QUBIC VII: The feedhorn-switch system of the technological demonstrator

F. Cavaliere^{1,2} A. Mennella^{1,2} M. Zannoni^{3,4} P. Battaglia⁵ E.S. Battistelli^{6,7} P. de Bernardis^{6,7} D. Burke⁸ G. D'Alessandro^{6,7} M. De Petris^{6,7} C. Franceschet^{1,2} L. Grandsire⁹ J.-Ch. Hamilton⁹ B. Maffei¹⁰ E. Manzan¹ S. Marnieros¹¹ S. Masi^{6,7} C. O'Sullivan⁸ A. Passerini^{3,4} F. Pezzotta^{1,2} M. Piat⁹ A. Tartari¹² S.A. Torchinsky^{9,13} D. Viganò^{1,2} F. Voisin⁹ P. Ade¹⁴ J.G. Alberro¹⁵ A. Almela¹⁶ G. Amico⁶ L.H. Arnaldi¹⁷ D. Auguste¹¹ J. Aumont¹⁸ S. Azzoni¹⁹ S. Banfi^{3,4} A. Baù^{3,4} B. Bélier²⁰ D. Bennett⁸ L. Bergé¹¹ J.-Ph. Bernard¹⁸ M. Bersanelli^{1,2} M.-A. Bigot-Sazy⁹ J. Bonaparte²¹ J. Bonis¹¹ E. Bunn²² D. Buzi⁶ P. Chanial⁹ C. Chapron⁹ R. Charlassier⁹ A.C. Cobos Cerutti¹⁶ F. Columbro^{6,7} A. Coppolecchia^{6,7} G. De Gasperis^{23,24} M. De Leo^{6,25} S. Dheilly⁹ C. Duca¹⁶ L. Dumoulin¹¹ A. Etchegoyen¹⁶ A. Fasciszewski²¹ L.P. Ferreyro¹⁶ D. Fracchia¹⁶ M.M. Gamboa Lerena^{26,27} K.M. Ganga⁹ B. García¹⁶ M.E. García Redondo¹⁶ M. Gaspard¹¹ D. Gayer⁸ M. Gervasi^{3,4} M. Giard¹⁸ V. Gilles^{6,28} Y. Giraud-Heraud⁹ M. Gómez Berisso¹⁷ M. González¹⁷ M. Gradziel⁸ M.R. Hampel¹⁶ D. Harari¹⁷ S. Henrot-Versillé¹¹ F. Incardona^{1,2} E. Jules¹¹ J. Kaplan⁹ C. Kristukat²⁹ L. Lamagna^{6,7} S. Loucatos^{9,30} T. Louis¹¹ W. Marty¹⁸ A. Mattei⁷ A. May²⁸ M. McCulloch²⁸ L. Mele^{6,7} D. Melo¹⁶ L. Montier¹⁸ L. Mousset⁹ L.M. Mundo¹⁵ J.A. Murphy⁸ J.D. Murphy⁸ F. Nati^{3,4} E. Olivieri¹¹ C. Oriol¹¹ A. Paiella^{6,7} F. Pajot¹⁸ H. Pastoriza¹⁷ A. Pelosi⁷ C. Perbost⁹ M. Perciballi⁷ F. Piacentini^{6,7} L. Piccirillo²⁸ G. Pisano¹⁴ M. Platino¹⁶ G. Polenta^{6,31} D. Prêle⁹ R. Puddu³² D. Rambaud¹⁸ E. Rasztocky³³ P. Ringegni¹⁵ G.E. Romero³³ J.M. Salum¹⁶ A. Schillaci^{6,34} C.G. Scóccola^{26,27} S. Scully^{8,35} S. Spinelli³ G. Stankowiak⁹ M. Stolpovskiy⁹ A.D. Supanitsky¹⁶ J.-P. Thermeau⁹ P. Timbie³⁶ M. Tomasi^{1,2} C. Tucker¹⁴ G. Tucker³⁷ N. Vittorio²³ F. Wicek¹¹ M. Wright²⁸ and A. Zullo⁷

- ¹Università degli studi di Milano, Milano, Italy
- ²INFN sezione di Milano, 20133 Milano, Italy
- ³Università di Milano Bicocca, Milano, Italy
- $^4\mathrm{INFN}$ sezione di Milano Bicocca, 20216 Milano, Italy
- $^5\mathrm{INAF}/\mathrm{OAS}$ Milano Section
- 6 Università di Roma La Sapienza, Roma, Italy
- $^7\mathrm{INFN}$ sezione di Roma, 00185 Roma, Italy
- ⁸National University of Ireland, Maynooth, Ireland
- ⁹Université de Paris, CNRS, Astroparticule et Cosmologie, F-75006 Paris, France
- ¹⁰Institut d'Astrophysique Spatiale, Orsay (CNRS-INSU), France
- ¹¹Laboratoire de Physique des 2 Infinis Irène Joliot-Curie (CNRS-IN2P3, Université Paris-Saclay), France
- $^{12}\mathrm{INFN}$ sezione di Pisa, 56127 Pisa, Italy
- ¹³Observatoire de Paris, Université Paris Science et Lettres, F-75014 Paris, France
- $^{14}\mathrm{Cardiff}$ University, UK
- ¹⁵GEMA (Universidad Nacional de La Plata), Argentina
- 16 Instituto de Tecnologías en Detección y Astropartículas (CNEA, CONICET, UNSAM), Argentina
- $^{17}\mathrm{Centro}$ Atómico Bariloche and Instituto Balseiro (CNEA), Argentina
- ¹⁸Institut de Recherche en Astrophysique et Planétologie, Toulouse (CNRS-INSU), France
- ¹⁹Department of Physics, University of Oxford, UK
- ²⁰Centre de Nanosciences et de Nanotechnologies, Orsay, France
- ²¹Centro Atómico Constituyentes (CNEA), Argentina
- ²²University of Richmond, Richmond, USA
- ²³Università di Roma "Tor Vergata", Roma, Italy
- ²⁴INFN sezione di Roma2, 00133 Roma, Italy
- ²⁵University of Surrey, UK
- ²⁶Facultad de Ciencias Astronómicas y Geofísicas (Universidad Nacional de La Plata), Argentina
- $^{27}\mathrm{CONICET},$ Argentina
- $^{28}\mathrm{University}$ of Manchester, UK
- ²⁹Escuela de Ciencia y Tecnología (UNSAM) and Centro Atómico Constituyentes (CNEA), Argentina
- ³⁰IRFU, CEA, Université Paris-Saclay, F-91191 Gif-sur-Yvette, France
- ³¹Italian Space Agency, Roma, Italy
- ³²Pontificia Universidad Catolica de Chile, Chile
- ³³Instituto Argentino de Radioastronomía (CONICET, CIC, UNLP), Argentina
- ³⁴California Institute of Technology, USA
- ³⁵Institute of Technology, Carlow, Ireland
- ³⁶University of Wisconsin, Madison, USA
- ³⁷Brown University, Providence, USA

E-mail: aniello.mennella@fisica.unimi.it, mario.zannoni@unimib.it

Abstract. We present the design, manufacturing and performance of the horn-switch system developed for the technological demonstrator of QUBIC (the Q&U Bolometric Interferometer for Cosmology). This system consists of 64 back-to-back dual-band (150 GHz and 220 GHz) corrugated feed-horns interposed with mechanical switches used to select desired baselines during the instrument self-calibration. We manufactured the horns in aluminum platelets milled by photo-chemical etching and mechanically tightened with screws. The switches are based on steel blades that open and close the waveguide between the back-to-back horns and are operated by miniaturized electromagnets. The measured electromagnetic performance of the feedhorns agrees with simulations. In particular we obtained a return loss around -20 dB up to 230 GHz and beam patterns in agreement with single-mode simulations down to -30 dB. The switches for this prototype were designed and built for the 150 GHz band. In this frequency range we find return and insertion losses consistent with expectations (< -25 dB and $\sim -0.1 \text{ dB}$, respectively) and an isolation larger than 70 dB. In this paper we also show the current development status of the feedhorn-switch system for the QUBIC full instrument, based on an array of 400 horn-switch assemblies.

Keywords: CMB experiments, CMB polarization, feedhorn arrays, microwave propagation

Contents

1	Introduction			
2	The QUBIC Instrument			
3	The	e horn-switch system	5	
4	Back-to-back feed horns			
	4.1	Feed horns requirements and design	5	
		4.1.1 Requirements	5	
		4.1.2 Mechanical and electromagnetic design	5	
		4.1.3 Electromagnetic simulations technique	7	
		4.1.4 Simulations results	8	
	4.2	The technological demonstrator feed horn array	10	
	4.3	Mechanical measurements and achieved tolerance	11	
		4.3.1 Experimental procedures	11	
		4.3.2 Impact of mechanical imperfections on electromagnetic performance	11	
		4.3.3 Results of metrological measurements	14	
	4.4	RF characterization	16	
		4.4.1 Experimental setup and procedures	17	
		4.4.2 Results	19	
	4.5	The full instrument feed horn array	20	
5	Swi	tch system	21	
	5.1	Switch requirements and design	21	
	5.2	Single channel prototype	27	
		5.2.1 Switch electromagnetic measurements	28	
	5.3	The technological demonstrator switch array	29	
		5.3.1 Switch manufacturing and control electronics	29	
		5.3.2 Switch cryogenic tests	31	
	5.4	The full instrument switch array	34	
6	Cor	clusions	34	

1 Introduction

This paper is part of a set describing the current status of QUBIC (Q and U Bolometric Instrument for Cosmology), an experiment based on the concept of bolometric interferometry [1, 2] that will observe the sky at 150 GHz and 220 GHz from the Argentinean Alto Chorrillo site and is designed to tightly constrain the *B*-mode polarization anisotropies of the Cosmic Microwave Background (CMB).

