se) and 3-6% (Te), while for the harzburgites blank
274 corrections for these elements strongly affected the obtained results (corrections \sim 10
275 - 26 % for S and up to 40 - 80% for Se and Te), given the very low measured
276 abundances.

277

278 *3.2 Oxygen fugacity*

279

280 Iron oxidation state in studied spinels was measured using the "flank method"
281 developed for the JEOL JXA-8200 electron microprobe at Freie Universität (Goncharov,
282 2018). The position shift and intensity variation of FeL α and FeL β x-ray lines were

283 investigated using as standards a collection of nineteen natural mantle spinels
284 previously characterised by Mössbauer spectroscopy for their $\text{Fe}^{3+}/\Sigma\text{Fe}$ at IPGG RAS
285 (St. Petersburg, see Goncharov & Ionov, 2012; Goncharov et al., 2015).

286 The analytical procedure was similar to the experiments performed over the last
287 decades to study $\text{Fe}^{3+}/\Sigma\text{Fe}$ in mantle garnets after the “flank” approach developed by
288 Höfer and Brey (2007).

289 The flank positions for the spectrometer were obtained from the difference spectrum
290 of almandine and andradite with known iron oxidation state in the wavelength range
291 related to $\text{FeL}\alpha$ and $\text{FeL}\beta$ lines. The intensities at flank positions near $\text{FeL}\alpha$ and $\text{FeL}\beta$
292 lines were collected as two fake elements, setting counting time at 300 s on both
293 measurements, with 3 repetitions in the core and in the rim respectively, and
294 considering 5 separate spinel grains within one thin section. Measurement conditions
295 were 15 kV and 60 nA using TAP crystal for intensities at flank positions and with the
296 remaining four spectrometers measuring chemical composition at the same spot
297 simultaneously. The studied spinels show no significant core to rim zoning in terms
298 of FeO content, Cr# and $\text{FeL}\beta/\text{FeL}\alpha$ ratio. Averaged $\text{FeL}\beta/\text{FeL}\alpha$ ratios for each sample
299 were used to calculate iron oxidation state of spinel using the equation with
300 coefficients obtained after investigation of the standard collection (see Table 2).

301

302 **4. Results**

303

304 *4.1 HSE and chalcophile elements in spinel and plagioclase lherzolites*

305

306 HSE and chalcophile element compositions of the New Caledonia lherzolites are
307 reported in Table 1 and displayed in Fig. 2, 3, 4. Spinel and plagioclase lherzolites are

308 relatively homogeneous in terms of HSE, Se, Te concentrations, abundance patterns
309 and Os isotopic compositions, with plagioclase-bearing samples showing
310 indistinguishable patterns from those of spinel lherzolites. The concentrations of the
311 highly siderophile and chalcophile elements are in the range of those observed for
312 modern abyssal and ophiolitic peridotites, displaying good correlation for Ir group PGE
313 (IPGE, e.g., Os vs. Ir and Ir vs. Ru) and more dispersed variations for the Pt group
314 PGE (PPGE, Fig. S4). In primitive mantle (PM) normalised concentration diagrams
315 (Fig. 3), the lherzolites exhibit flat or gently sloping negative patterns with similar PM-
316 normalised PGE concentrations and depletions in Au (except for sample BAB2B) and
317 Re compared to the PGE (with the exception of POU1A and POU3 for Re). Overall,
318 absolute contents of the PGE are similar or slightly lower than primitive mantle (PM)
319 values (Becker et al., 2006; Fischer-Gödde et al., 2011), overlapping the field of the
320 abyssal peridotites and peridotite tectonites from continental settings (e.g., Fig. 2 and
321 Becker and Dale, 2016). Ru/Ir and Pd/Ir ratios are suprachondritic, as observed for
322 other mantle lherzolites (e.g. Lorand et al., 1999; Rehkämper et al., 1999; Luguet et
323 al., 2003; Becker et al., 2006).

324 Initial $^{187}\text{Os}/^{188}\text{Os}$ ratios calculated at 53 Ma (i.e. the inferred age of initial magmatism
325 in the subduction system, e.g. Cluzel et al., 2006) vary from chondritic to slightly
326 suprachondritic (0.1273-0.1329, Fig. 4a), corresponding to $\gamma_{\text{Os}(53\text{Ma})}$ of 0.5 to 4.9.
327 These values overlap with data of abyssal peridotites and orogenic peridotites, but
328 tend to be somewhat higher than for other mantle lherzolites bearing comparable
329 depletion degrees (Fig. 4a and 4b). $^{187}\text{Re}/^{188}\text{Os}$ ranges from subchondritic to slightly
330 suprachondritic values (0.186-0.525, see Fig. 4c).

331 Se and Te are positively correlated in the lherzolites (Fig. 5a) and range between 54 -
332 91.3 ng/g and 7.3 - 13.8 ng/g, respectively. Se/Te ratios (5.9-7.1) are slightly lower
333 than the PM value and similar to the data previously obtained on depleted lherzolites

334 (Wang and Becker, 2013). Se and Te do not display any correlation with PGE
335 abundances, with the exception of Te, which shows a weak correlation with Pd (Fig.
336 5b). S contents are high and variable (202 - 1268 $\mu\text{g/g}$) compared to unserpentinised
337 peridotites, leading to high S/Se ratios (2703-16289, Fig 5c).

338

339 *4.2 HSE and chalcophile elements in harzburgites*

340

341 On the basis of HSE behaviour and Os isotopic compositions (Fig. 2, 3 and 6, Table 1),
342 the studied harzburgites can be grouped into two different sub-types: type-A and
343 type-B.

344 *Type-A harzburgites* (TI1, TI2, PO4 and YA1) are characterised by notably lower PGE
345 and chalcophile elements mass fractions (Figs. 2, 4) and more fractionated patterns
346 (Fig. 6a) compared to the lherzolites. Mass fractions of the PGE are 1.07-2.07 ng/g for
347 Os, 0.50-1.14 ng/g for Ir, 1.53-2.52 ng/g for Ru, 0.19-0.27 ng/g for Rh, 2.42-2.70
348 ng/g for Pt and 0.49-0.57 ng/g for Pd. Among the sub-group A, sample YA1 displays
349 distinct PGE abundances, showing much lower Os, Ir, Pt, Pd contents (0.55 ppb for
350 Os, 0.30 ppb for Ir, 0.27 ppb for Pt and 0.26 ppb for Pd), with exceptions for Ru, Rh,
351 Au and S.

352 HSE and chalcophile element diagrams of type-A harzburgites display fractionated
353 patterns, with concentrations decreasing from Os towards Re. Os and Ru are enriched
354 compared to Ir, leading to correlated suprachondritic Os/Ir and Ru/Ir ratios
355 (Os/Ir=2.9-7.9; Ru/Ir=1.8-2.1, Fig. 4). Pt and Au generally show positive spikes,
356 more pronounced for Au, with the exception of YA1, which displays a negative Pt
357 anomaly. Pd contents are low (< 0.1 PM values) and constant for PO4, TI1, TI2, with
358 Pd/Ir showing subchondritic ratios for all the studied samples. Positive correlations are
359 observed between IPGE (Fig. 2a-b-c) and Pt-Ir (not shown), and, somewhat

360 surprisingly, between IPGE and the fertility indicators (i.e. Al₂O₃ and CaO, not shown).
361 Mass fractions of Te, Se and S are low, often close to or below the detection limit,
362 again with the exception of the harzburgite YA1.

363 For all type-A harzburgites, Re concentrations are very low (about 0.02 ng/g), leading
364 to subchondritic ¹⁸⁷Re/¹⁸⁸Os ratios (0.045 to 0.196, the latter value also reflecting the
365 low Os concentrations in YA1). Os isotopic compositions are subchondritic to chondritic
366 (0.1203-0.1266, corresponding $-5 \leq \gamma_{\text{Os}(53 \text{ Ma})} \leq -0.1$) and do not define any
367 correlation with ¹⁸⁷Re/¹⁸⁸Os or incompatible element depletion indices (i.e. Al₂O₃, see
368 Fig. 4).

369
370 *Type-B harzburgites* comprise very fresh samples from Kopeto (KPT2, KPT3, KPT5),
371 Poro (PO3) and Poya (PY1) massifs. Compared to type-A harzburgites, these samples
372 have much lower HSE abundances and display variable and strong fractionations
373 among PGE and more incompatible chalcophile elements (Fig. 2 and 6b). In detail, Os,
374 Ir and Pt show positive correlations (Fig. 2) and are strongly depleted compared to
375 Ru, Rh and Pd (Os= 0.003-0.071 ppb, Ir= 0.015-0.079 ppb, with Os/Ru=0.01-0.26
376 and Ru/Ir= 2.5-20). For sample PY1, Pt is enriched relative to IPGE, Rh and Pd
377 (Pt/Rh=7.7; Pt/Pd=5.5). Pd, Re and S-Se-Te have similarly low normalised
378 abundances, with chalcophile elements often close to or below the detection limit. Au
379 exhibits positive spikes for PY1, KPT3 and PO3 samples and tends to be more enriched
380 than similar incompatible chalcophile elements (i.e. Pd and Re).

381 Measured ¹⁸⁷Os/¹⁸⁸Os ratios vary from chondritic to suprachondritic (0.1273-0.1534)
382 and are coupled with high and variable ¹⁸⁷Re/¹⁸⁸Os (1.62-32). Initial Os isotopic
383 compositions calculated at 53 Ma range from depleted to slightly suprachondritic
384 values (0.1181-0.1365, $-7 \leq \gamma_{\text{Os}(53\text{Ma})} \leq 3$).

385 Replicate analyses of samples PY1, KPT2, KPT5, PO3 yield quite similar results for Ru
386 (< 5 % relative deviation, except for sample KPT2) and Au (~ 6% for PY1 and KPT5)
387 and acceptable results for Pt for PY1-KPT5 (11.5 - 14.0 % relative deviation). Values
388 appear much more scattered for Os and Re (RSD> 30%) and less dispersed for Rh
389 and Pd ($7 \leq \text{RSD}\% \leq 32$). The relative deviation of chalcophile elements is more limited,
390 mostly < 15%.

391 The large variations of mass fractions of HSE and chalcophile elements in duplicate
392 samples reflects the very low mass fractions of these elements combined with the
393 inhomogeneity in the distribution of HSE carrier phases in gram-size quantities of rock
394 powder, an issue that has already been recognised in peridotitic rocks (e.g. Becker et
395 al., 2006; Luguet et al., 2007).

396

397 **4.3 Oxybarometry**

398

399 Results for calculations of f_{O_2} using coexisting olivine, orthopyroxene and spinel (e.g.,
400 Wood, 1990, see Table 2 and Table S2) with the reaction



402 (1)

403 are given in Table 2.

404 The calculation of f_{O_2} depends on the pressure and temperature of equilibrium. For
405 this study, we used olivine-spinel geothermometry (Li et al., 1995) as this
406 thermometer is based on the same elements and minerals (Fe and Mg content in
407 olivine and spinel) that are used to calculate f_{O_2} . As spinel peridotites lack a good
408 barometer, in order to maximise consistency with temperatures determined with Li et
409 al. (1995) thermometer, we have assumed a pressure of 1.0 GPa.

410 Ferric to total iron ratios ($\text{Fe}^{3+}/\Sigma\text{Fe}$) in spinel are similar within the harzburgites,
411 covering a limited range of values ($0.18 \leq \text{Fe}^{3+}/\text{Fe}^{3+} + \text{Fe}^{2+} \leq 0.21$, see Table 2) indicating
412 similar f_{O_2} values for the two harzburgite sub-groups ($\Delta\log\text{FMQ} = 0.35\text{-}0.92$).

413 In Fig. 7a-b spinel Cr# is reported against $\text{Fe}^{3+}/\Sigma\text{Fe}$ ratio and $\Delta\log\text{FMQ}$ for the
414 analysed samples. Literature data for mantle peridotites from different geodynamic
415 settings (i.e. MOR and supra-subduction zone) and arc xenoliths have also been
416 plotted for comparison. Overall, the harzburgites share remarkable similarities with
417 supra-subduction zone peridotites from the Izu-Bonin area (Parkinson and Pearce,
418 1998). Slightly lower Cr# values compared to the other harzburgites are displayed by
419 TI2 sample, which in turn falls between abyssal peridotites and arc xenoliths array
420 (Fig. 7a).

421

422 **5. Discussion**

423

424 *5.1 HSE and Re-Os systematics of the lherzolites*

425

426 Major element composition and lithophile trace element chemistry of spinel lherzolites
427 indicate moderately depleted compositions, inherited from moderate partial melting
428 degrees (8-9%) of a DMM source, whereas plagioclase lherzolites originated through
429 reactive melt percolation of spinel lherzolites by highly depleted, incremental melt
430 fractions of a DMM source in the shallow lithosphere (Secchiari et al., 2016). In the
431 following sections, the processes that may have affected HSE and Os isotopic
432 signature of the New Caledonia lherzolites will be discussed: low temperature
433 alteration, in particular serpentinisation, partial melting and the role of melt infiltration
434 and chemical disequilibrium of the HSE in mantle rocks.

435

436 *5.1.1 Effects of serpentinisation on HSE and $^{187}\text{Os}/^{188}\text{Os}$*

437

438 Serpentinisation is a widespread process of hydrothermal alteration in ultramafic
439 lithologies. However, its influence on HSE behaviour has not been investigated with
440 much detail, despite some authors have proposed it as a possible cause for ^{187}Os
441 ingrowth and Re variations in the upper mantle (Snow and Reisberg, 1995; Walker et
442 al., 1996; Standish et al., 2002). Recent experimental studies have shown that during
443 serpentinisation the formation of secondary sulphides, Fe-Ni alloys and native metals
444 (Au-Cu) is promoted by reducing f_{O_2} conditions (Klein and Bach, 2009; Foustoukos et
445 al., 2015) and thus, with the exception of Au, HSE may be retained in the host rock.
446 Comparison of partially serpentinised and unserpentinised peridotites displaying
447 similar major element features supports the notion that at least PGE ratios are not
448 significantly affected by moderate to strong serpentinisation (Becker and Dale, 2016).
449 The New Caledonia Iherzolites underwent intermediate serpentinisation degrees (LOI
450 = 6.4 to 10.7 %, see paragraph 2.1 and Table 1), which had limited effects on the
451 budget of fluid immobile moderately incompatible lithophile trace elements in these
452 rocks (Secchiari et al., 2016). Notably, PGE contents and ratios in the Iherzolites are
453 similar to other unaltered and serpentinised Iherzolites from the modern oceans and
454 ophiolitic complexes (see Fig. 2 and 3 e.g.; Snow et al., 2000; Luguet et al., 2001,
455 2004 ; Pearson et al., 2004; Alard et al., 2005; Becker et al., 2006; Fischer-Gödde et
456 al., 2011; Becker and Dale, 2016). This observation supports the hypothesis that PGE
457 abundances are comparable in fresh and variably serpentinised ultramafic rocks
458 (Becker et al., 2006; Liu et al., 2009; Fischer-Gödde et al., 2011; Marchesi et al.,
459 2013; Becker and Dale, 2016), implying that serpentinisation results in minor changes
460 in PGE ratios.

461 By contrast, the possible influence of serpentinisation on Au and Re is more difficult to
462 evaluate, as no study has systematically investigated its effect on the behaviour of the
463 aforementioned elements. In the lherzolites from New Caledonia, Au displays similar
464 normalised concentrations as Re and, with the exception of a few samples, both
465 elements are depleted relative to Pd, Te and Se. Au abundances tend to be somewhat
466 lower than abundances in other lherzolites with similar major element composition
467 (Fig. 3). Although the compositions can be entirely explained by magmatic
468 fractionation processes (see subsequent chapters), minor losses of Au due to
469 hydrothermal alteration cannot be ruled out (e.g. Lorand et al., 1999). The lack of
470 correlation with Al_2O_3 and Yb_N (not shown) could be a hint that Au abundances may
471 have been affected by a combination of magmatic processes and serpentinisation (i.e.
472 Fischer-Gödde et al., 2011). Rhenium is slightly depleted compared to the PGE for
473 most of the studied lherzolites, but displays higher concentrations than other mantle
474 lherzolites (Fig. 3). In addition, Re contents do not correlate with LOI and Re/Os ratios
475 cover the range generally reported for moderately depleted mantle rocks. The
476 samples with the lowest Re contents display the highest LOI values, suggesting that
477 no significant quantities of Re were added during the interaction with seawater during
478 serpentinisation, as also supported by the chondritic to subchondritic $^{187}\text{Re}/^{188}\text{Os}$ ratios
479 (see Table 1).

480 Likewise, the chondritic to slightly suprachondritic $^{187}\text{Os}/^{188}\text{Os}$ cannot be ascribed to
481 serpentinisation, as unrealistically high water-rock ratios ($\sim 10^3$ - 10^4) would be
482 required in order to perturb the whole rock $^{187}\text{Os}/^{188}\text{Os}$ at the % level or higher (e.g.
483 Becker and Dale, 2016). The lherzolite data also show mass fractions of Se and Te
484 and Se/Te that are similar to values in unserpentinised lherzolites (e.g., Wang and
485 Becker, 2013). In contrast, sulfur in most lherzolites from New Caledonia shows much

486 higher concentrations than typical for peridotites, which is readily explained by
487 contamination with seawater-derived sulfur during serpentinisation.

488 We thus conclude that the HSE (perhaps with the exception of Au), Se, Te and Re-Os
489 signature of the lherzolites offer no evidence to support that serpentinisation and
490 associated reactions affected these elements in a noticeable way.

491

492 *5.1.2 Partial melting and chemical disequilibrium of the HSE in the mantle*

493

494 Spinel and plagioclase lherzolites exhibit comparable HSE contents and patterns,
495 similar to other lherzolites from oceanic or continental settings that underwent low to
496 moderate degrees of melt extraction (Fig. 2 and 3).

497 Partial melting has often been invoked as a possible cause for HSE and $^{187}\text{Os}/^{188}\text{Os}$
498 variations in mantle rocks (e.g. Reisberg and Lorand, 1995; Meisel et al., 2001).

499 Studies of the behaviour of HSE during mantle melting and their abundances in
500 mantle rocks have supported the hypothesis that HSE concentrations in residual

501 peridotites result from sulphide-silicate partitioning during magmatic processes (i.e.
502 Becker and Dale, 2016; Brenan et al., 2016 and references therein). Experimental

503 studies have also highlighted that at temperatures relevant for mantle processes
504 sulphide liquid and, in special cases, sulphide solid solutions, coexist in equilibrium

505 with silicate melt, olivine, pyroxenes and an Al-rich phase (e.g. Rehkämper et al.,
506 1999; Mungall and Brenan, 2014; Brenan et al., 2016). Experimentally determined

507 sulphide melt-silicate melt partition coefficients ($D^{\text{sulph/sil}}$) for PGE have been shown to
508 be high and constant (10^5 to 10^6 , e.g. Mungall and Brenan 2014; Brenan et al. 2016),

509 while Au shows slightly lower $D^{\text{sulph/sil}}$ ($\sim 10^4$). Therefore, up to moderately high
510 degrees of melting, PGE behave as compatible elements and their inter-elemental

511 ratios remain similar as long as sulphide is present in the mantle residue (i.e. until 20-

512 25% of partial melting, depending on initial S content of the source, see Fonseca et
513 al., 2011; Mungall and Brenan, 2014). By contrast, Re is much less chalcophile
514 ($D^{\text{sulph/sil}} \sim 300-800$, e.g. Fonseca et al., 2007; Brenan, 2008) and is expected to
515 become more quickly depleted in the residual mantle.

516 Major element compositions of the sulphide phase observed in the spinel lherzolites
517 (Table S1) are consistent with a residual origin after incongruent melting processes
518 (e.g. Bockrath et al., 2004; Ballhaus et al., 2006). The occurrence of homogeneous
519 monosulphides also suggests relatively high cooling rate after the melting event.

520 Overall, the studied lherzolites are characterised by flat to gently sloping PGE
521 patterns, with similar PM-normalised abundances, no PPGE fractionation and nearly
522 constant ratios for IPGE (i.e. Os/Ir, Ru/Ir). By contrast, Au and Re display the
523 strongest depletion. These features imply that HSE, with the exception of Au and Re,
524 exhibit a similar compatible behaviour during mantle melting, as expected for low to
525 moderate melting degrees in presence of residual sulphide melt. This observation is
526 consistent with the previous estimates obtained through geochemical modelling
527 (Secchiari et al., 2016) and with the occurrence of a residual subsolidus sulphide
528 assemblage in spinel lherzolites.

529 Although the PGE patterns are nearly flat, with only slight depletion of Pd in a few
530 samples, the depletion of Au and Re, the range of chondritic to slightly suprachondritic
531 $\gamma^{187}\text{Os}_i$ and the higher mass fractions of Se and Te compared to Re and the other HSE
532 suggest a multi-stage history of the lherzolites. Notably, $\gamma^{187}\text{Os}_i$ do not correlate with
533 mass fractions of incompatible HSE such as Re, Re/Os nor with fertility indicators (Fig.
534 4), as was observed in some other suites of lherzolites (Becker and Dale, 2016).

535 The Os isotopic signature may be a pre-existing feature of the mantle source, i.e.
536 already present before the recent melt extraction event (Secchiari et al., 2016), as
537 supported by the dispersed Os isotopes-fertility indicators trends. In addition, the

538 remarkable absence of magmatic Cu-Fe-rich sulphides (e.g. see Lorand et al., 2013)
539 argue against a recent, post-melting sulphide addition. We thus speculate that the
540 bulk HSE, Se, Te and Os isotope compositions of the Iherzolites are the result of
541 partial melting event which affected a mantle source previously characterised by an
542 heterogeneous sulphide population including both residual and magmatic sulfides
543 precipitated along grain boundaries by infiltrating melts (Burton et al., 1999; Lorand
544 et al., 1999; Alard et al., 2000, 2002).

545

546 **5.2 Sulphur, Se and Te behaviour in the New Caledonia Iherzolites**

547 Sulfur mass fractions are variable in the Iherzolites from New Caledonia and typically
548 much higher compared to estimates of the depleted MORB mantle source (DMM ~
549 150-200 ppm, Mathez, 1976; Salters and Stracke, 2004) and unaltered Iherzolites
550 (e.g., Wang and Becker, 2013). In addition, sulphur does not correlate with fertility
551 indicators (i.e. Al_2O_3), as commonly observed in unserpentinised mantle tectonites
552 (e.g., Lorand and Alard, 2010; Wang and Becker, 2013). The high S concentrations
553 and the lack of correlation with melting indicators suggest that S experienced a late
554 addition during the evolution of these rocks. Sulphur budget of mantle peridotites can
555 be strongly affected by seawater-rock interaction, because of the high sulphate
556 content of seawater, leading to precipitation of hydrothermal sulphides and sulphate
557 (Alt and Shanks, 1998). We note that the major element chemistry indicates a
558 residual origin for the sulphide phases of the Iherzolites (see paragraph 4.1 and
559 5.1.2). Hydrothermal sulphides or sulphates were not identified during microprobe
560 analyses.

561 Recent geochemical works have demonstrated the role of serpentine as a sink of S
562 under various oxidation states (S^{2-} , S^- , S^0 and S^{6+} , Debret et al., 2017). These
563 studies have shown that S concentrations can be anomalously high in serpentinised

564 peridotites (up to 1 wt.%, see Alt et al., 2003), as S can be accommodated in
565 serpentine minerals, accounting from 60 to 100% of the sulphur budget of the
566 peridotite (Debret et al., 2017). In situ XANES analyses have also revealed that S can
567 be hosted in nano-phases associated with serpentine or trapped either via Si
568 substitutions in the tetrahedra, or as a sulphate ion in the network of the tetrahedral
569 sheet of serpentine minerals (Debret et al., 2017).

570 The addition of S during serpentinisation is also reflected in the high suprachondritic
571 S/Se ratios (up to 16500) and the excellent correlation observed between S
572 concentrations and S/Se ratios (see Fig. 5c). Despite the strong S enrichments, Se
573 and Te display 'normal' concentrations and Se/Te ratios are in the range of other
574 Iherzolites (see Wang and Becker, 2013). These data confirm that Se-Te contents and
575 ratios were not significantly impacted by serpentinisation, as previously observed for
576 other peridotites that experienced low to moderate serpentinisation degrees (e.g.
577 Wang and Becker, 2013; Marchesi et al., 2013).

578 Moreover, Se-Te show a good correlation between each other (Fig. 5a), implying that
579 they are controlled by the same mineral phases. In mantle peridotites, Se can replace
580 S as a chalcogen anion within the crystalline structure of sulphides (e.g. Bulanova et
581 al., 1996; Hattori et al., 2002; Helmy et al., 2010) or can form Se-rich micro phases,
582 while Tellurium, owing to its semi metal nature, tends to partition between sulphides
583 and late exsolved micrometric tellurides (Pt, Pd, Te, As, Bi phases). The latter are
584 thought to crystallise at low temperatures during cooling, once sulphide melt becomes
585 saturated with respect to Te (Luguet et al., 2004; Lorand et al., 2008; Lorand and
586 Alard, 2010). In the Iherzolites, Se and Te do not correlate with melting indicators,
587 but Te displays a rough positive correlation with Pd (Fig. 5b), which suggests that the
588 sulphide melt-bulk silicate partition coefficient of Te should be between Se and Pd
589 (e.g., Figs. 3, 4, 5).

590

591 **5.3 Type-A harzburgites: highly siderophile element systematics of a residual**
592 **sub-arc mantle section**

593

594 The distinct HSE patterns and Os isotopic signature testify that the New Caledonia
595 harzburgites recorded a different evolution compared to the northern Iherzolites.

596 Three of the four samples belonging to the sub-group A (TI1, TI2, PO4) show similar
597 HSE patterns and chalcophile elements depletion, suggesting that the same processes
598 contributed to the HSE and chalcophile element budget of these rocks. The low
599 chalcophile element concentrations, close to or below the detection limit, coupled with
600 low Pd/Ir ratios and subchondritic $^{187}\text{Os}/^{188}\text{Os}_i$, point out that type-A harzburgites are
601 residues after high degree of melt extraction, where sulphides melts must have been
602 nearly completely dissolved in the coexisting silicate melt.

603 Experimental studies have predicted a compatible behaviour for all the PGE during
604 mantle melting as long as the sulphide phase is retained in the peridotite (Mungall and
605 Brenan, 2014). Depending on the initial S content of the mantle rocks, ~17-20% of
606 melting is required for sulphide exhaustion (Fonseca et al., 2011; Mungall and Brenan,
607 2014). As melting proceeds, sulphides are progressively dissolved into the melt and
608 the PGE concentrate in the residual sulphide melt (Mungall and Brenan, 2014).
609 Assuming that chemical equilibrium is reached, subsequent melt increments should
610 lead to a slightly increase of whole rock PGE contents, leaving elemental ratios almost
611 constant. At the point when sulphide is completely removed, IPGE and Pt are
612 accommodated in metallic alloys, while Re, Au and Pd mass fraction should become
613 extremely low, as these elements are not hosted in any residual mantle phase
614 (Mungall and Brenan, 2014). Hence, for high melting degrees, HSE abundances in the
615 residue should reflect mineral-melt partitioning and the P-T and f_{O_2} -dependent
616 solubility of Pt and IPGE alloys in silicate melt (Fonseca et al., 2011, 2012; Mungall
617 and Brenan, 2014; Brenan et al., 2016).

618 The IPGE-PPGE fractionation and the resolvable fractionations between specific PGE
619 displayed by the type-A harzburgites bear witness of high melt extraction degrees,
620 which resulted in the formation of a S-free mantle residue. The fractionated Os-Ir-Ru-
621 Rh segments in the HSE patterns and the positive Pt anomalies are likely carried by
622 tiny residual sulphides (i.e. laurite) and metallic alloys (Os-Ir and Pt-Ir, see Lorand et
623 al., 1999; Luguet et al., 2001, 2007). These latter are thought to precipitate from
624 sulphide melt shortly before the complete exhaustion of sulphide (Mungall and
625 Brenan, 2014) or immediately after sulphide consumption, due to f_{S_2} lowering and
626 diminished metal-sulphide complexation in the silicate melt (Fonseca et al., 2012).
627 The variable but broadly systematic IPGE inter-elemental fractionation (high Os/Ir,
628 Ru/Ir and Ru/Rh) and the occurrence of positive Pt anomalies possibly suggest the
629 presence of different residual Ir-Pt alloy proportions and preferred Os-Ru retention
630 compared to Ir in the residual PGE alloys (e.g. Brenan and Andrews, 2001; Fonseca et
631 al., 2012).

632 The HSE fractionations observed for type-A harzburgites are different from HSE
633 patterns of modern harzburgites in MOR environments (Fig. 6a), as the latter are
634 characterised by flat or weakly-fractionated Os-Ir-Ru triplet, rarely displaying positive
635 Pt spikes (Snow and Schmidt, 1998; Luguet et al., 2001, 2003). By contrast, HSE
636 elemental fractionations of type-A harzburgites resemble those observed for some arc
637 xenoliths (Saha et al., 2005; Liu et al., 2015; Scott et al., 2019) or ophiolitic
638 peridotites bearing a supra-subduction zone affinity (see Büchl et al., 2002, 2004;
639 O'Driscoll et al., 2012). Notably, similar HSE fractionations have also been reported
640 for mantle xenoliths from the Chatam Islands (New Zealand, see Pearson et al.,
641 2004).

642 Accordingly, trace element geochemical modelling has shown that the extreme
643 depletion in trace element contents displayed by the New Caledonia harzburgites was
644 achieved through a polyphase evolution, including a first melting event in a mid-ocean

645 ridge setting, followed by fluid-assisted melting reaching clinopyroxene exhaustion in
646 a subduction system (see Secchiari et al., 2020). Such high melting degrees are
647 permissible in supra-subduction zone environments, where hydrous conditions at
648 relatively low pressures can produce melt fractions substantially exceeding 20%
649 without invoking extremely high temperature (e.g. see Ulmer, 2001). In addition,
650 LREE and FME (Sr, Ba, Pb) enrichments coupled with variable Pb isotope compositions
651 of the type-A harzburgites may be explained by syn- and post-melting interactions
652 with different subduction-related components (i.e. aqueous fluids and melts originated
653 in the forearc setting, see Secchiari et al., 2019, 2020).

654 We thus conclude that the HSE and chalcophile element signature displayed by TI1,
655 TI2 and PO4 predominantly reflect high degrees of melt extraction in a supra-
656 subduction zone environment. The positive Pt spikes suggest that Pt-rich alloys were
657 stable in the mantle residue and were only in part dissolved in the melt during melt
658 extraction.

659 The enrichments of Au are modest (0.2-1.3 ng/g Au) and may be related to fluid
660 overprint, either from slab-derived fluids (McInnes et al., 1999; Kepezhinskas et al.,
661 2002) or from low-T alteration, e.g. during obduction (e.g. Snow et al., 2000).

662 The harzburgite YA1 shows higher S and Se concentrations (Fig. 5a), which,
663 considering the significant LOI value of 6.83 wt.%, could be related to serpentinisation
664 and precipitation of secondary sulphides. The strongly fractionated HSE pattern and
665 the low concentrations of the incompatible HSE (i.e. Pd, Re) indicate that the HSE
666 budget of YA1 is also governed by melting, as for TI1, TI2 and PO4 harzburgites. The
667 low Os, Ir and Pt concentrations, are much closer to the values reported for type-B
668 harzburgites (see Table 1 and Fig. 5b). Sample YA1 can thus be seen as transitional
669 between type-A and type-B sub-group.

670

671

672 **5.4 Origin of type-B harzburgites – strong depletion followed by subduction**
673 **zone metasomatism?**

674

675 Type-B harzburgites mostly occur in the central massifs, however one sample (PO3)
676 has also been identified in the eastern zone, close to the area where one type-A
677 harzburgite (PO4) was sampled. Type-B harzburgites display remarkably different HSE
678 and Re/Os behaviour compared to type-A sub-group (Fig. 2 and 6).

679 Type-B harzburgites show low abundances of incompatible chalcophile elements, i.e.
680 Pd, S, Se and Re, with values in the range of type-A harzburgites. In principle such
681 low concentrations could be reconciled with high melting degrees and sulphide
682 exhaustion in the mantle source. The strong depletion of Os, Ir and Pt, (Fig. 6b)
683 coupled with slightly subchondritic to suprachondritic $^{187}\text{Re}/^{188}\text{Os}$ (0.23 to 37) and
684 $^{187}\text{Os}/^{188}\text{Os}$ (0.1239 to 0.302) are remarkable and do not occur in residual
685 harzburgites from convecting mantle, as represented by abyssal peridotites. By
686 contrast, type-B harzburgites share some similarities with mantle xenoliths from arc
687 settings, such as low Os contents associated with chondritic to suprachondritic
688 $^{187}\text{Os}/^{188}\text{Os}$ (Brandon et al., 1996, 1999; Saha et al., 2005; Widom, 2011). These
689 features have been ascribed to interactions with subduction zone fluids (i.e. fluids
690 from subducted altered oceanic crust and/or its sedimentary cover), which may have
691 induced sulphide breakdown, locally overprinting the Os isotopic signature of the
692 mantle (Wisdom et al., 2003). Such qualitative observations have also been supported
693 by experimental works, that highlighted the critical influence of oxygen fugacity (f_{O_2})
694 on sulphide and alloy stability (e.g. Andrews and Brenan, 2002; Fonseca et al., 2011,
695 2012; Mungall and Brenan, 2014). In addition, the strongly fractionated, IPGE-
696 depleted, HSE patterns of the type-B harzburgites closely resemble those observed in
697 some refractory harzburgites and replacive dunites from ophiolitic complexes that

698 underwent interaction with S-undersaturated melts (Büchl et al., 2002; Lorand et al.,
699 2004).

700 As metasomatism by subduction fluids and hydrous melts has been proposed for the
701 New Caledonia harzburgites based on isotopic and incompatible element studies, we
702 have determined oxygen fugacities on a set of five harzburgites, in order to test if the
703 HSE signature of the two sub-groups could reflect different oxygen fugacity
704 conditions.

705 As a whole, the harzburgites in our dataset bear witness of similar oxygen fugacity
706 conditions, showing no significant difference among sub-types A and B. Both groups
707 register oxygen fugacity values close to or only slightly higher than the FMQ buffer,
708 displaying a nearly horizontal trend in the Cr# vs $\log(f_{O_2})$ Δ FMQ variation diagram
709 (Fig. 7b).

710 Partial melting of a spinel peridotite source is expected to induce progressive $Fe^{3+}/\Sigma Fe$
711 lowering in the residual mantle, due to the preferential removal of Fe^{3+} in Al-rich
712 spinel and clinopyroxene during melt extraction (Woodland et al., 2006). Thus, as
713 melting proceeds and the aforementioned phases are removed from the mantle
714 assemblage, melt extraction in un-buffered conditions should result in f_{O_2} lowering in
715 the residual mantle, generating negative Cr# vs. $\log(f_{O_2})$ Δ FMQ trends (e.g. Brandon
716 and Draper, 1996). However, these correlations are rarely preserved in mantle
717 peridotites, as subsequent geochemical modifications, e.g. metasomatism, tend to
718 decouple f_{O_2} from melting depletion indexes. In particular, in the sub-arc region,
719 interaction of the refractory peridotite with subduction components is expected to shift
720 the oxygen fugacity state toward higher values, due to the high oxidation capacity of
721 slab-derived fluids and melts (e.g. Parkinson and Arculus, 1999; Brandon and Draper,
722 1996; Parkinson et al., 2003).

723 The subhorizontal trends observed in Fig. 7a-b, as well as the slightly oxidised f_{O_2}
724 values, thus bear evidence that oxygen fugacity was modified during melting and/or

725 post-melting evolution of the New Caledonia harzburgites. The measured f_{O_2} are
726 similar to those observed for the less oxidised peridotites from the Izu-Bonin region,
727 plotting at the lower end of the arc xenoliths domain (see Fig. 7b; Wood and Virgo,
728 1989; Brandon and Draper, 1996; Parkinson et al., 2003). Such oxygen fugacity
729 values most likely reflect limited interaction with slab-derived fluids and melts, in
730 agreement with the geodynamic scenario proposed for the New Caledonia archipelago
731 and the previous geochemical models (see Secchiari et al., 2020). In fact, Eocene
732 subduction is believed to have started close to or in correspondence of an active
733 oceanic spreading center, where hot and young ($\sim 6-9$ My old, Cluzel et al., 2016)
734 lithosphere was forced to subduct. In such a context, fluid fluxes from the downgoing
735 slab must have been limited, due to the young age of the subducted material and the
736 intra-oceanic nature of the subduction (Cluzel et al., 2016). In addition, post melting
737 metasomatism involved small fractions (0.5-1%) of depleted (boninitic) melts, which
738 may have not been able to shift mantle oxidation state toward more oxidised
739 conditions. In fact, while large oxygen fugacity variations in mantle peridotites are
740 permissible for low Fe^{3+} contents in spinel, more and more Fe^{3+} has to be added to
741 cause even a small increase in f_{O_2} when f_{O_2} approaches FMQ buffer (Woodland et al.,
742 2006).