In the landscape of CMB polarization experiments, QUBIC is currently the only instrument based on this concept that combines the sensitivity of transition edge sensor (TES) bolometers with the advantages of interferometers. The combination gives features including self-calibration, in-band spectral resolution, which we call spectral imaging [3, 4], and interferometric rather than photometric instrumental systematic effects. In this paper we describe the design, manufacturing and testing of the horn-switch array developed for the QUBIC technological demonstrator (TD), a prototype instrument that will demonstrate the application of bolometric interferometry to CMB polarization measurements with dedicated sky observations carried out from the QUBIC site.

As shown in figure 1, each horn in the array is connected to its twin with a straight circular waveguide section that can be kept open or closed by mean of a mechanical switch. When the instrument is in calibration mode it observes an artificial source in the far field and we can close subsets of horn pairs to select only certain baselines. This particular calibration scheme, that we call *self-calibration*, allows us to determine the instrument parameters with high precision thanks to the redundancy provided by the large number of available baselines. The interested reader can find all the details of the self-calibration concept applied to QUBIC in Bigot-Sazy et al. [5].

The QUBIC TD horn-switch system has 64+64 back-to-back horns and 64 switches, while in the QUBIC full instrument (FI) this number is raised to 400 (i.e., 400+400 horns and 400 switches). In this paper we describe in detail the manufacturing and testing of the horn-switch system of the QUBIC TD and present briefly the advances of the development of the array that will be installed in the QUBIC FI.

The main challenge in this work has been to realize large arrays of feed-horns and switches with a cost-effective and scalable approach. Concerning feed-horns, we have found the optimal solution in the platelet technique combined with chemical etching. This technique, first introduced by Haas et al. in 1993 [6, 7], allows one to build feed-horns by stacking metal plates that are drilled either by mechanical tools or by chemical etching. Several examples of the application of the platelet technique in the field of CMB polarization experiments can be found in the literature [8–14]. The waveguide switches are realized by insertion of a metallic blade in the guide connecting the back-to-back horns. The preliminary challenge is presented by realizing a switch mechanism capable of moving the shutters at cryogenic temperature in a reliable manner. Since we did not find similar devices described in the literature, we carried out a completely new research and development effort.

In our paper we start with a top-level description of the system in section 3, then we present the back-to-back horn system in section 4 and the switch array in section 5. In each section we review the main requirements, discuss the manufacturing and testing techniques, present the testing results and provide an overview of the development of the arrays for the QUBIC FI. Finally, section 6 summarizes the conclusions and future prospects of our work.

2 The QUBIC Instrument

QUBIC employs an optical system consisting of back-to-back horns that select the relevant baselines and an optical combiner focusing on a bolometric focal plane. The optical combiner forms interference fringes while the bolometers average their powers over timescales much larger than the period of the light waves. This is therefore the optical equivalent of a wide-band correlator in classical interferometry. The instrument operates at cryogenic temperatures thanks to a large cryostat described in Masi et al. [15].



Figure 1. Top: schematic of the QUBIC instrument. The filters in front of the half-wave plate are infrared filters described in table 1 of Masi et al. [15]. The cold stop is the septum between the two mirrors. *Middle*: sectional cut of the cryostat showing the same sub-systems in their actual accommodation. *Bottom*: detailed schematic of a pair of back-to-back horns connected by a circular waveguide that can be shut by the switch blade (shown by the gray rectangle). Notice that the blade is not placed in the middle of the waveguide section. The reason lies in the mechanical layout of the switch system (see section 5.1, figure 23).

Parameter	QUBIC-TD	QUBIC-FI
Frequency channels	$150 \mathrm{GHz}$	150 GHz & 220 GHz
Frequency range 150 GHz	[131-169] GHz	[131-169] GHz
Frequency range 220 GHz	-	[192.5-247.5] GHz
Window Aperture [m]	0.56	0.56
Number of horns	64	400
Number of detectors	248	992×2
Detector noise $[W/\sqrt{Hz}]$	2.05×10^{-16}	4.7×10^{-17}
Focal plane temp. [mK]	300	300
Sky Coverage	1.5%	1.5%
Synthesized beam FWHM $[{\rm degrees}]_{\cdot \cdot}$	0.68	0.39 (150 GHz), 0.27 (220 GHz)

 Table 1. QUBIC main parameters

A schematic of the design of QUBIC is shown in figure 1 and the main instrument parameters are listed in table 1. The sky signal first goes through a 56 cm diameter, 25 mm thick window made of Ultra-High-Molecular-Weight Polyethylene followed by a series of thermal blocking filters, which are used to limit the infrared loading on the cryo-system. For a detailed description of the infrared filter stack in the cryostat, see table 1 in Masi et al. [15].

The next optical component is the stepped rotating HWP which modulates incoming polarization [16]. A single polarization is then selected by the polarizing grid. This scheme makes the instrument largely insensitive to spurious polarization introduced by optical elements after the grid. A non ideal behavior of the system components could still introduce spurious polarization even after the polarizing grid (for example a fraction of the signal reflected by the polarizing grid could be reflected back in the main optical path, or the detectors could display some sensitivity to polarization). Laboratory measurements, however, have shown that the upper limit in spurious polarization is of the order of 0.4% (see figure 13 in Torchinsky et al. [17]) and that this limit is uniform over the detector focal plane (section 9.2 in D'Alessandro et al. [16]).

The next optical device is an array of 400 back-to-back corrugated horns made of an assembly of two 400-horn arrays, composed of 175 aluminium platelets (0.3 mm thick) chemically etched to reproduce the corrugations required for the horns to achieve the required performance. An array of mechanical shutters (RF switches) separates the two back-to-back horn arrays in order to be able to close or open horns for self-calibration [18]. The shutters are spring loaded and activated by applying a voltage to an induction coil. As a result, the shutter requires continuous electrical current in order to remain closed. To prevent heating of the 1 K horn array by the thermal dissipation by the shutters, no more than two switches may be simultaneously activated. However, due to the modularity of the switch electronics, a maximum of two switch per module can be operated at the same time. Both front and back horns are identical with a field of view of 13 degrees FWHM with secondary lobes below -25 dB.

As described in O'Sullivan et al. [19], the back-horns directly illuminate the two-mirror off-axis Gregorian optical combiner which focuses the signal onto the two perpendicular focal planes. A dichroic filter splits the incoming waves into two wide bands centred at 150 GHz for the on-axis focal plane and 220 GHz for the off-axis focal plane. The focal planes are each

equipped with 992 NbSi Transition-Edge-Sensors cooled down to 320 mK using a sorption fridge; for a detailed description see Piat et al. [20].

An isometric rendering of the cryostat is provided in the right-hand panel of figure 1. The cryostat weighs roughly 800 kg and is around 1.6 m high with a diameter of 1.4 m.

The QUBIC-TD uses the same cryostat, cooling system, filters and general sub-system architecture as described above but with only 64 back-to-back horns and smaller mirrors to match the illumination of the 8×8 horn-array. It has a single 248 TES bolometer array operating at 150 GHz.

3 The horn-switch system

In the left panel of figure 1 we show a schematic of the QUBIC optical layout. The sky signal enters the cryostat propagating through a high-density polyethylene (HDPE) window. Then, a rotating half-wave plate modulates the polarization and a polarizing grid selects one of the two linear polarization components. An array of 400 back-to-back corrugated horns collects the radiation and re-images it onto a dual-mirror optical combiner that focuses the signal onto two orthogonal focal planes populated by TES detectors that directly absorb the incoming radiation. A dichroic low-pass filter placed between the optical combiner and the focal planes selects the two frequency bands, centered at 150 GHz and 220 GHz. The output of each detector contains interference fringes that are the so-called "visibilities" of the selected Fourier modes. The detectors are unpolarized, so that the polarization of the detectors system in Piat et al. [20].

The right panel of figure 1 shows a sectional cut of the instrument solid model, while the bottom panel is a detailed view of a single element of the horn-switch array, consisting of two corrugated back-to-back horns connected by a circular, smooth waveguide that can be shut by the switch blade (gray rectangle in the figure). The particular mechanical layout of the switch system (see figure 23 in section 5.1) did not allow us to place the blade at the center of the connecting waveguide, but slightly closer to the input aperture.

4 Back-to-back feed horns

4.1 Feed horns requirements and design

4.1.1 Requirements

The back-to-back horn array has two objectives: (1) the front (sky) horns define the field of view (FoV) of the instrument and (2) the back horns specify the edge taper illumination on the reflective optics and define the coherent inputs to the spatial beam combiner. A drawback of this implementation is that half of the power is lost in reflection to the sky by the polarizing grid. In table 2 we list the primary requirements of the back-to-back horn array with notes detailing their relevance. In our requirements we do not include the horn cross-polarization since the array is placed behind the polarization modulation/separation stage, so that the intrinsic horn cross-polarization is not a critical requirement.

4.1.2 Mechanical and electromagnetic design

We built identical front and back horns with a design based on an initial study by Maffei et al. [21] that explored the performance of waveguide feed horns with a corrugated flare profile

Requirement	Value	Notes
Center-to-center dis- tance	$14\mathrm{mm}$	The smallest achievable, for the optimal sam- pling of the angular power spectrum.
Aperture	$12\mathrm{mm}$	Driven by the FoV and by the illumination of the optical combiner.
Return loss	$< -25\mathrm{dB}$	Over the 130–240 GHz bandwidth. This is the target RL value for the fundamental mode only, used to optimize the horn design.
Insertion loss	$< 0.1\mathrm{dB}$	To ensure a negligible contribution to optical losses, the overall transmission being higher than 60%.
Mass	$30{ m g/horn}$	Must be suspended on the top of the optical combiner and operated in a cryogenic environment.

Table 2. Main requirements for the QUBIC back-to-back horns array

inspired by a Winston concentrator geometry. This profile was modified and further developed for the project Clover [22] and also adopted for other instruments such as QUIJOTE [23], and BINGO [24].