743 The similar f_{O_2} values registered by the two harzburgite sub-types attest that HSE
744 behaviour was not critically influenced by different oxygen fugacity conditions. This
745 observation supports the hypothesis that the two types of patterns are not directly
746 linked to a specific process (i.e. higher melt extraction or enrichment degree or
747 interaction with an oxidised component). Rather, the occurrence of two distinct
748 patterns in the harzburgites may reflect an effect of source-control on the HSE
749 behaviour during the recent evolution, indicating the presence of geochemically
750 heterogeneous mantle domains below the New Caledonia archipelago before Eocene
751 subduction.

752

753 **5.5 Type-B harzburgites: a broader perspective**

754

755 Despite being similar in terms of chemistry or mineralogy, type-A and B harzburgites
756 possess distinct HSE signatures. In addition, the HSE signatures of type-B harzburgites
757 have not yet been identified in other mantle tectonites, either from modern oceanic
758 lithosphere or ophiolitic complex.

759 However, similar compositions have been recently reported for some moderately
760 depleted to highly refractory peridotites and mantle xenoliths from New Zealand
761 (Scott et al., 2019). The New Zealand mantle is composed of isotopically
762 heterogeneous mantle fragments with evolutionary histories extending over 2.75 Ga
763 (Os model ages= 0.1-2.75 Ga, with a broad peak at 1.2 Ga), and PGE systematics
764 decoupled from major element compositions (see Scott et al., 2019; Liu et al., 2015).
765 These features have been explained by accretion of Zealandia lithospheric mantle
766 from amalgamation of genetically unrelated convecting mantle fragments which were
767 swept together beneath the Gondwana subduction margin, variably re-melted and
768 laterally accreted (Scott et al., 2019).

769 Among the New Zealand peridotite suites, mantle xenoliths from Lake Moana and
770 Chatam Island show HSE patterns that are similar to our dataset (Fig. 8). Lake Moana
771 Cretaceous xenoliths include cpx-free harzburgites that experienced up to 30%
772 melting, while Eocene-aged Chatam Island harzburgites exhibit a less refractory
773 nature, as attested by the presence of primary clinopyroxene (up to 1.8% modal,
774 Scott et al., 2016). Such depletion degrees were achieved either by plume melting or
775 hydrous melting in an arc setting for the Lake Moana xenoliths, whereas Chatam
776 Island samples are thought to represent fragments of fore-arc lithosphere (Scott et al.,
777 2016, 2019).

778 Overall, HSE diagrams highlight that New Zealand mantle xenoliths reproduce with
779 good approximation both patterns observed in our harzburgites, namely type-A and
780 type-B (Fig. 8). IPGE patterns are broadly sub-parallel, with the New Caledonia
781 harzburgites falling within or at the lower range of values displayed by the New
782 Zealand samples, whereas Pt, Pd and Re exhibit greater variability.

783 For most of the New Zealand xenoliths, Ru concentrations are higher compared to the
784 contents in the New Caledonia harzburgites. Higher Ru contents cannot be ascribed to
785 different degrees of melting, as fertility indicators (i.e. Mg#(Ol), Cr#(Spl)) indicate
786 comparable depletion degrees for both peridotite suites (see Scott et al., 2016, 2019;
787 Secchiari et al., 2020). The increased Ru retention in the New Zealand samples may
788 be reconciled to the higher Cr-spinel content of these lithologies (up to 2.8%, see
789 Scott et al., 2016, vs. up to 0.8% for our harzburgites). Numerous studies have in
790 fact demonstrated that spinel can be a significant host for Ru ($D^{\text{spinel/melt}} \sim 20$,
791 Capobianco and Drake, 1990). In addition, increase in oxygen fugacity markedly
792 enhances Ru compatibility in Cr-rich spinel ($D^{\text{spinel/melt}}$ up to 500 for f_{O_2} of -0.5 FMQ,
793 Park et al., 2012), which can accommodate Ru within its crystal lattice (Pagé and
794 Barnes, 2016) or as laurite and/or Ru-rich alloy inclusions (Brenan and Andrews,
795 2001). Recent geochemical works have also illustrated the importance of Ru retention
796 in the sub-arc mantle for the HSE signature of arc lavas (Dale et al., 2012; Park et al.,
797 2013). These studies explain the low Ru concentrations and the high Pt/Ru ratios of
798 the volcanic products as related to Ru retention in the mantle source due to the
799 presence of Cr-rich spinel or PGM (see Dale et al., 2012; Park et al., 2013). Likewise,
800 positive Ru anomalies in our type-B harzburgites may reflect the presence of small
801 laurite or Ru-rich phase inclusions, which could have escaped the high melting
802 degrees due to the shield effect of spinel.

803 Other HSE (e.g. Pd and Re) in the New Caledonia harzburgites display a wider range
804 of values compared to the New Zealand samples. Enrichments in Pd have also been

805 recognised in other sub-arc mantle sections, where they have been attributed to slab-
806 derived fluids metasomatism, due to the high solubility of Pd in aqueous fluids
807 (McInnes et al., 1999; Park et al., 2013). Likewise, Re addition in mantle wedge
808 peridotites may be due to Re release in slab derived fluids during dehydration of the
809 mafic portion of the subducting slab (see Dale et al., 2009). Hence, we conclude that
810 Pd-Re (as well as Te-Se-S) re-enrichments in type-B harzburgites may have been
811 facilitated by minor sulphide precipitation from slab-derived fluids or melts.
812 By contrast, the widespread negative Pt anomalies may reflect destabilisation of a
813 pre-existing Pt-alloy phase, possibly related to melting after alloy saturation (see
814 Mungall and Brenan, 2016), and Pt release into the melt.

815

816 **5.6 Inferences from Re-Os systematics of the New Caledonia harzburgites**

817

818 All type-A harzburgites, including YA1, show unradiogenic $^{187}\text{Os}/^{188}\text{Os}_i$ and very low
819 Re/Os relative to the range of chondritic values (Walker et al., 2002) or primitive
820 mantle estimates (Meisel et al., 1996). The γ_{Os_i} overlap with data from depleted
821 abyssal peridotites and mantle sections from some ophiolites (e.g., Becker and Dale,
822 2016).

823 In order to obtain an estimate of the time of melt depletion, we calculated Re-Os
824 model ages (T_{MA}) and Re-depletion model ages (T_{RD}). T_{RD} is generally used to estimate
825 model ages of mantle peridotites that experienced high degrees of melt extraction,
826 leading to low Re contents and low Re/Os ratios. For mantle peridotites which are
827 pure residues of high partial melting degrees sufficient to remove all Re, then T_{MA} and
828 T_{RD} ages should be identical. If a minor amount of Re remains, the two values should
829 be relatively similar, with T_{RD} providing a minimum depletion age.

830 The Re-depletion ages for type-A harzburgites are quite homogeneous for three of our
831 samples (TI1, TI2, YA1) ranging between 0.4 and 0.7 Ga (see Table 1), while sample
832 PO4 yields an older Re-depletion age of 1.3 Ga. Slightly older ages are provided by
833 T_{MA} (0.5- 1.0 Ga for TI1, TI2, YA1 and 1.5 Ga for PO4), indicating that Re was not
834 effectively removed after partial melting.

835 The younger model ages recorded by the harzburgites are likely linked to the
836 evolution of the New Caledonia mantle in relationship to the eastern Australian
837 margin, from which the New Caledonia archipelago was separated *via* marginal rifting
838 about 90 Ma ago (Cluzel et al., 2001, 2012; Whattam, 2009). The ancient Re
839 depletion age of the harzburgite PO4, on the other hand, reflects a mantle domain
840 characterised by long-term low Re/Os ratio. This age is also mirrored by Nd isotopic
841 signature, which shows a highly radiogenic value ($\epsilon_{Nd_i} = +13.32$, Secchiari et al.,
842 2020), indicative of a mantle reservoir that underwent long-term depletion of Nd.

843 In addition, similar depletion ages (i.e. 1.2 Ga) occur in the New Zealand mantle rocks
844 (see Scott et al., 2019).

845 These results are consistent with recent Re-Os studies on abyssal peridotites and
846 mantle tectonites from ophiolitic sequences showing that the convecting mantle
847 contains harzburgite domains that underwent depletion events much older than the
848 age of peridotite processing under the ridge (e.g., Harvey et al., 2006).

849 In summary, the New Caledonia harzburgites show Re-Os systematics typical of
850 depleted upper mantle, showing Early Paleozoic Os isotopic equilibration and evidence
851 for ancient depletion events (> 1.0 Ga). This multi-stage history led to the depleted
852 nature of the harzburgites and possibly resulted in their complex HSE geochemical
853 signatures.

854

855 **6. Summary and conclusions**

856

857 A Re-Os isotopes, highly siderophile and chalcophile element investigation of the New
858 Caledonia peridotites was carried out to unravel the behaviour of the aforementioned
859 elements in lherzolites and harzburgites from the New Caledonia ophiolite.

860 The lherzolites exhibit subchondritic to slightly suprachondritic $^{187}\text{Os}/^{188}\text{Os}_i$ (0.1273-
861 0.1329). PM-normalised HSE abundance diagrams are characterised by gently sloping
862 patterns showing increasing depletion towards Re-Au, similar to lherzolites that
863 experienced low to moderate melt extraction. However, the lack of correlation
864 between HSE and fertility indicators, as well as the slightly suprachondritic Os isotopic
865 ratios, argue against a simple partial melting history. Rather, such features and the
866 presence of included and interstitial residual monosulphides possibly indicate that
867 melting involved a mantle domain that has previously experienced a melt percolation
868 event. The high S concentrations of the lherzolites (202-1268 ppm) most likely
869 resulted from late-stage seawater-rock reactions.

870 By contrast, the New Caledonia harzburgites record higher degrees of melt extraction,
871 as attested by the strikingly low, often below the detection limit, concentrations of
872 incompatible chalcophile elements. Despite their homogeneity in terms of
873 mineralogical and major element compositions, HSE patterns and Os isotopic
874 compositions indicate the occurrence of two distinct harzburgite sub-groups.

875 Type-A harzburgites are characterised by steeply plunging HSE patterns, showing
876 IPGE-PPGE and Os-Ir-Ru fractionation, coupled with low Re/Os ratios and
877 subchondritic $^{187}\text{Os}/^{188}\text{Os}_i$. The strongly fractionated HSE patterns and the positive Pt
878 anomalies, together with the high modelled melting degrees, indicate that melting
879 occurred under hydrous conditions in sub-arc mantle.

880 Type-B harzburgites display notably different HSE patterns, showing depleted Os-Ir
881 compared to Ru, positive anomalies and Pd-Re re-enrichments (relative to IPGE),
882 coupled with chondritic to strongly suprachondritic measured Os isotopic ratios

883 ($^{187}\text{Os}/^{188}\text{Os} = 0.127\text{-}0.153$). These features have not been yet identified in mantle
884 tectonites.

885 The similar and slightly oxidising oxygen fugacity values recorded by the two sub-
886 types, named type-A and B, do not convincingly support the hypothesis that type-B
887 patterns could be ascribed to massive interactions with an oxidising agent. By
888 contrast, this observation, coupled with the uniform mineralogical and geochemical
889 characters of the New Caledonia harzburgites, suggest that type-B patterns were not
890 the result of a specific depletion or enrichment process related to their recent tectonic
891 evolution in a subduction setting. Rather, these patterns could reflect the occurrence
892 of a geochemically heterogeneous mantle source below the New Caledonia
893 archipelago, similar to what was speculated for the New Zealand mantle. The HSE
894 signature carried by the studied peridotites, as well as the puzzling similarity observed
895 between the New Caledonia harzburgites and the New Zealand mantle xenoliths,
896 might attest the presence of a mantle source bearing a long lasting evolution (> 1
897 Ga), possibly linked to the Zealandia formation.

898

899 Acknowledgements

900

901 The authors are grateful to M. Feth, R. Milke and D. Cluzel for help and support during
902 clean laboratory work, microprobe analysis and sampling in New Caledonia,
903 respectively. A.G. wishes to thank Russian Science Foundation project 17-77-10103
904 for supporting iron oxidation state investigations at Freie Universität of Berlin. We are
905 indebted to Nadia Malaspina and Tiziano Catelani (University of Milano Bicocca) for
906 providing access to SEM-EDS analyses of sulphides.

907 We also thank C. Bonadiman and A. Woodland for editorial handling. Constructive
908 reviews by two anonymous reviewers are gratefully acknowledged.

909 This study represents a continuation of the PhD project of A.S. on the New Caledonia
910 ophiolite, that has been supported by DFG funding (SFB-TRR 170 Subproject B2) and
911 Italian-PRIN prot. 2015C5LN35. This work has also benefited from the equipment and
912 framework of the COMP-HUB Initiative, funded by the 'Departments of Excellence'
913 program of the Italian Ministry for Education, University and Research (MIUR, 2018-
914 2022).

915

916 References

917

918 Ackerman, L., Walker, R.J., Puchtel, I.S., Pitcher, L., Jelínek, E., Strnad, L., 2009.
919 Effects of melt percolation on highly siderophile elements and Os isotopes in
920 subcontinental lithospheric mantle: A study of the upper mantle profile beneath
921 Central Europe. *Geochimica et Cosmochimica Acta* 73, 2400–2414.

922 Aitchison, J.C., Clarke, L., Meffre, S., Cluzel, D., 1995. Eocene arc-continent collision
923 in New Caledonia and implications for regional Southwest Pacific tectonic evolution.
924 *Geology* 23, 161–164.

925 Alard, O., Luguet, A., Pearson, N.J., Griffin, W.L., Lorand, J.P., Gannoun, A., Burton,
926 K.W., O'Reilly S.Y., 2005. In situ Os isotopes in abyssal peridotites bridge the isotopic
927 gap between MORBs and their source mantle. *Nature* 436, 1005-1008.

928 Alt, J.C., Shanks, W.C., 1998. Sulfur in serpentinized oceanic peridotites:
929 Serpentinization processes and microbial sulfate reduction. *Journal of Geophysical*
930 *Research* 103, 9917–9929.

931 Alt, J.C., Shanks, W.C., 2003. Serpentinization of abyssal peridotites from the MARK
932 area, Mid-Atlantic Ridge: Sulfur geochemistry and reaction modeling. *Geochimica et*
933 *Cosmochimica Acta* 67, 641 – 653.

934 Andrews, D.R.A., Brenan, J.M., 2002. Phase-equilibrium constraints on the magmatic
935 origin of laurite + Ru–Os–Ir alloy. *Canadian Mineralogist* 40, 1705-1716.

936 Becker, H., Dale, C., 2016. Re–Pt–Os Isotopic and Highly Siderophile Element
937 Behavior in Oceanic and Continental Mantle Tectonites. *Reviews in Mineralogy and*
938 *Geochemistry*, 369-440

939 Becker, H., Horan, M.F., Walker, R.J., Gao, S., Lorand, J.P., Rudnick, R.L., 2006.
940 Highly siderophile element composition of the Earth 's primitive upper mantle :
941 Constraints from new data on peridotite massifs and xenoliths. *Geochimica et*
942 *Cosmochimica Acta* 70, 4528–4550.

943 Birck, J.L., Barman, M.R., Capmas, F., 1997. Re-Os Isotopic Measurements at the
944 Femtomole Level in Natural Samples. *Geostandards and Geoanalytical Research* 21,
945 19–27.

946 Birner, S.K., Warren, J. M., Cottrell, E., Davis, F. A., Kelley, K. A., & Falloon, T. J.
947 (2017). Forearc peridotites from Tonga record heterogeneous oxidation of the mantle
948 following subduction initiation. *Journal of Petrology*, 58(9), 1755–1780.

949 Brandon, A.D., Creaser, R.A., Shirey, S.B., Carlson, R.W., 1996. Osmium recycling in
950 subduction zones. *Science* 272, 861– 864.

951 Brandon, A.D., Draper, D.S., 1996. Constraints on the origin of the oxidation state of
952 mantle overlying subduction zones: an example from Simcoe, Washington, USA.
953 *Geochimica et Cosmochimica Acta* 60, 1739-1749.

954 Brandon, A.D., Norman, M.D., Walker, R.J., Morgan, J.W., 1999. 186Os–187Os
955 systematics of Hawaiian picrites. *Earth and Planetary Science Letters* 174, 25– 42.

956 Brandon, A.D., Snow J.E., Walker R.J., Morgan J.W., Mock, T.D., 2000. 190Pt-186Os
957 and 187Re-187Os systematics of abyssal peridotites. *Earth and Planetary Science*
958 *Letters* 177, 319-335.

959 Brenan, J.M., Andrews, D., 2001. High-temperature stability of laurite and Ru Os Ir
960 alloy and their role in PGE fractionation in mafic magmas. *Canadian Mineralogist* 39,
961 341–360.

962 Brenan, J.M., Bennett, N.R., Zajacz, Z., 2016. Experimental results on fractionation of
963 the highly siderophile elements (HSE) at variable pressures and temperatures during
964 planetary and magmatic differentiation. *Reviews in Mineralogy and Geochemistry* 81,
965 1–87.

966 Bryndzia, L.T., Wood, B.J., 1990. Oxygen thermobarometry of abyssal spinel
967 peridotites; the redox state and C-O-H volatile composition of the Earth's sub-oceanic
968 upper mantle. *American Journal of Science* 290, 1093–1116.

969 Brügmann, G., Batanova, V.G., Münker, C., Hofmann, A.W., 2002. Melt percolation
970 monitored by Os isotopes and HSE abundances : a case study from the mantle section
971 of the Troodos Ophiolite. *Earth and Planetary Science Letters* 204, 385–402.

972 Büchl, A., Brügmann, G. E., Batanova, V. G., & Hofmann, A. W., 2004. Os
973 mobilization during melt percolation: The evolution of Os isotope heterogeneities in
974 the mantle sequence of the Troodos ophiolite, Cyprus. *Geochimica et Cosmochimica*
975 *Acta*, 68, 3397-3408.

976 Bulanova, G.P., Griffin, W.L., Ryan, C.G., Shestakova, O.Y., Barnes, S.-J., 1996. Trace
977 elements in sulfide inclusions from Yakutian diamonds. *Contributions to Mineralogy
978 and Petrology* 124, 111–125.

979 Cluzel, D., Aitchison, J.C., Picard, C., 2001. Tectonic accretion and underplating of
980 mafic terranes in the Late Eocene intraoceanic fore-arc of New Caledonia (Southwest
981 Pacific): geodynamic implications. *Tectonophysics* 340, 23–59.

982 Cluzel, D., Maurizot, P., Collot, J., Sevin, B., 2012. An outline of the Geology of New
983 Caledonia ; from Permian – Mesozoic Southeast Gondwanaland active margin to
984 Cenozoic obduction and supergene evolution. *Episodes* 35, 72–86.

985 Cluzel, D., Meffre, S., Maurizot, P., Crawford, A.J., 2006. Earliest Eocene (53 Ma)
986 convergence in the Southwest Pacific; evidence from pre- obduction dikes in the
987 ophiolite of New Caledonia. *Terra Nova* 18, 395–402.

988 Cluzel, D., Ulrich, M., Jourdan, F., Meffre, S., Paquette, J.L., Audet, M.A., Secchiari,
989 A., Maurizot, P., 2016. Early Eocene clinoenstatite boninite and boninite-series dikes
990 of the ophiolite of New Caledonia; a witness of slab-derived enrichment of the mantle
991 wedge in a nascent volcanic arc. *Lithos* 260, 429–442.

992 Cohen, A.S., Waters, F.G., 1996. Separation of osmium from geological materials by
993 solvent extraction for analysis by thermal ionisation mass spectrometry. *Analytica
994 Chimica Acta* 332, 269–275.

995 Dale, C.W., Burton, K.W., Pearson, D.G., Gannoun, A., Alard, O., Argles, T.W.,
996 Parkinson, I.J., 2009. Highly siderophile element behaviour accompanying subduction
997 of oceanic crust : Whole rock and mineral-scale insights from a high-pressure terrain.
998 *Geochimica et Cosmochimica Acta* 73, 1394–1416.

999 Debret, B., Andreani, M., Delacour, A., Rouméjon S., Trcera, N., EIMF, Williams, H.,
1000 2017. E Assessing sulfur redox state and distribution in abyssal serpentinites using
1001 XANES spectroscopy. *Earth and Planetary Science Letters* 466, 1–11.

1002 Fischer-Gödde, M., Becker, H., Wombacher, F., 2011. Rhodium , gold and other highly
1003 siderophile elements in orogenic peridotites and peridotite xenoliths. *Chemical*
1004 *Geology* 280, 365–383.

1005 Fonseca, R.O.C., Mallmann, G., St.C.O'Neill, H., Campbell, I.H., 2007. How chalcophile
1006 is rhenium? An experimental study of the solubility of Re in sulphide mattes. *Earth*
1007 *and Planetary Science Letters* 260, 537–548.

1008 Fonseca, R.O.C., Mallmann, G., O'Neill, H.S., Campbell, I.H., Laurenz, V., 2011.
1009 Solubility of Os and Ir in sulfide melt: Implications for Re/Os fractionation during
1010 mantle melting. *Earth and Planetary Science Letters* 311, 339–350.

1011 Fonseca, R.O.C., Laurenz, V., Mallmann, G., Luguet, A., Hoehne, N., Jochum, K.P.,
1012 2012. New constraints on the genesis and long-term stability of Os-rich alloys in the
1013 the Earth's mantle. *Geochimica et Cosmochimica Acta* 87, 227-242.

1014 Goncharov, A.G., Ionov, D.A., 2012. Redox state of deep off-craton lithospheric
1015 mantle: New data from garnet and spinel peridotites from Vitim, southern Siberia.
1016 *Contributions to Mineralogy and Petrology*, 164 (5), 731-745.

1017 Goncharov, A.G., Nikitina, L.P., Borovkov, N.V., Babushkina, M.S., Sirotkin, A.N.,
1018 2015. Thermal and redox equilibrium conditions of the upper-mantle xenoliths from
1019 the Quaternary volcanoes of NW Spitsbergen, Svalbard Archipelago. *Russian Geology*
1020 *and Geophysics*, 56 (11), 1578-1602.

1021 Goncharov, A. , 2018. In-situ iron oxidation state in the upper mantle minerals using
1022 electron microprobe. Contributions to 3rd European mantle workshop. *Plinius* 44, p.
1023 172. DOI:10.19276/plinius.2018.03015

1024 Guo J., Griffin W. L., O'Reilly S. Y., 1999. Geochemistry and origin of sulphide
1025 minerals in mantle xenoliths: Qilin, Southeastern China. *Journal of Petrology* 40,
1026 1125–1149.

1027 Harvey, J., Gannoun, A., Burton, K.W., Rogers, N.W., Alard, O., Parkinson, I.J., 2006
1028 Ancient melt extraction from the oceanic upper mantle revealed by Re–Os isotopes in
1029 abyssal peridotites from the Mid-Atlantic ridge. *Earth and Planetary Science Letters*
1030 244, 606–621.

1031 Hattori, K.H., Arai, S., Clarke Barrie, D.B., 2002. Selenium, tellurium, arsenic and
1032 antimony contents of primary mantle sulfides. *Canadian Mineralogist* 40, 637–650.

1033 Helmy, H.M., Ballhaus, C., Wohlgemuth-Ueberwasser, C., Fonseca, R.O.C., Laurenz,
1034 V., 2010. Partitioning of Se, As, Sb, Te and Bi between monosulfide solid solution and
1035 sulfide melt – Application to magmatic sulfide deposits. *Geochimica et Cosmochimica*
1036 *Acta* 74, 6174–6179.

1037 Höfer H. E., Brey G. P., 2007. The iron oxidation state of garnet by electron
1038 microprobe: Its determination with the flank method combined with major-element
1039 analysis. *American Mineralogist* 92, 873–885.

1040 Kepezhinskas, P., Defant, M. J. and Widom, E., 2002. Abundance and distribution of
1041 PGE and Au in the island-arc mantle: implications for sub-arc metasomatism. *Lithos*
1042 60, 113–128.

1043 Klein, F., Bach, W., 2009. Fe–Ni–Co–O–S Phase Relations in Peridotite–Seawater
1044 Interactions. *Journal of Petrology* 50, 37–59.

1045 Lagabrielle, Y., Chauvet, A., Ulrich, M., Guillot, S., 2013. Passive obduction and
1046 gravity-driven emplacement of large ophiolitic sheets: The New Caledonia ophiolite
1047 (SW Pacific) as a case study? *Bulletin de la Société Géologique de France* 184, 545–
1048 556.

1049 Lassiter, J.C., Byerly, B.L., Snow, J.E., Hellebrand, E., 2014. Constraints from Os-
1050 isotope variations on the origin of Lena Trough abyssal peridotites and implications for
1051 the composition and evolution of the depleted upper mantle. *Earth and Planetary
1052 Science Letters* 403, 178–187.

1053 Lee, C.T.A., Leeman, W.P., Canil, D., Li, Z.X.A., 2005. Similar V/Sc systematics in
1054 MORB and arc basalts: Implications for the oxygen fugacities of their mantle source
1055 regions. *Journal of Petrology* 46, 2313–2336.

1056 Liu, C., Snow, J.E., Brüggmann, G., Hellebrand, E., Hofmann, A.W., 2009. Non-
1057 chondritic HSE budget in Earth 's upper mantle evidenced by abyssal peridotites from
1058 Gakkel ridge (Arctic Ocean). *Earth and Planetary Science Letters* 283, 122–132.

1059 Liu, J., Scott, J.M., Martin, C.E., Pearson, D.G., 2015. The longevity of Archean mantle
1060 residues in the convecting upper mantle and their role in young continent formation.
1061 *Earth and Planetary Science Letters* 424, 109–118.

1062 Lorand, J.-P., 1991. Sulphide Petrology and Sulphur Geochemistry of Orogenic
1063 Lherzolites: A Comparative Study of the Pyrenean Bodies (France) and the Lanzo
1064 Massif (Italy). *Journal of Petrology Special Volume*, 77–95.

1065 Lorand, J.-P., Pattou, L., Gros, M., 1999. Fractionation of Platinum-group Elements
1066 and Gold in the Upper Mantle: a Detailed Study in Pyrenean Orogenic Lherzolites.
1067 *Journal of Petrology* 40, 957–981.

1068 Lorand, J., Luguet, A., Alard, O., 2008. Platinum-Group Elements : A New Set of Key
1069 Tracers for the Earth's Interior. *Elements* 4, 247–252.

1070 Lorand, J.P., Alard, O., 2010. Determination of selenium and tellurium concentrations
1071 in Pyrenean peridotites (Ariege, France): New insight into S/Se/Te systematics of the
1072 upper in mantle samples. *Chemical Geology* 278, 120–130.

1073 Lorand, J.P., Alard, O., Luguet, A., Keays, R.R., 2003. Sulfur and selenium
1074 systematics of the subcontinental lithospheric mantle: Inferences from the Massif
1075 Central xenolith suite (France). *Geochimica et Cosmochimica Acta* 67, 4137–4151.

1076 Lorand, J.P., Delpech, G., Grégoire, M., Moine, B., O'Reilly, S.Y., Cottin, J.Y., 2004.
1077 Platinum-group elements and the multistage metasomatic history of Kerguelen
1078 lithospheric mantle (South Indian Ocean). *Chemical Geology* 208, 195–215.

1079 Luguet, A., Alard, O., Lorand, J.P., Pearson, N.J., Ryan, C.Y., Reilly, S.Y.O., 2001.
1080 Laser-ablation microprobe (LAM) -ICPMS unravels the highly siderophile element
1081 geochemistry of the oceanic mantle. *Earth and Planetary Science Letters* 189, 285–
1082 294.

1083 Luguet, A., Lorand, J.P., Seyler, M., 2003. Sulfide petrology and highly siderophile
1084 element geochemistry of abyssal peridotites : A coupled study of samples from the
1085 Kane Fracture Zone (45 ° W 23 ° 20N , MARK Area , Atlantic Ocean). *Geochimica et*
1086 *Cosmochimica Acta* 67, 1553–1570.

1087 Luguet, A., Lorand, J., Alard, O., Cottin, J., 2004. A multi-technique study of platinum
1088 group element systematic in some Ligurian ophiolitic peridotites , Italy. *Chemical*
1089 *Geology* 208, 175–194.

1090 Luguet, A., Shirey, S.B., Lorand, J.P., Horan, M.F., Carlson, R.W., 2007. Residual
1091 platinum-group minerals from highly depleted harzburgites of the Lherz massif

1092 (France) and their role in HSE fractionation of the mantle. *Geochimica et*
1093 *Cosmochimica Acta* 71, 3082–3097.

1094 Marchesi, C., Garrido, C.J., Godard, M., Belley, F., Ferré, E., 2009. Migration and
1095 accumulation of ultra-depleted subduction-related melts in the Massif du Sud ophiolite
1096 (New Caledonia). *Chemical Geology* 266, 171–186.

1097 Marchesi, C., Garrido, C.J., Harvey, J., 2013. Platinum-group elements , S , Se and Cu
1098 in highly depleted abyssal peridotites from the Mid-Atlantic Ocean Ridge (ODP Hole
1099 1274A): Influence of hydrothermal and magmatic processes. *Contributions to*
1100 *Mineralogy and Petrology* 166, 1521–1538.

1101 Mathez, E.A., 1976. Sulfur solubility and magmatic sulfides in submarine basalt glass.
1102 *Journal of Geophysical Research* 81, 4269–4276.

1103 McInnes, B. I. A., McBride, J. S., Evans, N. J., Lambert, D. D. & Andrew, A. S., 1999.
1104 Osmium isotope constraints on ore metal recycling in subduction zones. *Science* 286,
1105 512–516.

1106 Meisel, T., Walker, R.J., Irving, A.J., Lorand, J.-P., 2001. Osmium isotopic
1107 compositions of mantle xenoliths: a global perspective. *Geochimica et Cosmochimica*
1108 *Acta* 65, 1311–1323.

1109 Mungall J.E., Brenan, J., 2014. Partitioning of platinum-group elements and Au
1110 between sulfide liquid and basalt and the origins of mantle-crust fractionation of the
1111 chalcophile elements. *Geochimica et Cosmochimica Acta* 125, 265–289.

1112 O’Driscoll, B., Day, J.M.D., Walker, R.J., Daly, J.S., McDonough, W.F., Piccoli, P.M.,
1113 2012. Chemical heterogeneity in the upper mantle recorded by peridotites and
1114 chromitites from the Shetland Ophiolite Complex, Scotland. *Earth and Planetary*
1115 *Science Letters* 333–334, 226–237.

1116 Pagé, P., Barnes, S.-J., 2016. The influence of chromite on osmium, iridium,
1117 ruthenium and rhodium distribution during early magmatic processes. *Chemical*
1118 *Geology* 420, 51–68.

1119 Park, J., Campbell, I.H., Arculus, R.J., 2013. Platinum-alloy and sulfur saturation in an
1120 arc-related basalt to rhyolite suite : Evidence from the Pual Ridge lavas , the Eastern
1121 Manus Basin. *Geochimica et Cosmochimica Acta* 101, 76–95.

1122 Parkinson, I.J., Pearce, J.A., 1998. Peridotites from the Izu–Bonin–Mariana Forearc
1123 (ODP Leg 125): evidence for mantle melting and melt–mantle interaction in a supra-
1124 subduction zone setting. *Journal of Petrology* 9, 1577–1618.

1125 Parkinson, I.J., Arculus, R.J., Eggins, S.M., 2003. Peridotite xenoliths from Grenada,
1126 Lesser Antilles Island Arc. *Contributions to Mineralogy and Petrology* 146: 241–262.

1127 Pearson, D.G., Irvine, G.J., Ionov, D.A., Boyd, F.R., Dreibus, G.E., 2004. Re-Os
1128 isotope systematics and platinum group element fractionation during mantle melt
1129 extraction: A study of massif and xenolith peridotite suites. *Chemical Geology* 208,
1130 29–59.

1131 Pirard, C., Hermann, J., O’Neill, H.St.C., 2013. Petrology and geochemistry of the
1132 crust– mantle boundary in a nascent arc, Massif du Sud ophiolite, New Caledonia, SW
1133 Pacific. *Journal of Petrology* 0, 1–34.

1134 Rehkämper, M., Halliday, A.N., Alt, J., Fitton, J.G., Zipfel, J., Takazawa, E., 1999.
1135 Non-chondritic platinum-group element ratios in oceanic mantle lithosphere:
1136 petrogenetic signature of melt percolation? *Earth and Planetary Science Letters* 172,
1137 65–81.

1138 Reisberg, L., Lorand, J.-P., 1995. Longevity of sub-continental mantle lithosphere
1139 from osmium isotope systematics in orogenic peridotite massifs. *Nature* 376, 159–
1140 162.

1141 Saha, A., Basu, A.R., Jacobsen, S.B., Poreda, R.J., 2005. Slab devolatilization and Os
1142 and Pb mobility in the mantle wedge of the Kamchatka arc. *Earth and Planetary
1143 Science Letters* 236, 182–194.

1144 Salters, V.J.M., Stracke, A., 2004. Composition of the depleted mantle. *Geochemistry
1145 Geophysics Geosystems* 5 (5).

1146 Scott, J.M., Liu, J., Pearson, D.G., Harris, G.A., Czertowicz, T.A., Woodland, S.J.,
1147 Riches, A.J.V., Luth, R.W., 2019. Continent stabilisation by lateral accretion of
1148 subduction zone-processed depleted mantle residues; insights from Zealandia. *Earth
1149 and Planetary Science Letters* 507, 175–186.

1150 Scott, J.M., Liu, J., Pearson, D.G., Waight, T.E., 2016. Mantle depletion and
1151 metasomatism recorded in orthopyroxene in highly depleted peridotites. *Chemical
1152 Geology* 441, 280–291.

1153 Secchiari, A., Montanini, A., Bosch, D., Macera, P., Cluzel, D., 2020. Sr-Nd-Pb and
1154 trace element systematics of the New Caledonia harzburgites: tracking source
1155 depletion and contamination processes in a SSZ setting. In press on *Geoscience
1156 Frontiers*, special issue Ophiolites. <https://doi.org/10.1016/j.gsf.2019.04.004>

1157 Secchiari, A., Montanini, A., Bosch, D., Macera, P., Cluzel, D., 2019. Origin of the
1158 spinel-pyroxene symplectites in the harzburgites from the New Caledonia peridotite.
1159 *Ophioliti* 44, 31-42.

1160 Secchiari, A., Montanini, A., Bosch, D., Macera, P., Cluzel, D., 2018. The contrasting
1161 geochemical message from the New Caledonia gabbro-norites: insights on depletion

1162 and contamination processes of the sub-arc mantle in a nascent arc setting.
1163 *Contributions to Mineralogy and Petrology* 173, 66.

1164 Secchiari, A., Montanini, A., Bosch, D., Macera, P., Cluzel, D., 2016. Melt extraction
1165 and enrichment processes in the New Caledonia lherzolites: evidence from
1166 geochemical and Sr–Nd isotope data. *Lithos* 260, 28–43.

1167 Secchiari, A., 2016. Sr-Nd-Pb isotope investigation of the New Caledonia Ophiolite.
1168 PhD thesis, Parma and Montpellier University, 191 pp.

1169 Snow, J.E., Reisberg, L., 1995. Os isotopic systematics of the MORB mantle: results
1170 from altered abyssal peridotites. *Earth and Planetary Science Letters* 133, 411–421.

1171 Snow, J., Schmidt, G., Rampone E., 2000. Os isotopes and highly siderophile
1172 elements (HSE) in the Ligurian ophiolites, Italy. *Earth and Planetary Science Letters*
1173 175, 119–132

1174 Standish, J.J., Hart, S.R., Blusztajn, J., Dick, H.J.B., Lee, K.L., 2002. Abyssal
1175 peridotite osmium isotopic compositions from Cr-spinel. *Geochemistry Geophysics*
1176 *Geosystems* 3, (1).

1177 Ulmer, P., 2001. Partial melting in the mantle wedge — the role of H₂O in the genesis
1178 of mantle-derived ‘arc-related’ magmas. *Physics of the Earth and Planetary Interiors*
1179 127, 215–232.

1180 Ulrich, M., Picard, C., Guillot, S., Chauvel, C., Cluzel, D., Meffre, S., 2010. Multiple
1181 melting stages and refertilization as indicators for ridge to subduction formation: the
1182 New Caledonia ophiolite. *Lithos* 115, 223–236.