The profile optimized for QUBIC (see figure 2) has 0.5 mm deep corrugations to reduce the cross-polarization and return loss at 150 GHz. The horn waveguide has a corrugation depth of 0.7 mm and a minimum radius of 0.684 mm so that we select the HE₁₁ for the 150 GHz band with a cut-off frequency of about 130 GHz. The guide design also allows the propagation of four modes up to 240 GHz band, with the drawback of a larger cross-polarization at this frequency.

We also added a transition between the Winston and the waveguide sections of the horn which have a different corrugation depth, and a re-expansion that matches the corrugated waveguide to the smooth waveguide of the switch block.



Figure 2. The corrugations profile of the QUBIC horns.

As already noticed, although QUBIC is largely insensitive to any cross-polarization injected after the polarizing grid, we chose to implement a corrugated horn design that is generally used in single-mode optical systems with stringent requirements in cross-polarization and side-lobes. To understand our choice it is necessary to place it in the framework of the historical development of QUBIC and of the main changes in its design occurred over the years.

The choice of corrugated horns for QUBIC dates back to early phases of the project, described, e.g., in QUBIC collaboration et al. [1], when QUBIC was designed as an ensemble of single-band, dual-polarization instruments, with one dedicated cryostat per frequency band. At that time the polarizing grid was placed at 45° after the optical combiner to split the polarization states onto the two focal planes. In this design the feed-horn cross-polarization was an important issue and, therefore, we considered corrugated horns the best choice.

After the BICEP2 claim in 2014 [25], the importance to have a frequency channel dedicated to dust monitoring became clear and, consequently, the collaboration decided to modify the design to allow the detection of both 150 GHz and 220 GHz bands in the same module. This led to the current design, with the polarizer grid immediately below the rotating halfwave plate and a dichroic filter that splits the two frequency bands.

Meanwhile, a few prototypes of the corrugated horns had already been produced for a production chain for this design to be in place. For this reason we decided to modify just the most critical part of the horn, the mode converter, to allow the propagation of both the 150 GHz and the 220 GHz band, leading to the current geometry in which the horns are single-moded at 150 GHz and multi-moded at 220 GHz.

4.1.3 Electromagnetic simulations technique

We have simulated the fields propagating from the waveguide to the aperture of the QUBIC corrugated horn using the electromagnetic mode-matching technique [26, 27], depicted schematically in figure 3. This technique regards the corrugated structure as a sequence of smooth walled cylindrical waveguide sections, each of which can support a set of propagating transverse electric (TE) and transverse magnetic (TM) modes. At each corrugation the sudden change in the radius results in a scattering of power into backward propagating reflected modes in the left-hand side guide segment and forward propagating transmitted modes in the right-hand segment.

The power coupling between modes is given by the overlap integral $\int e_{n,l} h_{m,r} dA$, where $e_{n,l}$ is the transverse electric field of mode n on the left-hand side of the junction, $h_{m,r}$ is the magnetic field of mode m on the right-hand side of the junction and dA is a surface element on the transverse plane. The modes are then propagated through the length of waveguide section to the next scattering junction where the overlap integral between the modal components is computed again.

If \vec{A} and \vec{C} are column vectors of the mode coefficients of the fields incident from the left and the right, and \vec{B} and \vec{D} are the mode coefficients of the resulting reflected fields, then their relationship is described using a scattering matrix, **S**:

$$\begin{bmatrix} \vec{B} \\ \vec{D} \end{bmatrix} = \mathbf{S} \cdot \begin{bmatrix} \vec{A} \\ \vec{C} \end{bmatrix} = \begin{bmatrix} \mathbf{S}_{1,1} & \mathbf{S}_{1,2} \\ \mathbf{S}_{2,1} & \mathbf{S}_{2,2} \end{bmatrix} \cdot \begin{bmatrix} \vec{A} \\ \vec{C} \end{bmatrix}$$
(4.1)

whose elements are calculated using overlap integrals as described in Olver et al. [28]. The columns of the scattering matrix describe the amplitude of each output mode generated by a



Figure 3. Schematic of the mode-matching model implemented in the electromagnetic simulations.

unit-amplitude input mode. The scattering matrix for the horn as a whole, is computed by cascading the matrices for each uniform section and junction. We assume no scattering at the horn aperture, since it is electrically large (6λ) , so $\vec{C} = 0$.

The field at the aperture of the corrugated horn is then determined from $\vec{D} = \mathbf{S}_{2,1} \cdot \vec{A}$, where $\mathbf{S}_{2,1}$ is the sub-matrix that deals with the forward-propagating modes, and the reflected field is determined from $\vec{B} = \mathbf{S}_{1,1} \cdot \vec{A}$. The transmitted and reflected power are found by multiplying the complex elements of the relevant column vector by their complex conjugate and summing them.

In our analysis we used 60 waveguide modes (30 TE and 30 TM), which was more than sufficient for the accuracy levels shown in this paper. The TE and TM modes with power have a coherent phase relationship and in this case correspond to the single hybrid $\text{HE}_{1,1}$ mode.

In the 220 GHz band more than one column of the scattering matrices is non-zero and these represent possible independent modes of power transmission. We excite all modes equally at the input, $\vec{A} = [1, 1, 1, ...]^T$, and add the individual output fields incoherently (a detailed description can be found in Murphy et al. [29, 30]). The reflected power is calculated as a percentage of the power that could be transmitted by the number of propagating modes. Spikes can be observed at cut-off frequencies of higher order modes as these states transition from evanescent to propagating modes. These predicted features are also likely to depend on the exact details of the manufactured horns.

4.1.4 Simulations results

In figure 4 we show the return loss (left panel) and maximum value of the cross-polar beam (right panel) in the two QUBIC bands. We can see at a glance that the performance in the 150 GHz band is superior compared to the 220 GHz band. In fact the design was initially

tailored in the D-band and subsequently modified to accept also the higher band that could not be optimized in terms of performance like the lower frequency range.

The return loss at 150 GHz is, on average, around -25 dB, while in the higher frequency band it is compatible with -20 dB up to 230 GHz, and degrades to $\sim -10 \text{ dB}$ on the right hand side of the frequency interval. We assessed the potential impact of the poor return loss in the highest part of the 220 GHz band: a degradation of the return loss will induce a reduction of the horn transmission and therefore an overall decrease of the sensitivity. With a pessimistic -10 dB return loss over the whole 220 GHz, we estimate a degradation in the sensitivity of about 5%. We accepted this violation of the specified requirement because of the challenges in simultaneously optimizing the corrugated horns for both 150 GHz and 220 GHz bands without a significant increase in fabrication complexity.

The cross-polarization is very good ($\sim -35 \text{ dB}$) at 150 GHz, while it is around -5 dB at 220 GHz. This is in agreement with the design: at 150 GHz we have a single-mode corrugated horn, for which we expect excellent polarization purity, while at 220 GHz we have propagation of higher modes that do not preserve the polarization state. But this is not a problem for QUBIC, as already mentioned, because the polarization is selected before the radiation enters the horns. For this reason we show here the expected cross-polarization performance but we will not discuss it further in the rest of the paper.



Figure 4. Simulated return loss (left) and maximum cross-polarization (right) in the two QUBIC frequency bands.

In figure 5 we show the simulated normalized beam patterns at 150 and 220 GHz for the three main co-polar planes (E-plane, H-plane and 45° plane). The maximum directivity in the two frequency bands is 23.8 dBi at 150 GHz and 22.4 dBi at 220 GHz. At 150 GHz we can appreciate the typical Gaussian profile of single-mode corrugated horns, while at 220 GHz the main beam shape is a flat-top resulting from multi-mode propagation. The side-lobes are low, less than $\sim -30 \,\mathrm{dB}$ at angles larger than $\sim 30^{\circ}$.



Figure 5. Co-polar simulated beam patterns (E, H, and 45° planes) at 150 GHz (top row) and 220 GHz (bottom row).

4.2 The technological demonstrator feed horn array

We have chosen to develop the horn arrays in platelets because this technique represents the best option to realize large arrays of corrugated horns at our frequencies with the required performance and a sustainable cost, when compared with other options like electroforming, electrical discharge machining or even direct machining.

This technique, introduced by Haas in 1993 [6, 7], requires drilling circular holes into metal plates that are subsequently stacked and clamped or bonded. For QUBIC we have used two methods to drill the holes: chemical etching of 0.3 mm aluminum plates and computer controlled milling of top and bottom plates (3 mm and 6 mm thick, respectively). Then we mechanically clamped the plates with screws (applying a torque of $0.8 \text{ N} \cdot \text{m}$) and alignment pins. We have shown in other works [10, 14] that mechanical clamping is a valid alternative to bonding, does not impact negatively the performance and can be used also when the antennas operate in cryogenic conditions. In particular in [10] Del Torto et al. showed that a clamped platelet array of W-band feedhorns maintained its return loss performance after cooling to 4 K.

Each of the two 64-horns arrays is composed of 175 plates enclosed between the external flanges (panels (a), (b), and (c) of figure 6). We silver-plated each plate and both flanges to improve electrical conductivity along the horns profile. The thickness of the silver plating (quoted by the manufacturer) is $3.5 \,\mu\text{m} \pm 0.3 \,\mu\text{m}$ and the r.m.s. surface roughness measured in our workshop on one of the plates is $R_{\rm q} = 0.3 \,\mu\text{m}$. Indeed, on previous prototypes we verified the effect of silver plating by measuring the room-temperature insertion loss of platelet horns

with and without silver plating. We observed that the insertion loss decreased from about 0.45 dB to less than 0.1 dB when the aluminum plates were silver coated.

Each metal plate has an overall size of $112 \text{ mm} \times 112 \text{ mm}$, with a thickness of 0.3 mm. The plates also include four 3 mm diameter holes for alignment pins and 77 holes for the ERGAL (7075 aluminum alloy) M3 screws that pack the array between the two flanges.

The top flange has an overall size of $120 \text{ mm} \times 120 \text{ mm}$ with a thickness of 3 mm and it contains the 64 horn apertures with a diameter of 12.33 mm. It also contains countersink holes for the tightening screws so that the screw heads lie flush with the array planar surface. The bottom flange has an overall size of $133 \text{ mm} \times 133 \text{ mm}$ with a thickness of 6 mm and it contains circular waveguide segments with a diameter of 1.91 mm for each of the 64 horns.

The bottom flange couples to the switch array and is machined in a shape similar to that of the UG-387 waveguide flange (MIL-DTL-3922/67C) to obtain anti-cocking¹ interfaces between each waveguide pair. It also includes holes for the module tightening screws. Half of the screws run from the top to the bottom flange and hold each of the two arrays together (therefore allowing us to handle and test the arrays separately), the remaining group extends further to hold the horn and switch modules together (see panel (d) of figure 6).

4.3 Mechanical measurements and achieved tolerance

4.3.1 Experimental procedures

We tested the mechanical tolerance according to two different procedures. First we visually inspected the inner profile of a 4-elements horn array manufactured in brass that was cut to allow us to magnify the shape of the antenna teeth and grooves. Then we used a metrological machine (Werth ScopeCheck 200) to measure the position, diameter and deviation from circularity of each hole in the platelets of the final array.

Visual inspection. The left panel in figure 7 shows a section of the brass prototype. The enlargement in the right panel highlights the presence of cusps ($\leq 0.06 \text{ mm high}$) on the profile of all the corrugations that are the effect of a non uniform erosion of the metal during the etching process.

This non uniformity is a limitation which is inherent in the chemical etching process, so that we can expect that the antennas produced with this method present imperfections in their corrugated profile. In section 4.3.2 we discuss the impact of these defects on the feed-horn performance for QUBIC.

Metrology. We carried out metrological measurements using the Werth ScopeCheck 200 instrument, that performs precision measurements using either an optical or a tactile device. In our setup we used the optical sensor, which can be moved in three dimensions over a glass work plane where we laid our platelets. In section 4.3.3 we discuss the results of the metrological measurements.

4.3.2 Impact of mechanical imperfections on electromagnetic performance

We assessed the impact of the imperfections in the feed-horn profile caused by the etching process by computing the return loss and the co-polar radiation patterns on the E and H

¹The "cocking" effect refers to the presence of a non-parallel flange mate or gap at the interface, which can lead to signal loss and mechanical damage. The geometry of the "anti-cocking" interface is designed to minimize this possibility during installation of the waveguide flange fasteners [31].



(a)



(b)



Figure 6. The QUBIC TD feed-horns array: (a) chemically etched aluminum platelets during the stacking process; (b) 3-D model of the antenna array; (c) one of the two antenna modules after integration; (d) the complete integrated feed-horn-switch system.

planes considering two cases: (i) the nominal profile, and (ii) a profile modified inserting a step-like defect on teeth and grooves of all the corrugations (see figure 8).

We notice that the step-like defect in figure 8 is not fully representative of the cusp-like structure shown in figure 7. On the other hand we had to use the step-like approximation because the mode-matching code we used to simulate the multi-moded beams does not allow the simulation of cusp-like structures. We validated the use of this simplification by comparing single-mode beam simulations carried out with two codes: our mode-matching code for the step-like model and the SRSR software [32] for the cusp-like structure.



Figure 7. Left. A section of the first feed-horn prototype. Right. Detail of the cusps on teeth and grooves resulting from non uniform chemical erosion.



Figure 8. Sketch (drawn to scale) of the model used to simulate the imperfections in the feed-horn profile.

Return loss. Figure 9 shows the effects of the defects on the return loss, and we see that the narrower effective size of the etched holes shifts the resonances in frequency. We see that they do not change significantly the overall level, but shifts some of the resonances in frequency. In general, however, we can consider the impact on the return loss negligible.

Radiation pattern at the center frequency. Figure 10 shows the simulated radiation patterns (E-plane, H-plane, and 45° plane) at 150 GHz (top rows) and 220 GHz (bottom rows) for the two cases studied. The bottom plot in each figure shows the difference in dB of the beam patterns for the two cases. In the main beam region $(-15^{\circ} < \theta < 15^{\circ})$ the difference is less than 0.05 dB, and over all the $-90^{\circ} \le \theta \le 90^{\circ}$ range the difference is within $\pm 2 \text{ dB}$.



Figure 9. Impact of defects on the feed-horn return loss.

4.3.3 Results of metrological measurements

We have carried out metrological measurements of both feed-horn arrays and compared the manufacturing precision with the maximum achievable tolerance of the chemical etching process, which is ± 0.05 mm. In this section we will refer to the two arrays as array-1 and array-2.

We measured the holes of each antenna and alignment pin for all the aluminum plates, compared the measured positions and diameters with their nominal values and calculated the form tolerance (FT) of each hole (see the sketch in figure 11 for a definition of this parameter). This rich set of measurements allowed us to obtain the actual mechanical profiles of all the feeds in the array that we used to simulate their actual electromagnetic behavior. We then compared this family of simulations with the electromagnetic parameters measured in the laboratory, as explained at the beginning of section 4.4.

The box-plots of figure 12 show the deviation of the measured Cartesian center coordinates of the antenna holes from their nominal value (Δx , top-left, and Δy , top-right), the deviation between the measured and nominal hole diameters (bottom-left), and the corresponding distribution of FT values (bottom-right). The red line corresponds to a null deviation, while the green area highlights the expected manufacturing tolerance. The measurements are relative to the array-1.

As one can see, all the antenna positions comply with the manufacturing tolerance, while more than 90% of the antenna diameters are out of specification, generally larger than expected and distributed around two peaks: $\Delta d_1 = 0.07 \,\mathrm{mm}$ and $\Delta d_2 = 0.15 \,\mathrm{mm}$, with a maximum deviation of 0.25 mm. The measured FT show no significant deviation from circularity.

We obtained similar results for the alignment pins of array-1 and for both antenna holes and pins of array-2, but they are not reported here for simplicity. We measured also the



Figure 10. Impact of defects on the co-polar radiation patterns. The black line in the bottom panel shows the difference between the two beam patterns. The top plots refer to 150 GHz simulations, the bottom plots refer to 220 GHz simulations.

top and bottom plates of the arrays and they are in compliance with the milling precision tolerance of $0.03 \,\mathrm{mm}$.

The out-of-spec was due to a loose control of the chemical etching time. Indeed, one can see that the measurements are grouped in blocks of plates and the average deviations from the nominal diameters follow a bi-modal distribution. Discussing with the company that performed the etching we understood that the plates were treated in batches and that the etching time was not carefully controlled among the various batches for reasons independent of QUBIC. These problems were solved in the production of the horns for the FI (see section 4.5) by strictly controlling the etching time.

In section 4.4 we discuss the effect of this out-of-spec on the TD horns electromagnetic performance, where we compare the measured return loss and beam patterns with simulations run with the nominal and measured antenna profiles.



Figure 11. Definition of the form tolerance (FT) parameter.

The box-plots in figure 12 also highlight an oscillatory, almost sinusoidal pattern in the measurements of the holes centers coordinates as a function of the antenna number. This is likely a systematic effect in our measurement. In fact, this behavior correlates with the row-by-row scanning of the antenna holes in the square antenna array. We found a similar oscillatory pattern also in the diameter measurements, with an amplitude much smaller than the spread in the deviation of the diameters from their nominal value. Because we could not clearly identify this effect, neither in the measurement strategy nor in the measurement machine, this remains a reasonable hypothesis that is not demonstrated yet. For this reason we preferred not to decorrelate this effect from the data and to leave it as an additional source of uncertainty.

4.4 RF characterization

In this section we present the measured electromagnetic performance of the QUBIC TD feedhorns compared with simulations obtained using the measured profiles of all the feed-horns in the array. In our measurements we tested a subset of feed-horns ($\sim 40\%$ for array 1 and $\sim 14\%$ for array 2), each identified with a pair of numbers corresponding to the row and column position in the array.

In figure 13 we show four grids summarizing the measurements that we have conducted on the TD arrays. Each grid represents the 8×8 horn array in each module. On the top of each grid we have specified the number of the module and the type of measurements performed. In each grid we identify with black cells the horns tested at 150 GHz, with red cells the horns tested at 220 GHz and with black-red cells the horns tested at both frequencies. Notice that we do not report return loss measurements for the TD array at 220 GHz. In this case the measurements presented in this paper were carried out on the two central horns of the two modules of the QUBIC FI that share the identical electromagnetic design and manufacturing technique of the TD (see section 4.5).



Figure 12. Results of the metrological measurements of the antenna holes of array 1. We do not report the results for the second array because they tell essentially the same story. Top-left and top-right: deviation of measured x and y center coordinates from the nominal value. Here for each antenna (horizontal axis) we display the spread of the deviations, Δx and Δy , of the holes center coordinates. The box represents the 1- σ spread, the dashed line the peak-to-peak. The red line corresponds to a null deviation. Bottom-left: deviation of measured diameters from the nominal value. Here for each plate (horizontal axis) we show the spread of the deviations, Δd , of the holes diameters. Again, the box represents the 1- σ spread, the dashed line the peak-to-peak. The red line corresponds to a null deviation, while the green area highlights the expected manufacturing tolerance. Bottom-right: measured form tolerances for all the holes in the antenna array platelets. See figure 11 for the definition of form tolerance.

4.4.1 Experimental setup and procedures

The experimental setup consisted of a vector network analyzer (VNA) equipped with millimeter extensions for full two-port characterization in the 110–170 GHz and 170–260 GHz bands. With this setup we measured the horn return loss and radiation patterns and compared our measurements with simulations. We connected the array to the VNA by means of a cascade of adapters (figure 14) to connect the standard rectangular waveguide of the VNA to the custom circular one of the horns.