1183 Walker, R.J., Hanski, E., Vuollo, J., Liipo, J., 1996. The Os isotopic composition of
1184 Proterozoic upper mantle: evidence for chondritic upper mantle from the Outokumpu
1185 ophiolite, Finland. *Earth and Planetary Science Letters* 141, 161–173.

1186 Wang, Z., Becker, H., 2013. Ratios of S, Se and Te in the silicate Earth require a
1187 volatile-rich late veneer. *Nature* 499, 328–331.

1188 Wang, Z., Becker, H., Gawronski, T., 2013. Partial re-equilibration of highly
1189 siderophile elements and the chalcogens in the mantle : A case study on the
1190 Baldissero and Balmuccia. *Geochimica et Cosmochimica Acta* 108, 21–44.

1191 Whattam, S. A., 2009. Arc-continent collisional orogenesis in the SW Pacific and the
1192 nature, source and correlation of emplaced ophiolitic nappe components. *Lithos* 113,
1193 88–114.

1194 Widom, E., Kepezhinskas, P., and Defant, M., 2003, The nature of metasomatism in
1195 the sub-arc mantle wedge: Evidence from Re-Os isotopes in Kamchatka peridotite
1196 xenoliths: *Chemical Geology* 196, 283–306.

1197 Widom, E., 2011. Recognizing recycled Osmium. *Geology* 39, 1087–1088.

1198 Wood, B.J., Virgo, D., 1989. Upper mantle oxidation state: ferric iron contents of
1199 lherzolite spinels by ^{57}Fe Mössbauer spectroscopy and resultant oxygen fugacities.
1200 *Geochimica et Cosmochimica Acta* 53, 1277–1291.

1201 Woodland, A.B., Kornprobst, J., Tabit., A., 2006. Ferric iron in orogenic lherzolite
1202 massifs and controls of oxygen fugacity in the upper mantle. *Lithos* 89, 222-241.

1203

1204 Figure captions

1205 Fig. 1. a) Present-day structures of the Southwest Pacific region modified after Cluzel
1206 et al. (2012). Dark orange, land; light orange, continental plateau; white, oceanic
1207 basins (LHR: Lord Howe Rise, NR: Norfolk ridge, LR: Loyalty ridge, HP: Hikurangi
1208 Plateau); b) simplified geological map of New Caledonia showing distribution of the
1209 Peridotite massifs (modified after Cluzel et al., 2012).

1210

1211 Fig. 2. Variations of Os, Ru, Rh, Pt, Pd, Au and Re vs. Ir for the New Caledonia
1212 peridotites. Abyssal peridotites (Kane fracture zone: Snow and Schmidt, 1998;
1213 Brandon et al., 2000; Luguet et al., 2001, 2003; Marchesi et al., 2013; MAR: Harvey
1214 et al., 2006; Lena trough: Lassiter et al., 2014) and ophiolitic peridotites (IL-EL:
1215 Internal and External Ligurides, Snow et al., 2000; Luguet et al., 2004; Fischer-Godde
1216 et al., 2011; Lanzo: Becker et al., 2006; Pyrenees: Becker et al., 2006; Luguet et al.,
1217 2007) are shown for comparison.

1218

1219 Fig. 3. Primitive mantle normalised HSE and chalcogen patterns for the New Caledonia
1220 spinel and plagioclase lherzolites. Grey shaded area includes oceanic lherzolites from
1221 Mid-Atlantic and South West Indian ridges (Snow and Schmidt, 1998; Luguet et al.,
1222 2001; Luguet et al., 2003) and ophiolitic lherzolites from the Ligurian Units (Snow et
1223 al., 2000; Luguet et al., 2004; Fischer-Gödde et al. 2011). Normalising values after
1224 Becker et al. (2006), Fischer-Gödde et al. (2011) and Wang and Becker (2013).

1225

1226 Fig. 4. a) Al_2O_3 (wt.%) $^{-187}\text{Os}/^{188}\text{Os}_i$ and b) Os - $^{187}\text{Os}/^{188}\text{Os}_i$ and c) $^{187}\text{Re}/^{188}\text{Os}$ -
1227 $^{187}\text{Os}/^{188}\text{Os}$ diagrams showing data from New Caledonia lherzolites and type-A
1228 harzburgites in comparison to PM compositions. PM data from Meisel et al. (1996).
1229 Data for abyssal peridotites are from Harvey et al., (2006) for Atlantic peridotites,
1230 Lassiter et al., (2014) for Lena through and Liu et al. (2015) for Gakkel ridge. See Fig.
1231 2 for ophiolitic peridotites references.

1232

1233 Fig. 5. a) Te vs. Se, b) Te vs. Pd and c) S vs. S/Se correlation diagrams for New
1234 Caledonia spinel and plagioclase lherzolites. Data for orogenic lherzolites from Wang
1235 and Becker (2013) are also plotted for comparison.

1236

1237 Fig. 6. a) Primitive mantle normalised HSE and chalcogen abundances in type-A
1238 harzburgites. Light grey shaded field encompasses the area of modern MOR
1239 harzburgites (Snow and Schmidt, 1998; Luguet et al., 2001, 2003; Harvey et al.,
1240 2006; Marchesi et al., 2013); b) Primitive mantle normalised diagram showing HSE
1241 and chalcogen patterns of type-B harzburgites. Normalising values are after Becker et
1242 al. (2006), Fischer-Gödde et al. (2011) and Wang and Becker (2013).

1243

1244 Fig. 7. a) $\text{Fe}^{3+}/\Sigma\text{Fe}$ ratios in spinel and b) $\log(f_{\text{O}_2}) \Delta\text{FMQ}$ vs Cr# in spinel for samples
1245 from this study. Error bars in a = relative error expressed as % (see Table 2). Error
1246 for $\log(f_{\text{O}_2}) \Delta\text{FMQ}$ includes error related to f_{O_2} sensor (0.5 log units) plus uncertainties
1247 derived from $\text{Fe}^{3+}/\text{Fe}_{\text{total}}$ ratio measurements (see Table 2).

1248 Background fields represent literature data for mantle peridotites: pink= abyssal
1249 peridotites (Bryndzia and Wood, 1990); grey = arc xenoliths (Wood and Virgo, 1989;
1250 Brandon and Draper, 1996; Parkinson et al., 2003); light green and light yellow =
1251 Toroshima and Conical seamount respectively (Parkinson and Pearce, 1998). Modified
1252 after Birner et al. (2017).

1253

1254 Fig. 8. Primitive mantle normalised HSE abundances of a) type-A and b) type-B
1255 harzburgites compared to the HSE composition displayed by New Zealand mantle
1256 xenoliths (Scott et al., 2019). See text for further detail.

1257

Figure 1
[Click here to download high resolution image](#)

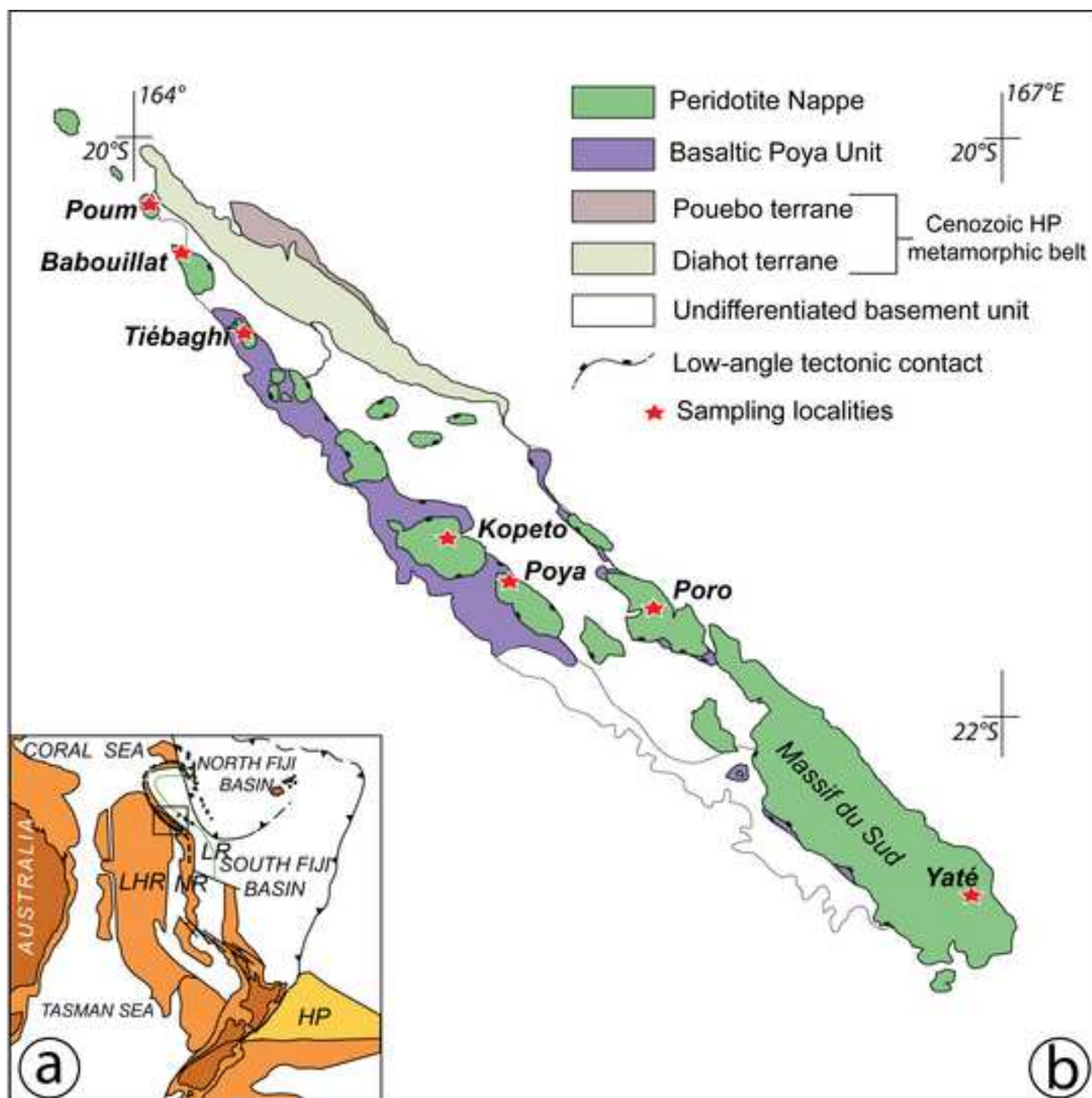


Figure 2

[Click here to download high resolution image](#)

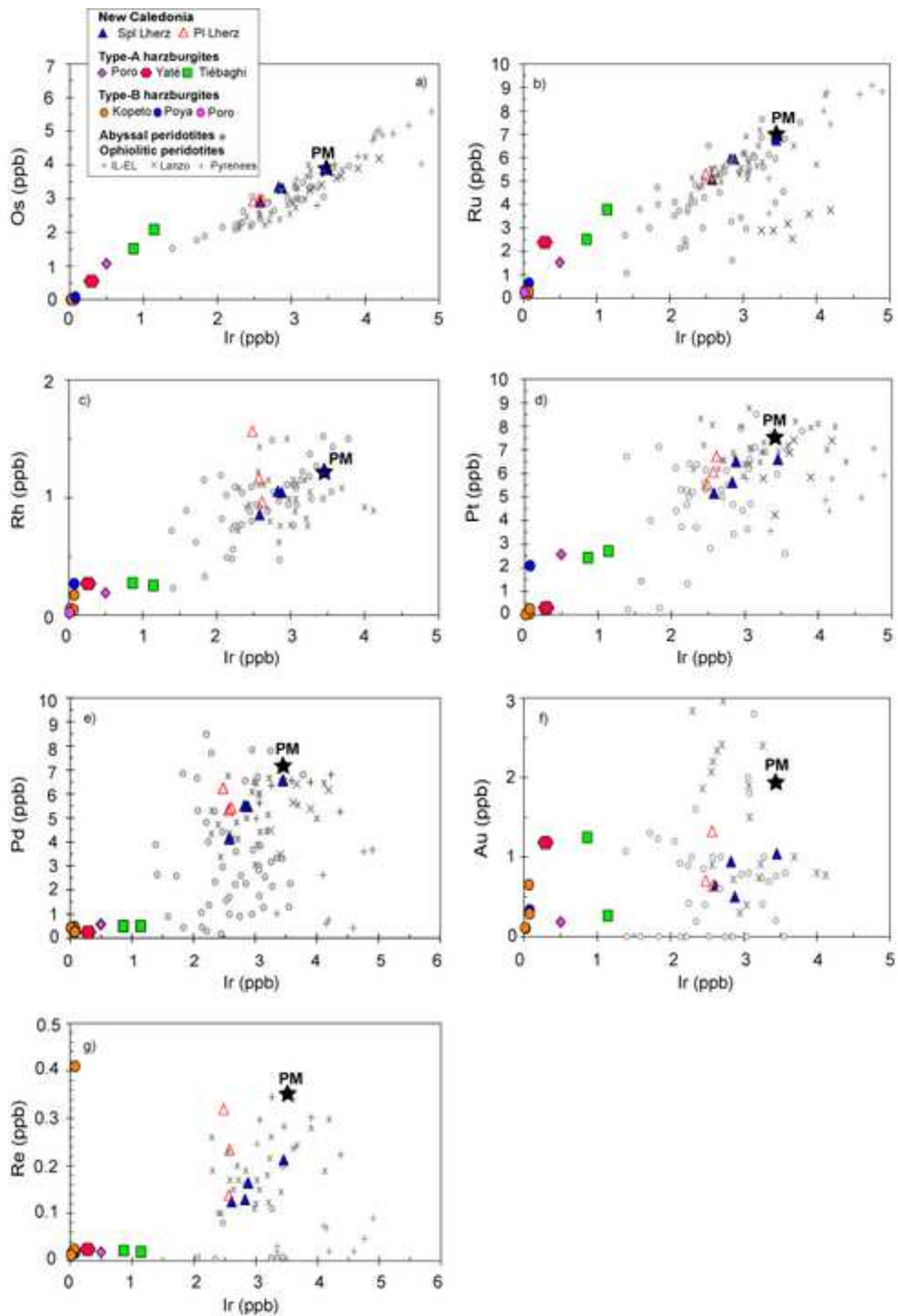


Figure 3
[Click here to download high resolution image](#)

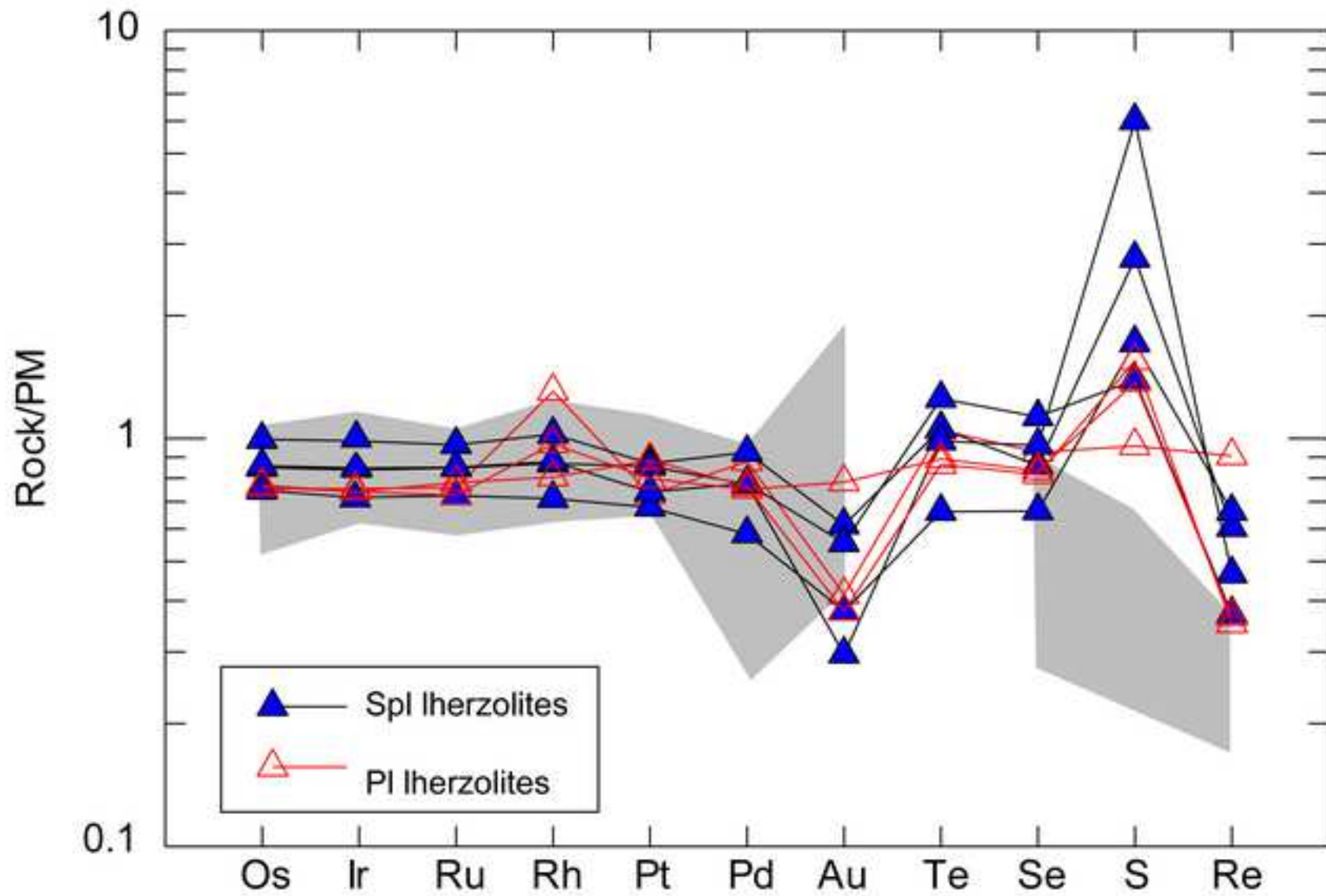


Figure 4

[Click here to download high resolution image](#)