The reference plane for the VNA measurements is at the WR06 millimeter-wave extension waveguide port and calibrated with thru-reflect-line (TRL) standards. The time domain window gating function of the network analyzer system is used to isolate the horn backscattered signal from the undesired reflections associated with the test adapters. Because the VNA setup is sensitive to only the dominant propagation mode we could not test the multi-mode behavior of the system in the 220 GHz band. Consequently also the results of the simulations displayed in figures 15, 16 and 17 regard single-mode propagation.

To sample the beam patterns, the TD feed-horn array was mounted on a goniometer



Return loss measurements

Figure 13. Grids summarizing the electromagnetic measurements carried out on the TD feed-horn arrays. See text for further details.

fixed on an optical bench. The horn array was illuminated in the 110–170 GHz band by a corrugated circular standard gain horns manufactured by Radiometer Physics, while in the 170–250 GHz band it was illuminated by a smooth wall standard gain horn manufactured by Custom Microwave. In these measurements we moved the device under test (DUT) in azimuth with an angular step of 1° and selected the proper reference plane (*E*-plane or *H*-plane) by



Figure 14. *Left*: detail of the adapter cascade used in the 110–170 GHz measurements. After the VNA golden millimeter extension where the TRL calibration plane is defined, there is a golden rectangular to circular waveguide taper. The golden taper is connected to the horn array by means of silver flange adapter. *Right*: detail of the adapter cascade used in the 170–260 GHz measurements. A rectangular taper from WR4.3 to WR06 is attached to the VNA golden waveguide millimeter extension (TRL calibration plane). After that, the same components of the lower band set-up is used.

properly rotating the launcher and the DUT.

4.4.2 Results

In this section we summarize the measured return loss and beam patters with simulations and show that we obtain an overall match within the uncertainties given by the mechanical differences among the horns.

Return loss. In figure 15 we show the results of the return loss measurement in both bands compared with the simulations. The orange area is the envelope of the return loss simulated for all the 128 feed-horns in the array, each with its own measured profile, while the blue area is the envelope of the measured return loss for all the tested horns (refer to figure 15).

We see that the measured reflection matches the simulation, within the scatter given by the mechanical differences among the horns. We also see that the average achieved return loss at 150 GHz lies around $-20 \,\mathrm{dB}$, while in the higher band it is around $-25 \,\mathrm{dB}$ up to 230 GHz and then degrades to about $-10 \,\mathrm{dB}$ as expected. We believe that the higher-than-expected return loss at 150 GHz and the scatter among simulations is caused by the out-of-spec in the mechanical tolerance discussed in section 4.3. Given the improvements adopted in the manufacturing procedure we think that this scatter and the overall return loss are reduced in the FI horn array.

The reader may notice that the measured return loss between 190 and 230 GHz is about -25 dB, therefore 5 dB lower than the value resulting from the simulation of the nominal feedhorn (see the left panel of figure 4). This is because the measurements and the simulations displayed in figure 15 are relative to single-mode propagation, while the simulation in figure 4 considers all the possible modes that can propagate in the 220 GHz band.

Beam patterns. We show our beam pattern measurements compared with simulations in figures 16 and 17. Also in these figures the orange area is the envelope of the simulated



Figure 15. Measured return loss compared with simulation and requirement listed in table 2.

patterns for all the feed-horns in the arrays and the blue area is the envelope of the measured patterns.

In the 150 GHz band (figure 16) we measured E- and H-plane diagrams in the entire D band (110-170 GHz) but for sake of simplicity reported them for only three frequencies: 145, 150 and 155 GHz. Measurements match simulations very well (with a few dB discrepancy) down to about -30 dB. The scatter increases at larger angles, where the detected power is smaller and the measurement becomes sensitive to signal reflections.

We have obtained similar results in the 220 GHz band (figure 17). In this case we reported only the *H*-plane diagram at five frequencies, equally spaced between 190 and 230 GHz. Also in this band there is a very good match between measurements and simulations down to -30 dB.

We have also measured the cross-polar beam pattern in the 150 GHz band for two horns of the array 1. We show these patterns in figure 18, where the measured data are compared with simulations. We see that measurements and simulations match at the level of the pattern general level and trend, although our setup was not optimized against reflections to allow a fine-grained match. This is particularly apparent at the highest frequency, where we see a mismatch of about 10 dB in the maximum cross-polarization.

4.5 The full instrument feed horn array

Here we present briefly the feed-horn array developed for the QUBIC FI, shown in figure 19. The FI horn system has the same electromagnetic design of the TD and it was manufactured with the same technique: the inner part by chemically etching 0.3 mm thick aluminum sheets and the front and back flanges by mechanical milling 3 mm and 6 mm thick aluminum plates. Like for the TD array we silver-coated all plates before the final integration.



Figure 16. Measured co-polar beam patterns at 150 GHz compared with simulations. *Top: E*-plane. *Bottom: H*-plane.

In this array we measured the geometrical profile of a large subset of horns in the two modules. This allowed us to simulate the expected performance of this subset and to compare it with laboratory measurements.

The detailed discussion of the development and testing of this part is out of the scope of this paper so that we defer the full discussion of the QUBIC FI horns to a forthcoming dedicated paper.

5 Switch system

5.1 Switch requirements and design

The QUBIC switch array is used to select the baselines under test during the calibration phase. Theoretically the self-calibration procedure requires to acquire data from a known calibration source for all possible baselines, each obtained by closing all the horns apart from the pair of the selected baseline. This approach, however, would cause a huge variation of the radiative load unto the focal plane, shifting the detectors' temperature well below the design one. In QUBIC we apply a different, but mathematically equivalent approach [5] in which we close only the horns of the baseline under test, thus maximizing the signal stability during calibration.



Figure 17. Measured co-polar (H-plane) beam patterns in the 220 GHz frequency band.

In table 3 we list the main requirements of our switch system. In particular the insertion loss must be as low as possible, even if the 1 K working temperature helps at least in reducing its thermal noise, while we require high isolation between on and off states and fast commutation for accurate and efficient baseline calibration.

From an electromagnetic point of view, the switch design was based on a best effort approach where the length of the waveguide is set by the shutter mechanism dimensions plus the thickness of the plates housing the threaded holes to mate the two horns arrays. The waveguide diameter matches the input waveguide of the horn. The switch waveguide was designed with a choke trap to short-circuit the gap when the switch is open. Such a structure was originally designed for the lower band, but extended simulations showed that it was effective also for the upper band. In figure 20 we report the results of the simulation of the transmission and reflection of the main electromagnetic mode (TE11) for 25 mm waveguide length and 200 μ m choke gap. In the two QUBIC bands the reflection is always below -20 dB.

We have simulated the insertion loss at room temperature with an effective bulk re-



Figure 18. Measured and simulated cross-polar (plane $\phi = 45^{\circ}$) beam patterns in the 150 GHz frequency band. The patterns are normalized to the 0 dB maximum of the co-polar patterns.



Figure 19. The back-to-back feed-horn system of the QUBIC FI. *Left:* Stacked platelets of a quarter array. *Right:* one of the two 400 horn arrays.

sistivity of $16 \mu\Omega$ cm, which takes into account the effect of surface roughness (O. Peverini, private communication). The simulated loss (reported in figure 20) is a factor two larger than the requirement, although it is consistent with the loss of a waveguide of identical length

Requirement	Value	Notes	
Insertion loss	$< -0.1\mathrm{dB}$	Driven by minimization of signal loss and ther- mal noise. This is not met by the TD which is a factor 2 worse.	
Isolation	$> 50 \mathrm{dB}$	Driven by maximizing contrast in fringe pat- terns during calibration.	
Return loss	$< -20\mathrm{dB}$	Over the 130–240 GHz bandwidth.	
Switch commuta- tion time	$\sim \mathrm{ms}$	To be negligible in the calibration duty cycle.	
Heat load	as low as possible	To minimize thermal drifts of the 1 K stage.	
Mass	as low as possible	Must be suspended on the top of the optical combiner.	

 Table 3. Main requirements for the QUBIC switch array at room temperature

without the choke (-0.18 dB). For comparison, an idealized loss-less waveguide (actually $\rho = 10^{-6} \mu \Omega$ cm for computational reasons) with the same gap and choke has an insertion loss of -0.02 dB. This is the reactive mis-match loss associated with the junction geometry. We have also simulated the insertion loss for an ideally smooth waveguide using the bulk resistivity of Al6061-T6 ($\rho = 4.19 \ \mu \Omega \text{ cm}$) as reported in Clark et al. [33] and in ASM HANDBOOK [34]. Since the return loss is dominated by the switch geometry, the output of the simulation for the three resistivities are, in this case, practically identical.

At higher frequencies the loss is definitely higher than desired. This can be partially due to the almost double length of the waveguide, if measured in wavelength units. Another reason why the loss is higher in the upper band is because the choke itself is less effective. While in the lower band the discontinuity of the waveguide is made invisible to the incoming radiation by a choke trap which lowers the reflected power at the gap below -40 dB (for the waveguide and choke geometry see figure 24), in the higher band the reflected power rises at -12 dB. Moreover, starting from 190 GHz TE11 begins to convert to higher order modes and the simulation is less revealing.

In order to verify that the higher modes are not reflected back by the waveguide gap, we run a simulation of the return loss including all the modes able to propagate within our band, i.e. from the TM01 up to the TM21 mode. The result, reported in figure 21, shows

that on average this return loss is less than 10%.



Figure 20. Top: Simulation of TE11 mode insertion loss of the switch waveguide for a choke gap of 200 μ m. The red solid line is for an effective resistivity $\rho = 16 \ \mu\Omega$ cm, the green dashed line is for a perfect conductor ($\rho = 10^{-6} \ \mu\Omega$ cm for computational reasons) and the orange dashed line is for Al-6061 resistivity at room temperature ($\rho = 4.19 \ \mu\Omega$ cm). Bottom: Simulation of TE11 mode return loss of the switch waveguide for a choke gap of 200 μ m. Since the return loss is dominated by the switch geometry, the cases for the three resistivities ($\rho = [0, 4.19, 16] \ \mu\Omega$ cm) are practically identical. The dashed vertical lines are placed where the higher order modes start to propagate. We used the same color code as in figure 21.

Since QUBIC is designed to observe the low/medium ℓ multipoles, the horns' spacing is one of the smallest possible. As a matter of fact, a triangular lattice is even more compact, but it would have left no room for the platelets tightening screws and for the shutters' mechanism. The QUBIC's square lattice is a compromise that leaves enough room available to realize the mechanical shutters (switches). We designed each switch around the smallest commercially available coil manufactured by Line Electric (model TO-5S figure 23 *left*). The area occupied by the whole device is so small that the minimum center-to-center distance between horns is not determined by the switch, but by the horn mouth diameter. The total thickness of the switch block is 24.75 mm and this is also the waveguide length. The TD switch array in its final form can be seen in figure 22, with and without the lid assembled. Top panel shows the array of newbwaveguides and threaded holes to fix the top horn array. The two DB-37 connectors are used to bias the coils. Bottom panel shows the internal part of the switch array. All the coils, shutters and waveguides are clearly visible.

The shutter is realized by means of a thin stainless steel blade entering the waveguide through a thin gap and acting as a short circuit. The gap is 200 μm wide and the blade is 100 μm . The back and forth blade movement is actuated by the coil pulling a ferrite which is



Figure 21. Simulation of the return loss induced by the 200μ m gap where the shutter enter to short circuit the waveguide.



Figure 22. *Left*: the switch array closed and ready to be integrated with the horn arrays. *Right*: the inner part of the switch assembly with solenoids, shutters and waveguides clearly visible.

soldered to a hook attached to the shutter. The shutter makes an angular movement around a pivot. The shutter returns to its original position thanks to a spring pushing back the ferrite. The center of figure 23 shows a detail of the TD solid model, with its main components highlighted. On the right of the same figure, a microscope picture of the device is reported.

The coil nominal ohmic impedance is 80Ω at room temperature, but lowers to 19Ω at 4 K. We use a constant current of 90 mA to hold the ferrite inside the coil, resulting in a Joule dissipation of 150 mW per shutter. As already stated, we decided to limit to two the number of switches closed at the same time. However, from a thermal point of view the system can



Figure 23. Left: the TO5-S coil used to move the shutter inside and outside the waveguide. Right: detail of the 3D model where single components of a unit cell are identified.

manage four switches closed at the same time. As a matter of fact, if we limit to four the number of simultaneously biased switches we reduce the extra heat load to 600 mW. Thanks to the large mass of the horns plus switches system, the heat released increases horn-switch block temperature on a timescale much longer than the RF load change without any impact on the self-calibration effectiveness. As a matter of fact, during the test of the assembled TD, we found that the heat released increased horn-switch block temperature by less than 2 mK in a timescale of 2 minutes. This effect is not detected by the focal-plane detectors because of a Neutral Density (ND) filter located on the 1-K cold stop, as described in fig.1 caption. This ND filter was added for the testing campaign in the lab in order to reduce the background power on the detectors. When observing the sky the ND filter will be removed.

Another reason not to exceed the number of four active switches (two per electronic module, which are two in the TD) is because the minimum number of wires necessary to drive N devices is (N + 1), where the extra wire is used for current return of the N coils. Since the driving current per coil is not negligible, the return wire must be made of copper instead of phosphor-bronze or manganine (this is discussed in detail in section 5.3.1). To minimize the conductive heat load, the copper wire must be thin ($\leq 200 \,\mu$ m), reducing the number of coils active at the same time to two per return wire. Since there are two switch banks, operated by two electronic boards, there are a maximum of four switch shut at the same time. However, since the self calibration procedure is effective even with only two horns closed, we limit to two the active switches operated at the same time.

5.2 Single channel prototype

We developed a single channel prototype at the APC Laboratory in Paris [35] to test the electromagnetic performance, the effectiveness of the linear motion design and the compatibility with the cryogenic environment. We realized both the single channel prototype and the full array for the QUBIC TD in Al-6061T6. The 3-D sketch in figure 24 shows the single channel switch without the lid to highlight the coil pulling the shutter inside a 200 μ m gap, in between the two sections of the circular waveguide. Details of the choke trap are also given. When the shutter is outside the waveguide, the switch is on and the coil is not biased. The shutter movements are limited by a couple of stainless steel alignment pins.

The prototype was tested four times in liquid nitrogen (LN2) at 77 K to verify the coil lifetime. In each thermal cycle the switch worked at 50 Hz, accumulating hundred thousands



Figure 24. Solid model of the single channel prototype: the lid was removed to reveal the solenoid and the shutter in their open position. The detail magnifies the waveguide with the choke trap. A schematic cross section of the waveguide and choke with the drawing dimension is shown too.

operations. The functionality of the back and forth movement of the shutter was verified by means of an optical fiber. This prototype never showed any issue in LN2.

5.2.1 Switch electromagnetic measurements

We characterized the single channel prototype at room temperature in both the QUBIC bands: 135–165 GHz and 190–220 GHz using the VNA available at the laboratory of the University of Milano-Bicocca. We used a pair of rectangular-to-circular transitions to connect the VNA to the switch. Since the expected insertion loss of the switch was low, we measured and subsequently removed the contribution of the tapered transitions.

When the switch is open, the insertion loss at 150 GHz is IL = -0.2 ± 0.05 dB at room temperature; see figures 25 and 26 for which left panels report the insertion and return loss in the QUBIC lower band and the right panels report the same quantities for the upper band. The ohmic contribution to the loss is anticipated to be a monotonically increasing function of the guide surface roughness relative to the skin depth, which are $\sim 1 \,\mu\text{m}$ and $0.27 \,\mu\text{m}$ (computed for $\rho = 4.19 \,\mu\Omega$ cm), respectively, at 150 GHz [36, 37]. As already said in Section 5.1, the surface roughness associated with machining can be phenomenologically accounted for using an effective bulk resistivity of $16 \,\mu\Omega$ cm (O. Peverini, private communication) instead of the more conventional $\sim 4.19 \,\mu\Omega$ cm value for room temperature Al 6061-T6 [33, 34].

As reported in 5.1, there is a good agreement between simulations and measurements of the insertion loss at room temperature (IL=-0.2 dB at 150 GHz) while the simulation above 190 GHz can be inaccurate because of mode conversion from TE11 to TM11 (compare fig-

ure 20 with figure 26). After cooling, the ohmic contribution to the insertion loss decreases and the contribution associated with surface roughness mode conversion remains practically unchanged. Since the Al 6061-T6 resistivity decreases from $4.19 \,\mu\Omega$ cm at room temperature down to 1.38 $\mu\Omega$ cm at 10 K [33, 34], we can estimate the insertion loss at cryogenic temperature to be $-0.18 \,\mathrm{dB}$ at 150 GHz and $-0.73 \,\mathrm{dB}$ at 220 GHz. It is reasonable that also impedance mismatch and mode conversion are contributors to the insertion loss, but at the moment we don't have a way to estimate their contribution.

The measured return loss, shown in the left bottom plot of figure 25, confirms that reflections are well within the requirement of -20 dB as specified in table 3. At 220 GHz the insertion and return loss requirements are not met, but this was expected because the prototype was designed and manufactured taking into account only the QUBIC lower band. Moreover, the TD works at 150 GHz, so the 220 GHz performance is not critical. In the FI the waveguides will be manufactured adopting all the techniques able to reduce the surface roughness like a combination of electrical discharge milling, high-grade reamers and chemical polishing to approach the desired electromagnetic specifications. In case a cryogenic characterization will be crucial, we will adopt the setup described in Zannoni et al. [38].



Figure 25. Measured insertion loss (top) and return loss (bottom) of the switch prototype in the band 135–165 GHz. Left: Switch ON. Right: Switch OFF. $1-\sigma$ confidence bands are plotted in blue.

5.3 The technological demonstrator switch array

5.3.1 Switch manufacturing and control electronics

The TD switch array design is based on the single channel prototype. It reflects the 8×8 structure of the feed-horn array which was chosen as a trade-off between the required filling



Figure 26. Measured insertion loss (top) and return loss (bottom) of the switch prototype in the band 190–250 GHz. *Left:* Switch ON. *Right:* Switch OFF. $1-\sigma$ confidence bands are plotted in blue.

factor and the possibility to leave enough room for the screws used to pack the platelets and connect the horns to the switch array. It was designed at APC in Paris and manufactured at the machine-shop of the University of Milano-Bicocca using Al-6061T5 for the two main shells. The main body is made by two parts. The first is a base housing most of the waveguide, the printed circuit board (PCB) and the coils+shutters. The bottom part of the base has the threaded holes to interface with the bottom horn array. The second part is a lid with the rest of the waveguide length and the threaded holes to interface with the top horn array. The shutters are a replica (64 times) of the single channel shutter. They are mounted on a custom PCB sharing the same footprint of the waveguides. Due to the limited cooling power at 1 K, the number of wires reaching the 1 K zone is reduced to a minimum. The heat dissipated by the coils while biased to close the shutters is also non negligible. So, in the 8×8 TD it is possible to close only a small fraction of the shutters and not in an arbitrary configuration.