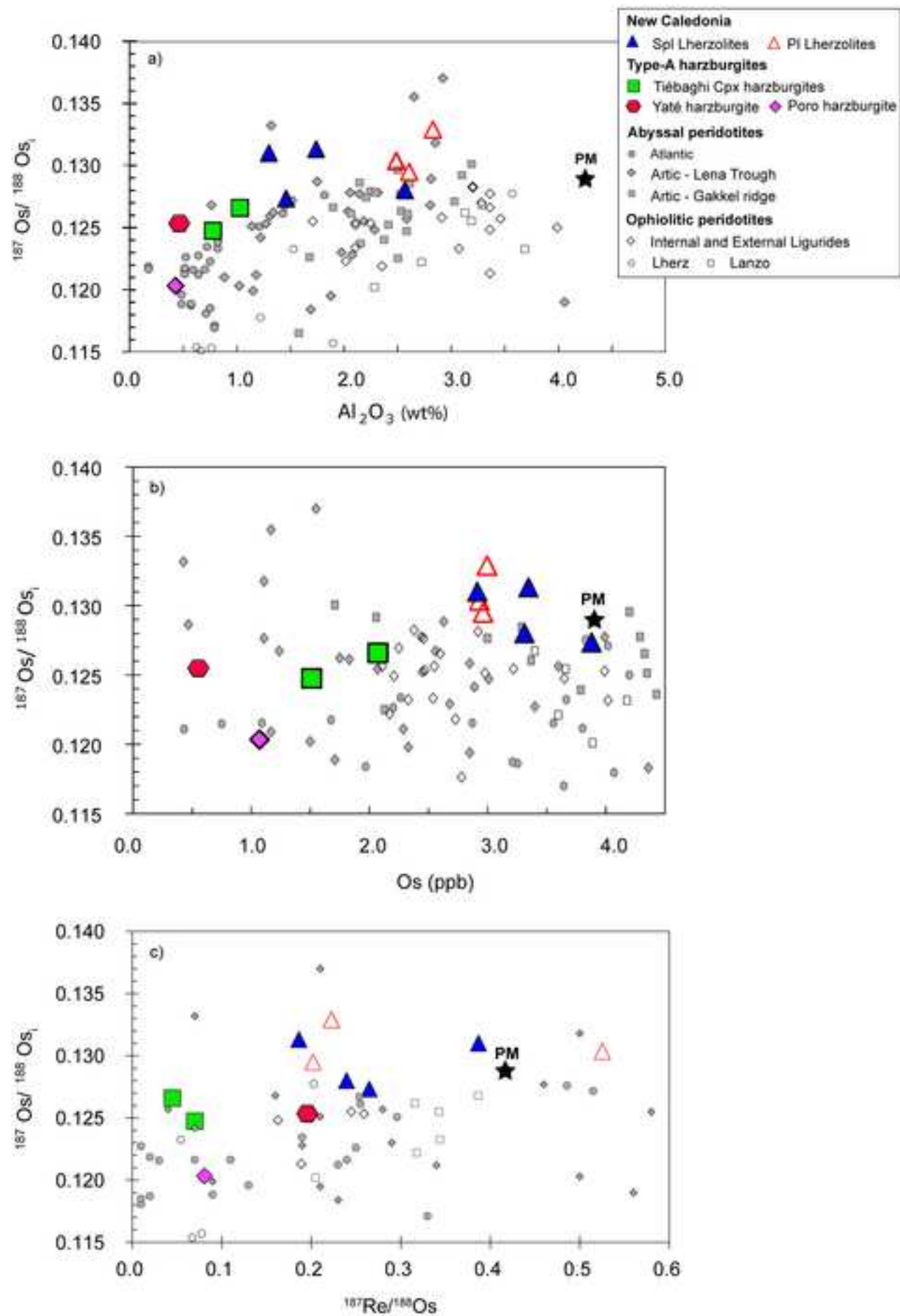


Figure 5

[Click here to download high resolution image](#)

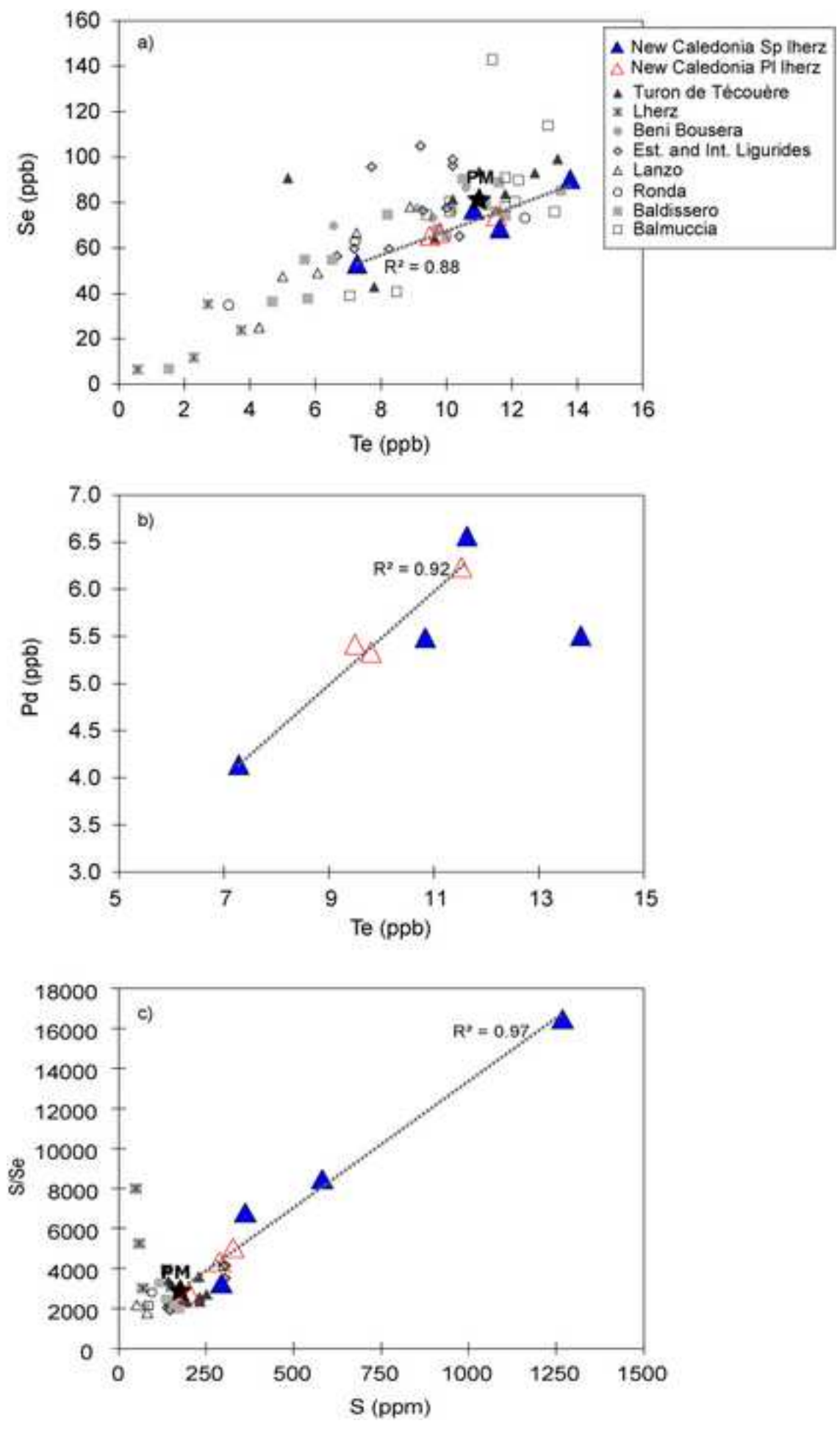


Figure 6
[Click here to download high resolution image](#)

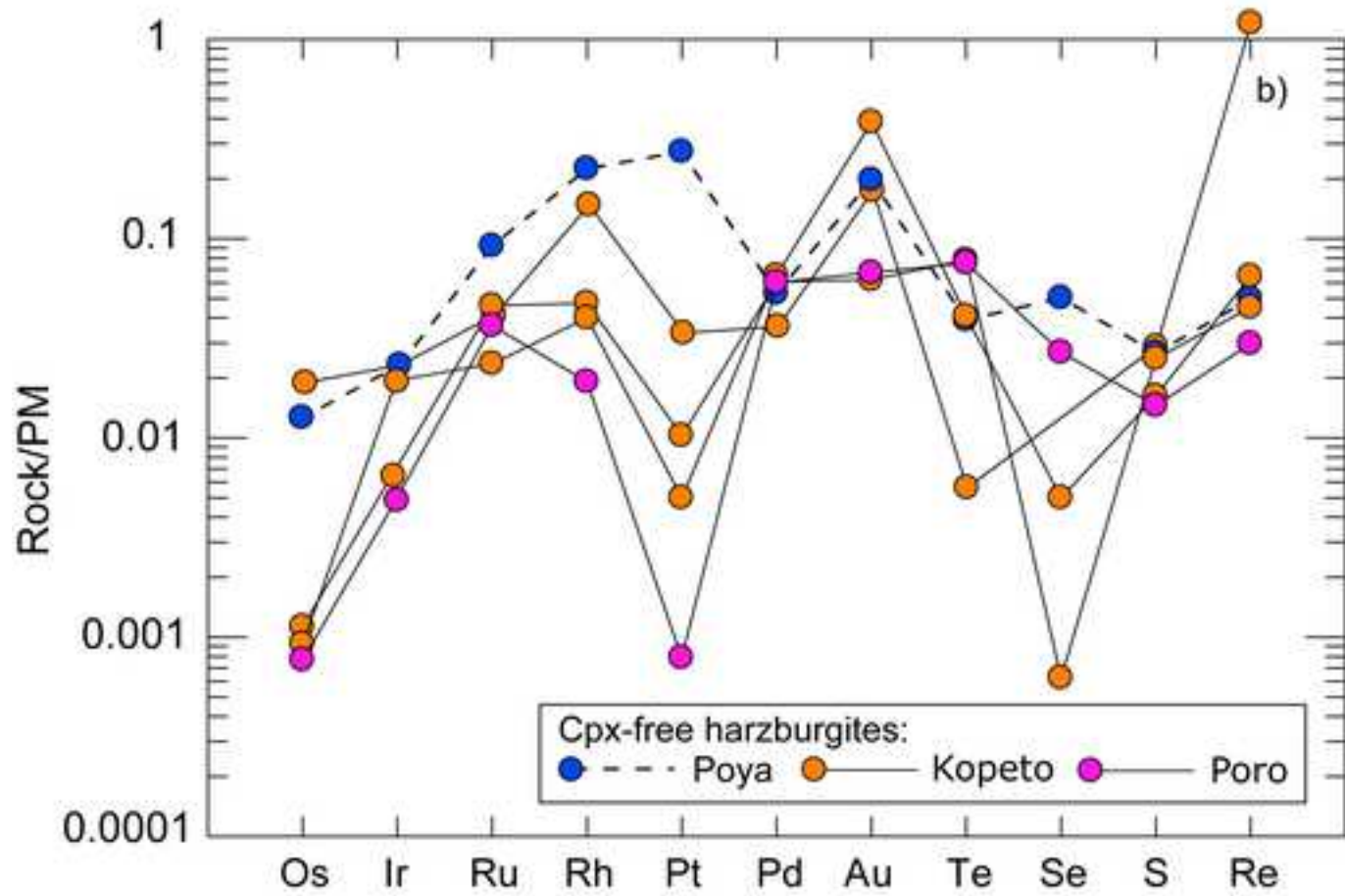
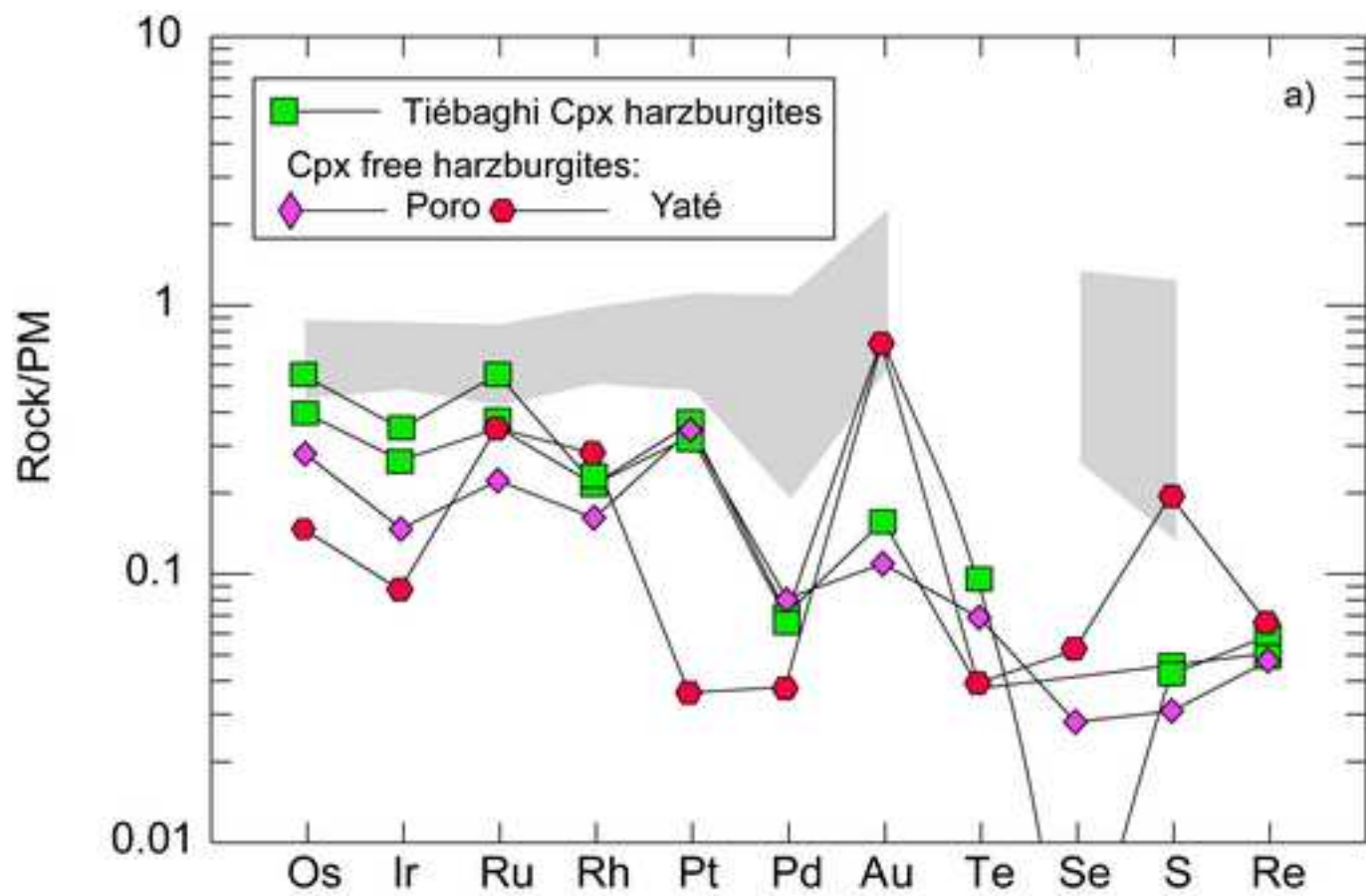


Figure 7
[Click here to download high resolution image](#)

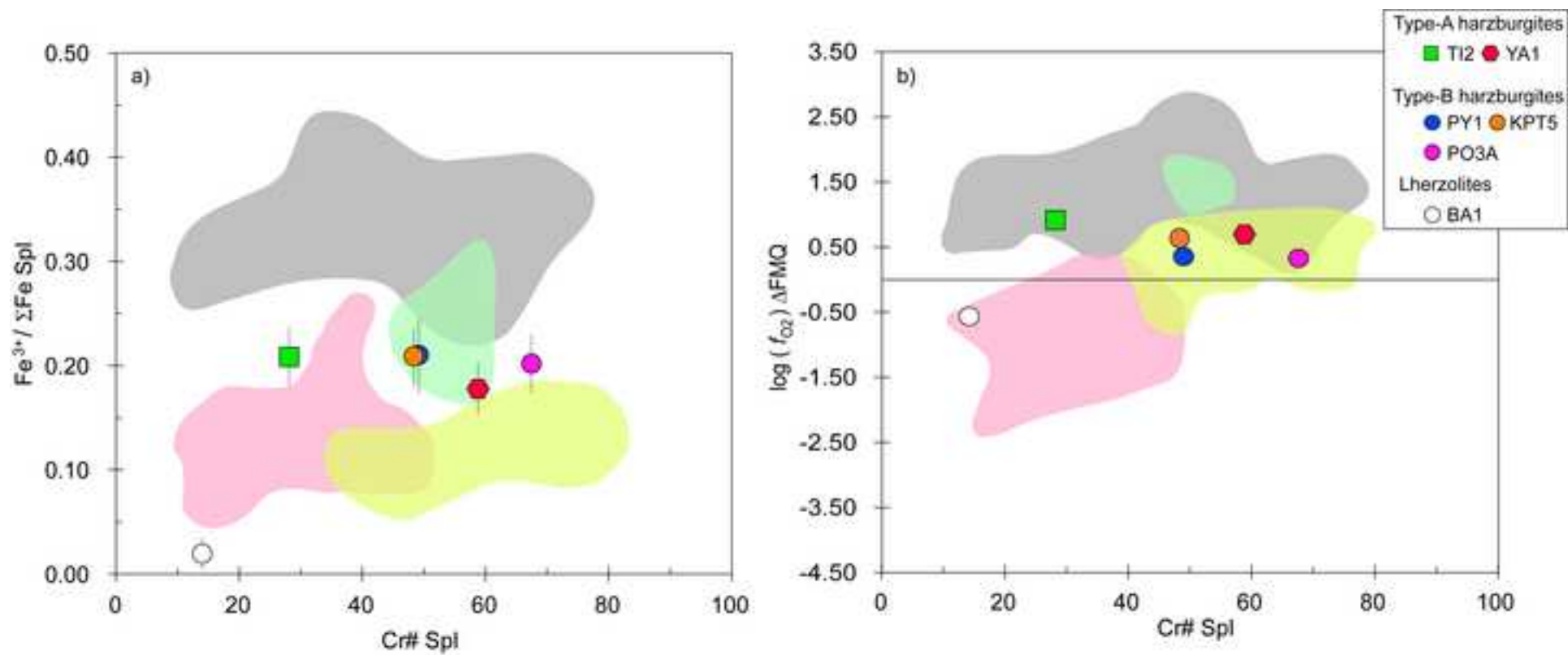
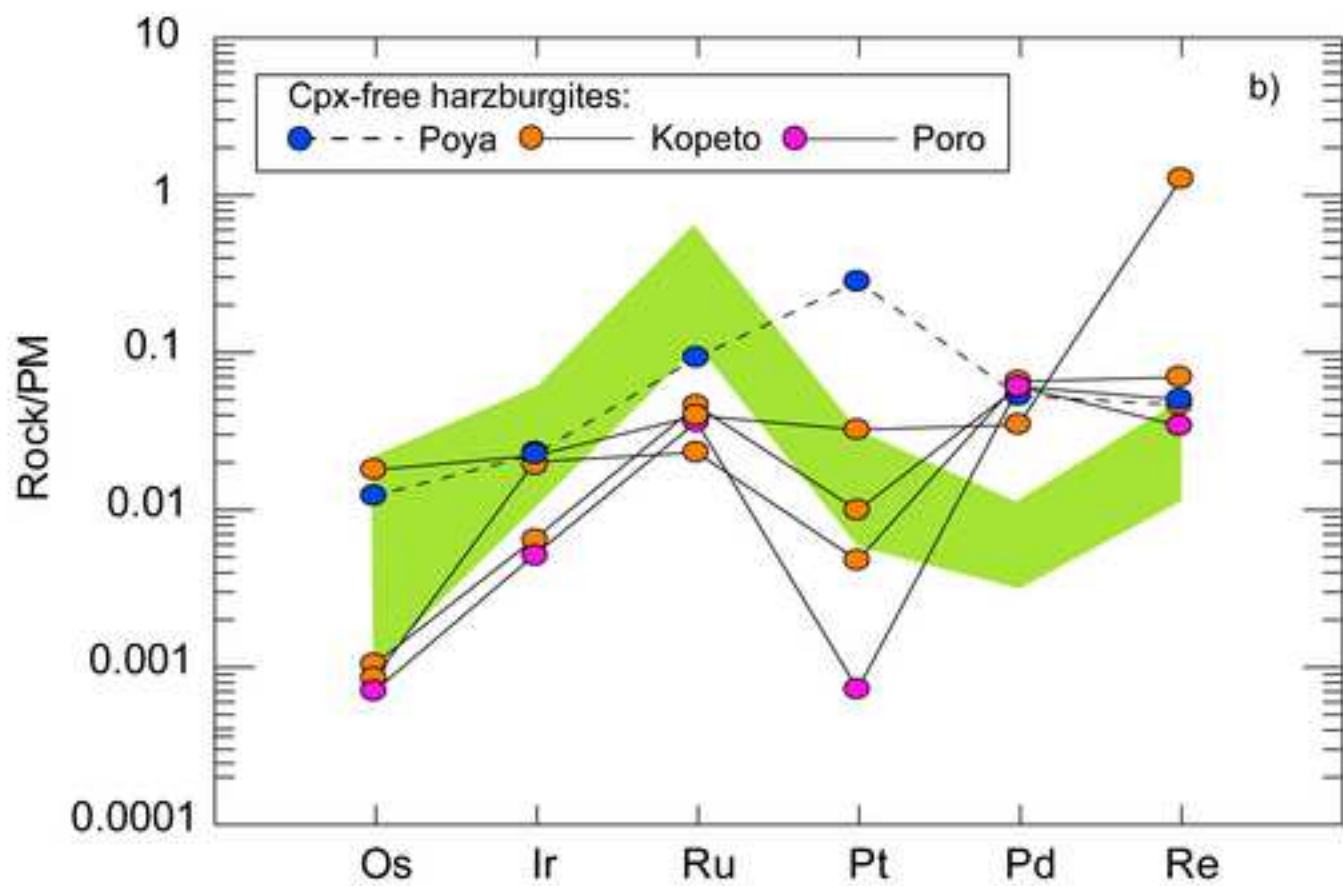
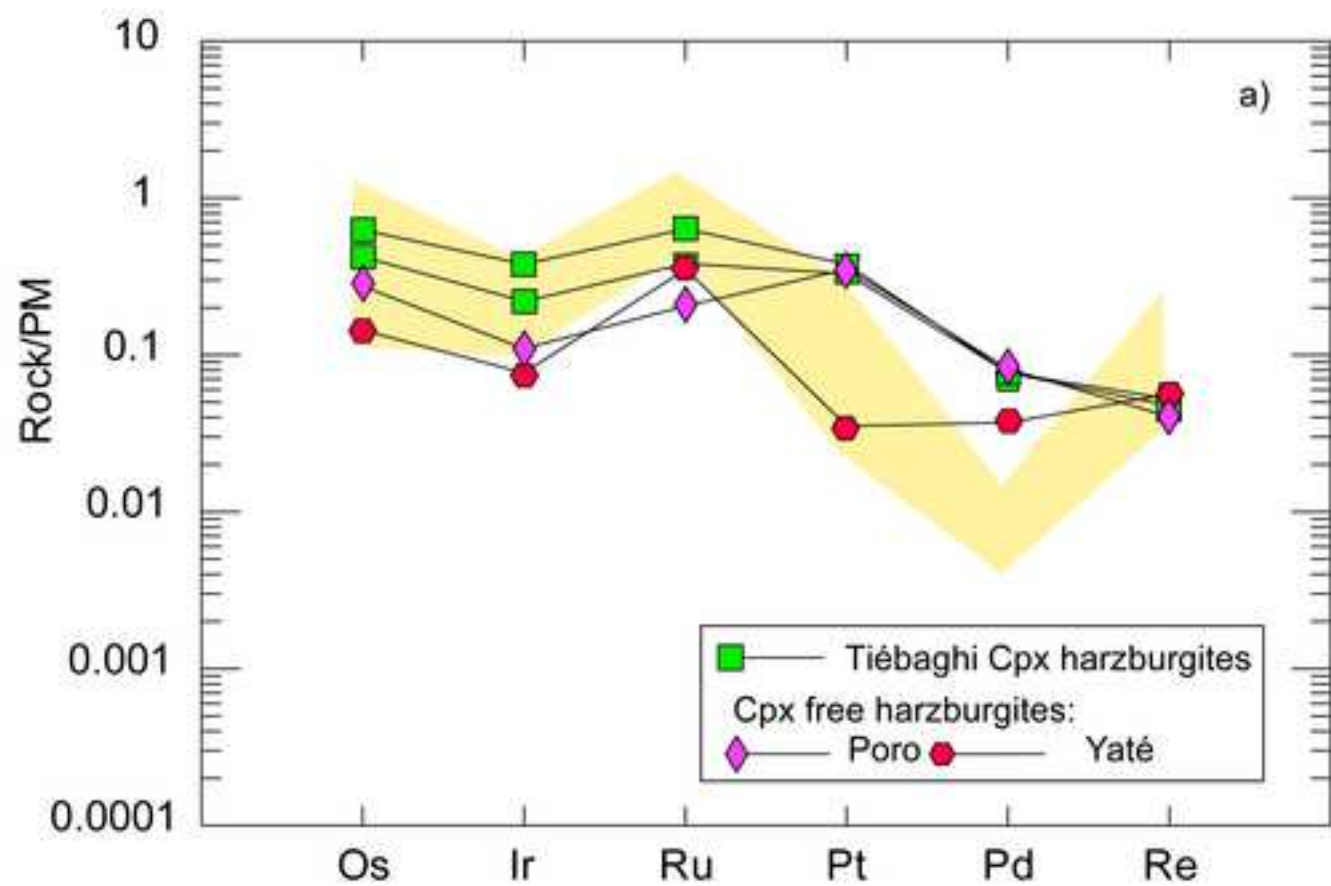
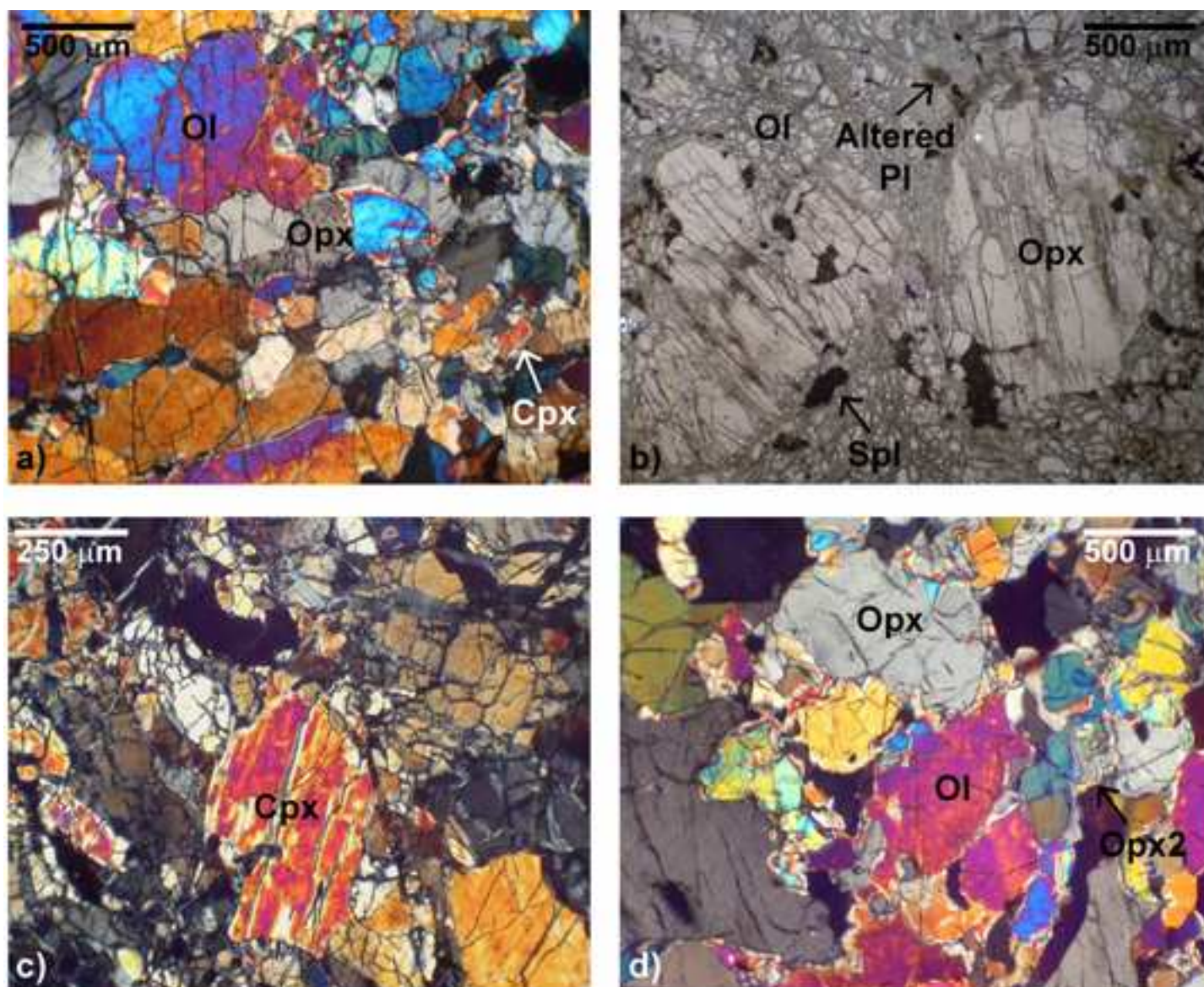
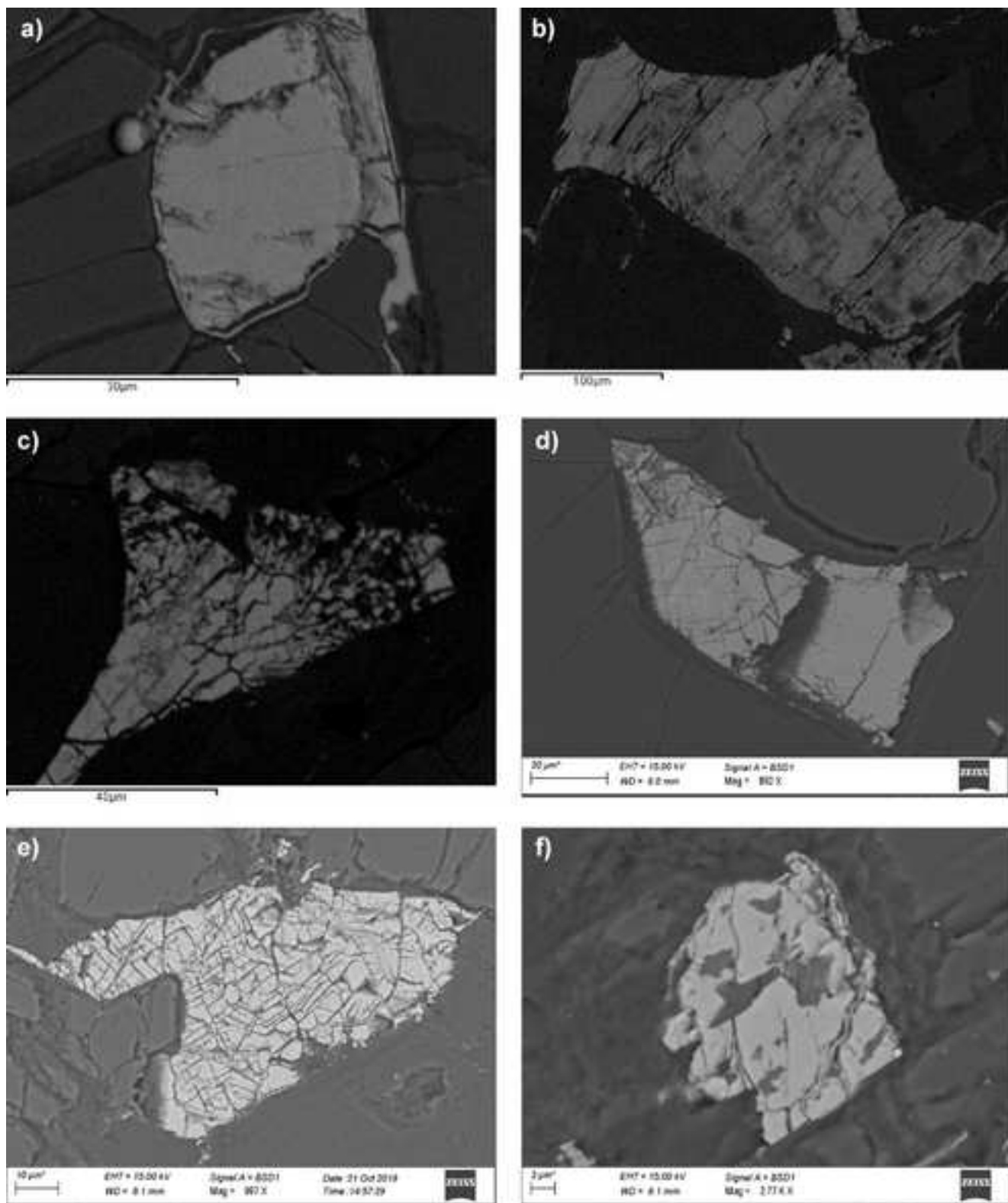


Figure 8
[Click here to download high resolution image](#)





Microphotographs of the New Caledonia peridotites: a) porphyroclastic texture marked by deformed and stretched olivine and orthopyroxene crystals in spl-lherzolite POU2; b) plagioclase-lherzolite thin section (BAB2A, parallel polars); c) strongly exsolved, primary clinopyroxene occurrence in T12 harzburgite; d) interstitial orthopyroxene (opx2) formed at the expense of primary olivine (KPT5 harzburgite).



BSE images of sulphide phases in the New Caledonia ilherzolites: a) small, rounded, interstitial sulphide in BA1 (Mss); b-c) sulphide grains included in porphyroclastic olivine (POU2 ilherzolite, b= pentlandite, c= Mss); d) irregular-shaped, interstitial, Mss grain (BAB1B); e) interstitial sulphide (pentlandite, POU3); f) Mss grain enclosed in olivine (POU1A).

Figure S3

[Click here to download high resolution image](#)