A modular electronics is used to operate the shutters and read their position. Every electronic module is in charge of fifty shutters, being able to operate two of them at a time. Every module is composed of two arbitrary current generators to bias the coils. A small sinusoidal modulation of the coil current is used to read the shutter position, using the different phase delay between current and voltage when the ferrite is outside (shutter open) or inside (shutter closed) the coil. A set of six Analog to Digital Converters capture the alternate current (AC), AC voltage and direct current (DC) voltage during the 8 ms after the command to close a shutter is set. The DC voltage is used to evaluate the ohmic load of the lines+coil system, in order to verify that there is electrical continuity. The AC signals are used to calculate the phase delay in a way similar to what is done in phase locked loop (PLL) circuits. The sinusoidal current and voltage signals are digitized as follows: when the signal is above its average value, it is recorded as 1, otherwise as 0. In this way two squared waves are generated, one for the voltage and one for the current, which phase delay is easily computed as π times the average value of a third square wave obtained by applying XOR operator to the first two. When a switch is set to *closed*, a current pulse of 350 mA lasting ~ 5ms is used to energize the switch coil and the ferrite is attracted inside it. Once the ferrite is inside, the current is reduced to 90 mA, which is enough to hold in place to shutter reducing the heat load.

An alternative method to compute the phase difference was tested too. It is based on the fits of the digitized signals (voltage and current) and the consequent computation of the phase delay. This approach, however, has a computational cost much higher than the XOR method because of the necessity to digitally filter the data and run fitting algorithms. A direct comparison of the two methods showed us that the two are perfectly compatible within the statistical uncertainties.

5.3.2 Switch cryogenic tests

We tested the TD switch array at 5 K in the Milano Bicocca Millimeter Lab using a custom cryofacility. The cryofacility is composed of an aluminum vacuum chamber operated by Cryomech PT407 pulse tube and equipped with a Varian Navigator turbo pump assisted by an Agilent Scroll roughing pump. A SRS CTC-100 cryogenic temperature controller is in charge of the temperature readings and stabilization by means of load resistors. Temperature sensors are calibrated DT-470 Lake Shore diodes. The base temperature with no heat load is $3.5 \,\mathrm{K}$.



Figure 27. Technological demonstrator switch and horn array assembled and placed in the cryofacility for cryogenic functional tests.

Once we assembled the TD switch array with the two horn arrays, we placed it in the cryofacility which reached 5 K instead of 4 K because of the heat load of the numerous needed harness wires. We considered 5 K to be representative of the behaviour of the device at its nominal working temperature (1 K). The aim of these tests was to find the effective current values to move the shutters and maintain them in closed position in a harness configuration similar to that in the QUBIC cryostat and with the shutter release spring stiffness due to cryogenic temperature, very similar to the operative one. Moreover, an important goal of the tests was to verify the switch functionality and reliability at cryogenic temperature.

Figure 28 shows the phase between voltage and current for every switch and for increasing values of exciting current. We applied the excitation current from 20 mA up to 350 mA. All



Figure 28. Technological demonstrator switch excitation current optimization. All the switches move (and the phase changes) for current as high as 300 mA. A conservative value of 350 mA was adopted to assure the closing movement.

the switches did not move till a current value of 250 mA, for which part of them closed, and the rest stayed open. At 300 mA all the switches moved to closed position. We set a default conservative value of 350 mA. We also found that once a switch is closed, a much lower holding current can be applied. Experimental verification led to a value of 90 mA as the holding current capable to keep every switch in closed position. A typical excitation pulse at cryogenic temperature with the superimposed oscillating current can be seen in figure 29. The phase value doesn't change from 350 mA to 90 mA, proving that the switch is kept in closed position even with the lower current.

Once we found the proper excitation current, we tested all the switches one by one, starting from room temperature. Three of them were defective (numbers 5, 26 and 60 because open or short-circuited) and two of them (n. 1 and n. 32) were mechanically stuck in the open



Figure 29. Technological demonstrator switch excitation and readout current example. The excitation continuous current, plus the monitoring sinusoidal one, is kept at 20 mA for the first 6 ms, then a 350 mA pulse lasting 4 ms is applied to make the ferrite enter the coil and move the shutter. After 10 ms the current is lowered to 90 mA to reduce heat load.

position. These last two switches were stuck probably because of a non perfectly plane shutter. However, at cold temperature n.1 resulted operative. The result of this test is reported in figure 30.



Figure 30. Phase between voltage and current measured by XOR algorithm. *Left*: shutters in open position *Right*: shutters in closed position. Switch n. 1 was operative at 5 K. N. 5, 26 and 60 were defective and n. 32 was stuck open.

Finally we performed a statistical test operating each working switch 100 times. We

computed the phase delay in open and closed position using both the functional fit and the XOR algorithm, applied to the data acquired by the AC coupled ADCs. Both methods gave consistent values, as reported in table 4.

FIT		XOR		
Open	Closed	Open	Closed	
1.52 ± 0.01	$1.32{\pm}0.01$	$1.52{\pm}0.01$	$1.33 {\pm} 0.01$	

Table 4. Phase delay statistics in radians for 60 switches and 100 cycles each

However, every switch shows a distinctive mean phase delay with a standard deviation which is one third of the one of the full population, reflecting on one side the intrinsic different impedance of each single coil, and on the other side the intrinsic accuracy of the phase measurement of our electronics which is of the order of 0.003 rad (10'). For this reason, we defined a standard calibration procedure to be run when the focal plane is cold, to compile a look-up table of the phase delay of every switch in open and closed position, based on a statistical analysis of 100 switch operations. The look-up table will then be used to decide if the switch moves in the right position or not and, if needed, to repeat the command.

Despite the overall good behaviour of the TD switch array, the mechanical tolerances needed to operate the shutters resulted in a certain level of unreliability. The main issue is that a few shutters didn't completely enter the wave-guide, even if their ferrites were totally inside the coil. The main reason of this is the hook soldered on the ferrite responsible for the back and forth movement which was not sufficiently controlled to enable the ideal displacement of the shutter. The design of the stainless steel blade was consequently redesigned to reduce this uncertainty and improve the reliability of the switch motion.

5.4 The full instrument switch array

The Full Instrument switch array was designed replicating the TD and adapting it to fill the circular aperture sampled by the 400 horns. Unfortunately, the TO5-shaped coils were out of stock and were no longer manufactured. Due to size constraints, selection of other switches was limited. We selected a bi-stable micro-shutter (the BOS7/10, manufactured by the Japanese company Takano Co. LTD) and modified the design to fit the volume available. The driving electronics that was developed for the previous shutters was readily adapted to the new devices. Because the new shutters are bi-stable they do not dissipate energy except for the few milliseconds when changing state. They can be operated, in sequence, in any number from 1 to 400. Since this model is designed for room temperature operations, we tested it at 5 K to confirm that it remains operative at cryogenic temperatures. We collected several thousands of movements without any problem, simply adjusting the excitation current with temperature. The Full Instrument switch array is currently being developed.

6 Conclusions

In this paper we have described the design, manufacturing and testing of the feed-horn-switch system of the QUBIC technological demonstrator (TD), which will demonstrate the concept of bolometric interferometry by observing the polarized microwave sky from the high-altitude Alto Chorillo site in Argentina. The TD is the precursor of the QUBIC full instrument (FI), that will measure the CMB polarization from the same site.



Figure 31. Open view of the Final Instrument switch array partially populated by the new bi-stable shutter.

The TD horn-switch system is composed by a square array of 64 back-to-back corrugated feed-horns interspersed by 64 mechanical switches that can open and close the connecting circular waveguides. We designed the horns to allow the operation within two wide bands centered at 150 and 220 GHz, and manufactured them in platelets that were drilled with a combination of photo-chemical etching (inner, 0.3 mm plates) and mechanical milling (outer, 3 mm and 6 mm flanges).

We fully characterized the mechanical profile of all horns and found that the hole diameters of the inner plates were, on average, larger than the expected tolerance. The reason was the etching time that was not properly controlled during the process. We eliminated this problem in the full instrument (FI) horns and found that this out-of-spec was not critical for the objectives of the TD.

The measured electromagnetic performance agrees with simulations. In particular we obtained a return loss around -20 dB up to 230 GHz and beam patterns in agreement with single-mode simulations down to -30 dB.

Regarding the switches, we performed electromagnetic characterizations on the single channel prototype, finding return and insertion losses at 150 GHz consistent with expectations $(< -25 \text{ dB} \text{ and } \sim -0.1 \text{ dB}, \text{ respectively})$ and an isolation larger than 70 dB (specification was 50 dB). At 220 GHz, return and insertion losses specifications are not met, but this is expected because the prototype was designed and manufactured for the QUBIC lower band, which is the only one operative in the TD. We successfully tested the switch array at room temperature and inside the QUBIC cryostat, at 5 K, close to the nominal final one (1 K). We developed a readout system able to monitor the actual switch positions in operative conditions (1 K). Since the micro-miniaturized coils of the TD are not commercially available anymore, we

found an alternative for the FI which forced us to redesign the whole mechanism around a room temperature commercial bi-stable shutter which has been already tested at 5 K. This new shutter remained operative at cryogenic temperature simply adjusting the bias current.

We have completed the development of the feed-horn arrays while the switch system is currently being developed. After their mutual integration the whole system will be ready for the upgrade from the TD to the FI.

Acknowledgements

QUBIC is funded by the following agencies. France: ANR (Agencie Nationale de la Recherche) 2012 and 2014, DIM-ACAV (Domaine d'Interet Majeur-Astronomie et Conditions d'Apparition de la Vie), Labex UnivEarthS (Universitè de Paris), CNRS/IN2P3 (Centre National de la Recherche Scientifique/Institut National de Physique Nuclèaire et de Physique des Particules), CNRS/INSU (Centre National de la Recherche Scientifique/Institut National de Sciences de l'Univers). Italy: CNR/PNRA (Consiglio Nazionale delle Ricerche/Programma Nazionale Ricerche in Antartide) until 2016, INFN (Istituto Nazionale di Fisica Nucleare) since 2017. Argentina: MINCyT (Ministerio de Ciencia, Tecnología e Innovación), CNEA (Comisión Nacional de Energía Atómica), CONICET (Consejo Nacional de Investigaciones Científicas y Técnicas). Ireland: D. Burke and J.D. Murphy acknowledge funding from the Irish Research Council under the Government of Ireland Postgraduate Scholarship Scheme. D. Gayer and S. Scully acknowledge funding from the National University of Ireland, Maynooth. D. Bennett acknowledges funding from Science Foundation Ireland.