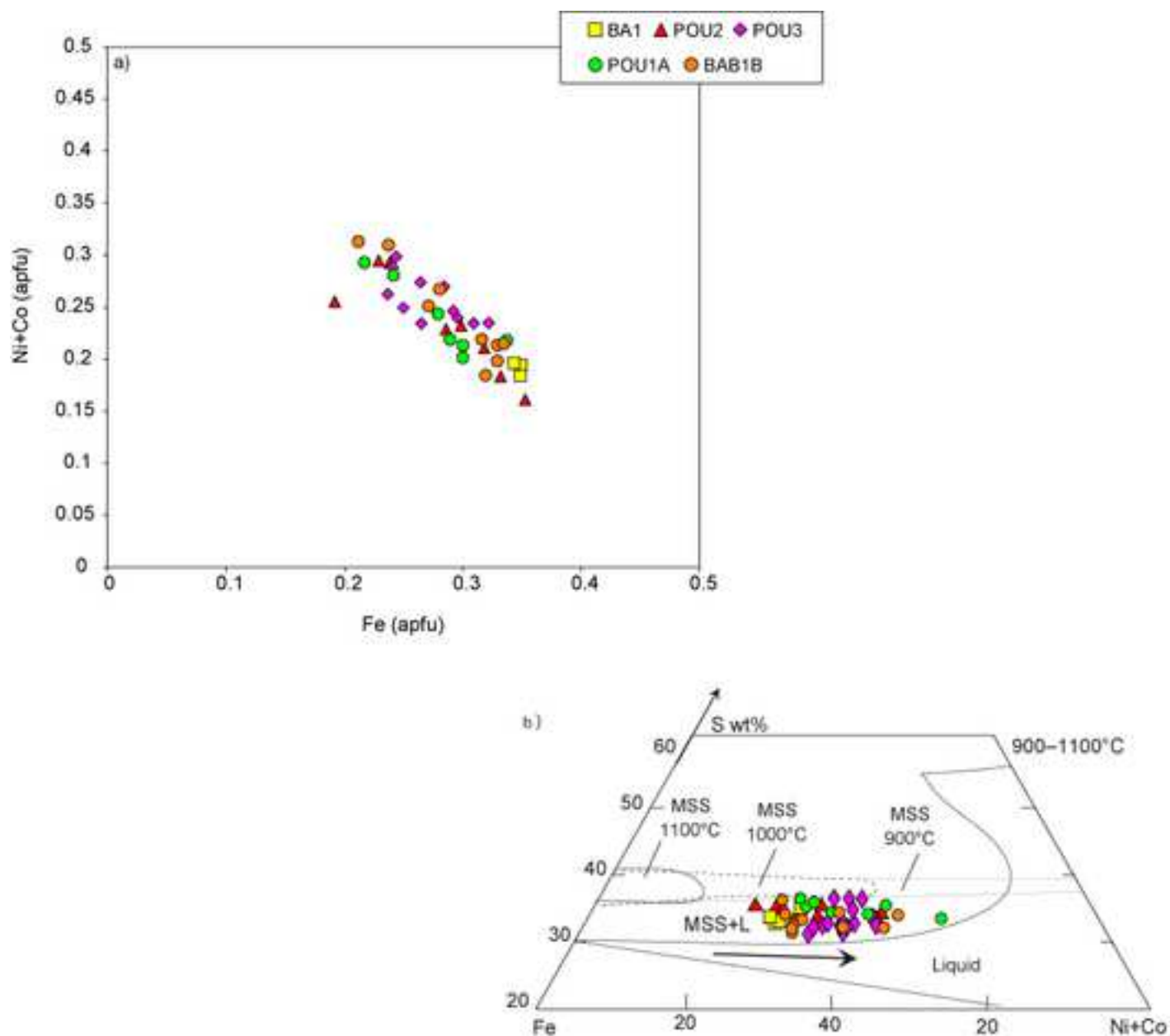
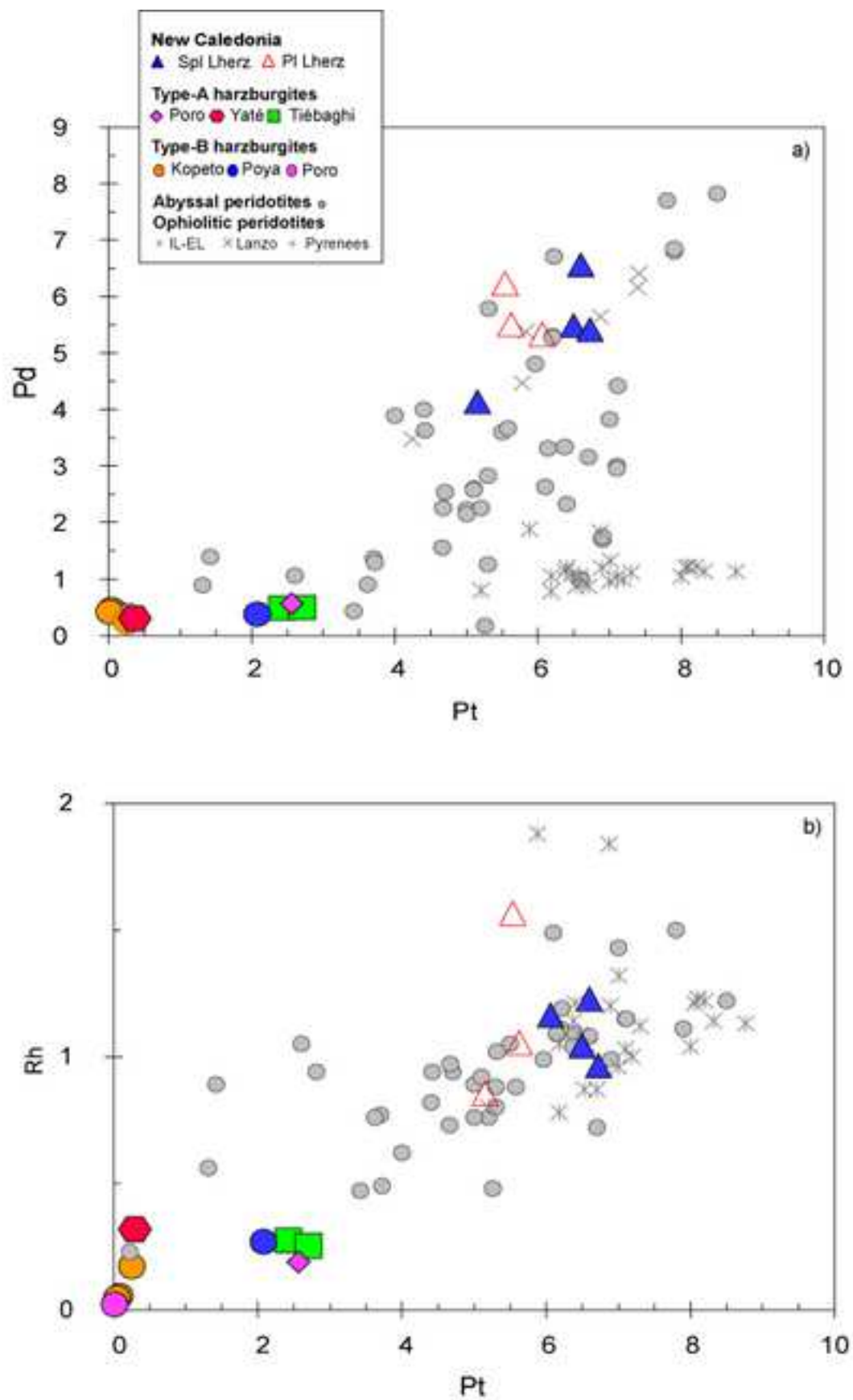


Fig. S3. a) Plot of Fe (at%) v. (Ni + Co) (at %) showing compositional variations of sulphides in the New Caledonia hercynites; b) projection in the Fe-(Ni +Co)-S system of sulphide compositions. Phase relations from Kullerud et al. (1969) and Craig (1973).

Figure S4

[Click here to download high resolution image](#)



Pt vs. a) Pd and b) Rh variation diagrams for the New Caledonia peridotites. Abundances are in ppb.

Table 1
[Click here to download Table: Table 1.xlsx](#)

Table 1
 Concentrations of the HSE, S, Se and Te, Os isotopes, selected major elements and ratios for the New Caledonia peridotites

Sample	Rock	Al ₂ O ₃ wt.%	LOI %	Os (ng/g)	Ir (ng/g)	Ru (ng/g)	Rh (ng/g)	Pt (ng/g)	Pd (ng/g)	Au (ng/g)	Re (ng/g)	Os _N /Ir _N	Os _N /Ru _N	Pd _N /Ir _N	Ru _N /Ir _N	Pt _N /Ir _N	Pt _N /Ru _N	¹⁸⁷ Re/ ¹⁸⁸ Os (2SE)	¹⁸⁷ Os/ ¹⁸⁸ Os measured	2SE	¹⁸⁷ Os/ ¹⁸⁸ Os _i	γOs _(33Ma)	T _{MA} (PM) Ga	T _{RD} (PM) Ga	T _{RD2} (PM) Ga	S (μg/g)	Se (ng/g)	Te (ng/g)	S/Se	Se/Te		
POU1A	PI L	2.49	7.56	2.93	2.48	5.31	1.57	5.54	6.23	0.704	0.319	1.06	0.99	1.24	1.07	1.03	0.96	0.525(1)	0.130822	9.1E-06	0.130358	2.9	0.8	f	f	202	74.9	11.5	2703	6.5		
POU2	Sp L	1.46	8.63	3.88	3.45	6.76	1.23	6.60	6.56	1.04	0.213	1.01	0.18	0.94	5.71	0.88	0.15	0.265(1)	0.127574	6.8E-06	0.127340	0.5	0.7	f	0.3	582	69.6	11.6	8364	6.0		
POU2B	Sp L	2.57	10.69	3.31	2.88	5.93	1.04	6.49	5.49	0.503	0.165	1.03	1.09	0.94	0.95	1.04	1.09	0.240(1)	0.128227	8.0E-06	0.128016	1.1	0.4	f	0.2	1268	77.8	10.8	16289	7.2		
POU3	Sp L	1.30	9.60	2.91	2.58	5.07	0.86	5.15	4.14	0.642	0.234	1.01	1.05	0.79	0.97	0.92	0.95	0.387(1)	0.131356	9.1E-06	0.131015	3.4	f	f	f	362	54.0	7.3	6691	7.4		
BA1	Sp L	1.74	6.39	3.34	2.85	5.95	1.05	5.62	5.51	0.943	0.129	1.06	0.27	0.96	3.88	0.92	0.24	0.186(1)	0.131485	7.9E-06	0.131320	3.7	f	f	f	294	91.3	13.8	3218	6.6		
BAB1B	PI L	2.61	6.98	2.96	2.61	5.46	0.965	6.73	5.42	0.641	0.125	1.02	0.24	1.02	4.30	1.19	0.28	0.202(1)	0.129664	8.9E-06	0.129485	2.2	f	f	0.0	327	66.2	9.5	4943	7.0		
BAB2B	PI L	2.83	8.43	3.00	2.57	5.08	1.16	6.06	5.33	1.33	0.138	1.05	1.67	1.02	0.63	1.09	1.73	0.223(1)	0.133084	8.9E-06	0.132887	4.9	f	f	f	289	67.6	9.8	4278	6.9		
TI1	H	0.78	9.01	1.51	0.864	2.52	0.275	2.42	0.488	1.25	0.022	1.57	1.08	0.28	1.46	1.29	0.89	0.070(25)	0.12479	1.7E-05	0.12473	-1.5	0.8	0.7	0.7	9	1.1	1.0	8188	1.0		
TI2	H	1.03	6.04	2.07	1.14	3.78	0.254	2.70	0.499	0.264	0.019	1.63	0.98	0.21	1.65	1.09	0.66	0.045(18)	0.12662	1.3E-05	0.12658	-0.1	0.5	0.4	0.4	10	0.7	0.4	14053	1.7		
PO4	H	0.43	0.18	1.07	0.499	1.53	0.191	2.56	0.570	0.160	0.018	1.93	1.26	0.56	1.53	2.36	1.54	0.081(35)	0.12040	2.4E-05	0.12033	-5.0	1.5	1.3	1.3	6	3.1	0.8	2084	4.1		
YA1	H	0.46	6.83	0.554	0.297	2.35	0.326	0.268	0.263	1.17	0.022	1.67	0.42	0.44	3.95	0.42	0.11	0.196(68)	0.12551	4.7E-05	0.12534	-1.0	1.0	0.6	0.6	40	4.9	0.4	8117	11.4		
PY1	H	0.78	0.00	0.049	0.077	0.643	0.269	2.08	0.378	0.336	0.016	0.58	0.14	2.43	4.19	2.97	2.97	1.62(8)	0.1299	5.3E-04	0.1284	1.4	0.0	0.0	0.2	6	0.9	0.9	6103	1.1		
Duplicate				0.023	0.133	0.693	0.171	2.44	0.588	0.311	0.062							13.0(1)	0.127	1.0E-03	0.115	-8.9				53	0.3	1.0	196197	0.3		
KPT2	H	0.70	3.03	0.004	0.022	0.323	0.057	0.078	0.431	0.107	0.018	0.18	0.02	9.50	7.22	0.22	0.22	19(1)	0.148	5.6E-03	0.131	3.6	0.1	f	f	5	1.3	0.4	4164	2.8		
Duplicate				0.036	0.017	0.211	0.040	0.210	0.620	0.144	0.010							1.4(1)	0.302	1.3E-03	0.301	137.8				7	1.3	8.2	4947	0.2		
KPT5	H	0.74	0.12	0.004	0.066	0.165	0.048	0.038	0.464	0.654	0.024	0.05	0.04	3.46	1.25	0.21	0.21	32(2)	0.147	6.2E-03	0.118	-6.6	0.0	f	f	3	3.0	0.7	1133	4.1		
Duplicate				0.007	0.062	0.165	0.054	0.046	0.512	0.603	0.056							37(1)	0.160	3.2E-03	0.127	0.5				3	0.7	1.3	4453	0.5		
PO3	H	0.41	0.00	0.003	0.015	0.256	0.023	0.006	0.428	0.084	0.012	0.18	0.02	14.12	8.57	0.02	0.02	19(2)	0.153	8.4E-03	0.136	8.0	0.1	f	f	3	1.4	0.6	2129	2.2		
Duplicate				0.055	0.018	0.243	0.016	0.009	0.573	0.293	0.0078							0.69(7)	0.1239	4.7E-04	0.1233					bdl	bdl	0.8	-	-		
KPT3	H	0.67	0.67	0.072	0.079	0.280	0.172	0.248	0.250	0.288	0.410	0.81	1.87	0.17	0.19	6.92	6.92	28(1)	0.1273	3.5E-04	0.1204	-5.1	0.0	0.3	1.3	6	bdl	bdl	-	-		

Duplicate: replicate digestion of the same sample powder

¹L= plagioclase lherzolite, Sp L= spinel lherzolite, H= harzburgite

Values of PM ¹⁸⁷Os/¹⁸⁸Os = 0.1296 and ¹⁸⁷Re/¹⁸⁸Os = 0.434 used for calculation of T_{MA} and T_{RD} ages (Meisel et al., 2001); f= future model ages. T_{RD2}(PM) indicates depletion ages calculated taking into account Re addition that may have occurred during peridotite evolution.

Table 2[Click here to download Table: Table 2.xlsx](#)**Table 2**

Equilibration temperature, pressure, and oxygen fugacity calculated for selected peridotite samples. Average spinel compositions are also reported.

Sample	Type	Fe ³⁺ /ΣFe ¹	Fe ³⁺ /Σfe (rel. err.%) ²	T Ol-Spl (°C) ³	P (GPa)	Wood ⁴ (1990)	
						log f(O ₂)	ΔFMQ
TI2	Type-A harz	0.21 (0.03)	15.0	870	1.0	-11.37	0.92
YA1	Type-A harz	0.18 (0.03)	14.5	815	1.0	-12.66	0.67
KPT5	Type-B harz	0.21 (0.03)	16.0	840	1.0	-12.20	0.65
PO3A	Type-B harz	0.20 (0.03)	14.8	940	1.0	-10.76	0.33
PY1B	Type-B harz	0.21 (0.03)	16.6	930	1.0	-10.90	0.35

Sample	Type	Spinel											
		SiO ₂	TiO ₂	Al ₂ O ₃	Cr ₂ O ₃	FeO	Fe ₂ O ₃	MnO	MgO	Total	Cr#	Mg#	Lβ/Lα ⁵
TI2	Type-A harz	0.04	0.03	43.45	25.35	10.20	2.99	0.15	16.93	99.13	0.281	0.747	0.804
YA1	Type-A harz	0.02	0.01	21.56	45.88	16.76	4.03	0.29	10.81	99.35	0.588	0.535	0.896
KPT5	Type-B harz	0.01	0.01	29.66	41.41	11.31	3.31	0.21	14.41	100.33	0.484	0.695	0.825
PO3A	Type-B harz	0.02	0.00	17.27	53.48	12.58	3.54	0.25	12.32	99.47	0.675	0.636	0.833
PY1B	Type-B harz	0.01	0.01	29.35	41.87	11.20	3.32	0.20	14.42	100.38	0.492	0.696	0.826

¹ Fe³⁺/(Fe²⁺+Fe³⁺) in spinel determined by "flank"-method.Fe³⁺/(Fe²⁺+Fe³⁺) ratio has been calculated using the equation : Lα/Lβ x 0.42+ Cr# x 0.03 - (FeO + (Fe₂O₃/1.113)) x 0.01 obtained after standardisation.² Relative error in % calculated as a total of measurement errors for Fe, Cr, Al and X-ray intensities at flank positions.³ T Ol-Spl = Equilibration temperature calculated using Li et al. (1995) formulation.⁴ Oxygen fugacity estimates are from Wood et al. (1990) and are reported as Δlog fO₂ from the quartz-fayalite-magnetite (FMQ) buffer using the Fe³⁺/ΣFe values of the analysed spinels. See text for further details.⁵ Current normalized ratio of the intensities at flank position for Lα and Lβ Fe lines

Table S1[Click here to download Table: Table S1.xlsx](#)**Table S1**

Major element composition for selected sulphides from the New Caledonia lherzolites

Sample Mineral Occurrence	POU2 <i>mss</i> <i>Encl</i>	POU2 <i>mss</i> <i>Interst</i>	POU2 <i>mss</i> <i>Encl</i>	POU2 <i>pn</i> <i>Encl</i>	POU2 <i>pn</i> <i>Interst</i>	POU2 <i>pn</i> <i>Interst</i>	POU2 <i>pn</i> <i>Interst</i>	BA1 <i>mss</i> <i>Interst</i>	BA1 <i>mss</i> <i>Interst</i>	BA1 <i>mss</i> <i>Interst</i>	BA1 <i>mss</i> <i>Encl</i>	BA1 <i>mss</i> <i>Encl</i>
<i>wt%</i>												
S	34.32	34.39	32.74	33.24	41.19	32.68	40.82	33.22	33.08	32.47	32.53	33.13
Fe	41.23	43.80	36.65	28.09	23.90	29.14	24.68	36.85	42.99	43.25	42.19	43.02
Co	0.64	0.59	0.78	0.46	1.93	1.91	2.26	0.46	0.50	0.55	0.80	0.58
Ni	23.28	20.44	29.19	37.55	32.99	35.74	32.23	25.77	23.44	24.72	24.54	23.27
Cu	0.52	0.57	0.64	0.66	bdl	0.53	bdl	3.70	bdl	bdl	bdl	bdl
Total	99.99	99.79	100.00	100.00	100.01	100.00	99.99	100.00	100.01	100.99	100.06	100.00
Fe/Ni	1.8	2.1	1.3	0.7	0.7	0.8	0.8	1.4	1.8	1.7	1.7	1.8

Interst. =interstitial; encl = enclosed

Pn= pentlandite; mss= monosulphide solid solution

Table S2[Click here to download Table: Table S2.xlsx](#)

Average major element composition of olivine from the analysed harzburgites

Sample	Type	SiO ₂	FeO	MnO	MgO	NiO	Total	Mg#
TI2	Type-A harz	41.79	7.96	0.25	49.88	0.51	100.39	0.918
YA1	Type-A harz	41.39	8.14	0.24	50.03	0.37	100.16	0.914
KPT5	Type-B harz	41.83	7.54	0.23	50.01	0.47	100.07	0.923
PO3A	Type-B harz	41.56	8.60	0.08	49.54	0.41	100.19	0.914
PY1B	Type-B harz	41.34	8.26	0.09	50.25	0.33	100.27	0.915

Average major element composition of orthopyroxene from the analysed harzburgites

Sample	Type	SiO ₂	Al ₂ O ₃	Cr ₂ O ₃	FeO	MnO	MgO	CaO
TI2	Type-A harz	55.69	4.13	0.87	6.03	0.18	32.93	0.56
YA1	Type-A harz	57.51	1.42	0.76	5.01	0.09	34.32	0.90
KPT5	Type-B harz	57.16	1.91	0.71	5.17	0.18	34.01	0.88
PO3A	Type-B harz	57.98	0.94	0.44	5.70	0.25	34.88	0.35
PY1	Type-B harz	57.22	2.37	0.96	5.47	0.10	33.63	0.71