References

- QUBIC Collaboration, E. Battistelli, A. Baú, D. Bennett, L. Bergé, J. P. Bernard et al., QUBIC: The QU bolometric interferometer for cosmology, Astroparticle Physics 34 (2011) 705 [1010.0645].
- [2] A. Mennella, P. Ade, G. Amico, D. Auguste, J. Aumont, S. Banfi et al., QUBIC: Exploring the Primordial Universe with the Q&U Bolometric Interferometer, Universe 5 (2019) 42.
- [3] J. C. Hamilton, L. Mousset et al., QUBIC I: Overview and ScienceProgram, arXiv e-prints (2020) arXiv:2011.02213 [2011.02213].
- [4] L. Mousset, M. M. Gamboa Lerena et al., QUBIC II: Spectro-Polarimetry with Bolometric Interferometry, arXiv e-prints (2020) arXiv:2010.15119 [2010.15119].
- [5] M.-A. Bigot-Sazy, R. Charlassier, J. Hamilton, J. Kaplan and G. Zahariade, Astrophysics Self-calibration : an efficient method to control systematic effects in bolometric interferometry, Astronomy & Astrophysics 59 (2013) 1.
- [6] R. W. Haas, D. Brest, H. Mueggenburg, L. Lang and D. Heimlich, Fabrication and performance of mmw and smmw platelet horn arrays, International Journal of Infrared and Millimeter Waves 14 (1993) 2289.
- [7] R. W. Haas, Further development of mmw and smmw platelet feed horn arrays, in Multi-Feed Systems for Radio Telescopes, D. T. Emerson and J. M. Payne, eds., vol. 75 of Astronomical Society of the Pacific Conference Series, p. 99, January, 1995.
- [8] J. J. Bock, J. Gundersen, A. T. Lee, P. L. Richards and E. Wollack, Optical coupling, in Journal of Physics Conference Series, vol. 155 of Journal of Physics Conference Series, p. 012005, March, 2009, DOI.

- [9] J. W. Britton, J. P. Nibarger, K. W. Yoon, J. A. Beall, D. Becker, H.-M. Cho et al., Corrugated silicon platelet feed horn array for cmb polarimetry at 150 ghz, in Millimeter, Submillimeter, and Far-Infrared Detectors and Instrumentation for Astronomy V, W. S. Holland and J. Zmuidzinas, eds., vol. 7741 of Society of Photo-Optical Instrumentation Engineers (SPIE) Conference Series, p. 77410T, July, 2010, DOI.
- [10] F. Del Torto, M. Bersanelli, F. Cavaliere, A. De Rosa, O. D'Arcangelo, C. Franceschet et al., W-band prototype of platelet feed-horn array for cmb polarisation measurements, Journal of Instrumentation 6 (2011) 6009 [1107.1157].
- [11] F. Del Torto, C. Franceschet, F. Villa, P. Battaglia, M. Bersanelli, F. Cavaliere et al., A 7X7 elements Q-band feed-horn array for the LSPE-STRIP instrument, in 36th ESA Antenna Workshop on Antennas and RF Systems for Space Science, 2015.
- [12] M. D. Niemack, J. Beall, D. Becker, H. M. Cho, A. Fox, G. Hilton et al., Optimizing feedhorn-coupled tes polarimeters for balloon and space-based cmb observations, Journal of Low Temperature Physics 167 (2012) 917.
- [13] C. Bischoff, A. Brizius, I. Buder, Y. Chinone, K. Cleary, R. N. Dumoulin et al., The q/u imaging experiment instrument, Atrophysical Journal 768 (2013) 9 [1207.5562].
- [14] S. Mandelli, E. Manzan, A. Mennella, F. Cavaliere, D. Viganò, C. Franceschet et al., A chemically etched corrugated feedhorn array for D-band CMB observations, Experimental Astronomy (2021) 1.
- [15] S. Masi et al., QUBIC V: Cryogenic system design and performance, arXiv e-prints (2020) arXiv:2008.10659 [2008.10659].
- [16] G. D'Alessandro, L. Mele, F. Columbro et al., QUBIC VI: cryogenic half wave platerotator, design and performances, arXiv e-prints (2020) arXiv:2008.10667 [2008.10667].
- [17] S. A. Torchinsky et al., QUBIC III: Laboratory Characterization, arXiv e-prints (2020) arXiv:2008.10056 [2008.10056].
- [18] S. A. Torchinsky, J. C. Hamilton, M. Piat, E. S. Battistelli, C. Chapron, G. D'Alessandro et al., QUBIC III: Laboratory Characterization, arXiv e-prints (2020) arXiv:2008.10056 [2008.10056].
- [19] C. O'Sullivan, M. De Petris et al., QUBIC VIII: Optical design and performance, arXiv e-prints (2020) arXiv:2008.10119 [2008.10119].
- [20] M. Piat, G. Stankowiak et al., QUBIC IV: Performance of TES Bolometers and Readout Electronics, arXiv e-prints (2021) arXiv:2101.06787 [2101.06787].
- [21] B. Maffei, E. Gleeson, J. A. Murphy and G. Pisano, Study of corrugated Winston horns, in Millimeter and Submillimeter Detectors for Astronomy II, J. Zmuidzinas, W. S. Holland and S. Withington, eds., vol. 5498, pp. 812 – 817, International Society for Optics and Photonics, SPIE, 2004, DOI.
- [22] L. Piccirillo, P. Ade, M. D. Audley, C. Baines, R. Battye, M. Brown et al., The ClOVER experiment, in Millimeter and Submillimeter Detectors and Instrumentation for Astronomy IV, W. D. Duncan, W. S. Holland, S. Withington and J. Zmuidzinas, eds., vol. 7020 of Society of Photo-Optical Instrumentation Engineers (SPIE) Conference Series, p. 70201E, July, 2008, DOI.
- [23] A. Gomez, G. Murga, B. Etxeita, R. Sanquirce, R. Rebolo, J. A. Rubiño-Martin et al., QUIJOTE telescope design and fabrication, in Ground-based and Airborne Telescopes III, L. M. Stepp, R. Gilmozzi and H. J. Hall, eds., vol. 7733 of Society of Photo-Optical Instrumentation Engineers (SPIE) Conference Series, p. 77330Z, July, 2010, DOI.
- [24] C. A. Wuensche, L. Reitano, M. W. Peel, I. W. A. Browne, B. Maffei, E. Abdalla et al., Baryon acoustic oscillations from Integrated Neutral Gas Observations: Broadband corrugated horn construction and testing, Experimental Astronomy 50 (2020) 125 [1911.13188].

- [25] BICEP2 Collaboration, P. A. R. Ade, R. W. Aikin, D. Barkats, S. J. Benton, C. A. Bischoff et al., Detection of b-mode polarization at degree angular scales by bicep2, Phys. Rev. Lett. 112 (2014) 241101 [1403.3985].
- [26] G. L. James, Analysis and design of te/11/-to-he/11/ corrugated cylindrical waveguide mode converters, IEEE Transactions on Microwave Theory Techniques 29 (1981) 1059.
- [27] P. Clarricoats and A. Olver, Corrugated Horns for Microwave Antennas. Peter Peregrinus Ltd, 1984.
- [28] A. Olver, P. Clarricoats, A. Kishk and L. Shafai, Microwave horns and feeds. IEEE Press, 1994.
- [29] J. Murphy and R. Padman, Radiation patterns of few-moded horns and condensing lightpipes, Infrared Physics 31 (1991) 291.
- [30] J. Murphy, R. Colgan, C. O'Sullivan, B. Maffei and P. Ade, Radiation patterns of multi-moded corrugated horns for far-ir space applications, Infrared Physics & Technology 42 (2001) 515.
- [31] A. R. Kerr and S. Srikanth, The Ring-Centered Waveguide Flange for Submillimeter Wavelengths, in 20th International Symposium on Space Terahertz Technology, 2009.
- [32] A. Berthon and R. P. Bills, Integral equation analysis of radiating structures of revolution, IEEE Transactions on Antennas and Propagation 37 (1989) 159.
- [33] A. Clark, G. Childs and G. Wallace, Electrical resistivity of some engineering alloys at low temperatures, Cryogenics 10 (1970) 295.
- [34] ASM HANDBOOK Properties and Selection: Nonferrous Alloys and Special-Purpose Materials. ASM International, 1990, DOI: https://doi.org/10.31399/asm.hb.v02.a0001060.
- [35] G. Bordier, Développements de composants millimétriques pour la caractérisation de la polarisation du fond diffus cosmologique, phd, Ècole doctorale ED127: Astronomie & Astrophysique d'Île de France Laboratoire Astroparticule et Cosmologie (APC), 2014.
- [36] S. P. Morgan, Effect of Surface Roughness on Eddy Current Losses at Microwave Frequencies, Journal of Applied Physics 20 (1949) 352.
- [37] C.-D. Chen, C.-K. C. Tzuang and S. T. Peng, Full-wave analysis of a lossy rectangular waveguide containing rough inner surfaces, IEEE Microwave and Guided Wave Letters 2 (1992) 180.
- [38] M. Zannoni, A. Baù, A. Passerini, A. Tartari, M. Gervasi and L. Valenziano, A cryogenic set-up for accurate measurements of S-parameters, in Millimeter, Submillimeter, and Far-Infrared Detectors and Instrumentation for Astronomy VI, W. S. Holland and J. Zmuidzinas, eds., vol. 8452 of Society of Photo-Optical Instrumentation Engineers (SPIE) Conference Series, p. 84522R, Sept., 2012, DOI.