

What drives Alpine Tethys opening: clues from the review of geological data and model predictions

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Abstract

Permo-Triassic remnants (300-220 Ma) of high-temperature metamorphism associated with large gabbro bodies occur in the Alps and indicate a high thermal regime compatible with lithospheric thinning. During the Late Triassic-Early Jurassic an extensional tectonics leads to the break-up of Pangea continental lithosphere and the opening of Alpine Tethys Ocean (170-160 Ma), as testified by the ophiolites outcropping in the Central-Western Alps and Apennines. We revise geological data from the Permian to Jurassic of the Alps and Northern Apennines, focusing on continental and oceanic basement rocks, and predictions of existing numerical models of post-collisional extension of continental lithosphere and successive rifting and oceanization. The aim is to test whether the transition from the Permo-Triassic extensional tectonics to the Jurassic opening of Alpine Tethys occurred. We enforce the interpretation that a forced extension of 2 cm/yr of the post-collisional lithosphere results in a thermal state compatible with the Permo-Triassic high-temperature event suggested by pressure and temperature conditions of metamorphic rocks and widespread igneous activity. Extensional or transtensional tectonics is also in agreement with the generalized subsidence indicated by the deposition of sedimentary successions with deepening upward facies occurred in the Alps from the Permian to Jurassic. Furthermore, a rifting developed on a thermally perturbed lithosphere agrees with a hyperextended configuration of the Alpine Tethys rifting and with the duration of the extension necessary to the oceanization. The review supports the interpretation of Alpine Tethys opening developed on a lithosphere characterized by a thermo-mechanical configuration inherited by the post-Variscan extension which affected Pangea during the Permian and Triassic. Therefore, a long-lasting period of active extension can be envisaged for the breaking of Pangea supercontinent, starting from the unrooting of the Variscan belts, followed by the Permo-Triassic thermal high, and ending with the crustal break-up and the formation of the Alpine Tethys Ocean.

Keywords: Alpine Tethys Rifting, Numerical modeling, Permo-Triassic high thermal regime, Variscan collision

1 INTRODUCTION

Continental crustal slices preserving pre-Alpine metamorphic imprints are widely described in the Alps and Apennines. Variscan-age eclogites (430-326 Ma), generated from continental, oceanic and mantle rocks belonging to these slices, suggest a pre-Alpine burial of continental crust in a context of oceanic lithosphere subduction underneath a continental upper plate, followed by continental collision (e.g., Marotta & Spalla, 2007; Spalla et al., 2014; Spalla & Marotta, 2007; von Raumer, Bussy, Schaltegger, Schulz, & Stampfli, 2013). Permo-Triassic records (300-220 Ma) of high-temperature metamorphism mainly occur within Austroalpine and Southalpine domains of the Alps (Figure 1) and are associated with widespread emplacement of large gabbro bodies (Figure 2) and acidic intrusives, together with spinel-plagioclase bearing peridotites. This accounts for an increase in the thermal regime (e.g., Lardeaux & Spalla, 1991; Marotta, Spalla, & Gosso, 2009; Schuster & Stüwe, 2008; Spalla et al., 2014) related to asthenospheric upwelling and lithospheric thinning (e.g., Beardmore & Cull, 2001; Sandiford & Powell, 1986; Thompson, 1981). Late Triassic-Early Jurassic extensional tectonics leads to the break-up of the Pangea continental lithosphere and the opening of the Alpine Tethys Ocean, testified by the occurrence of ophiolitic sequences in the Alps and Apennines (Figure 3).

The geodynamic significance of the Permo-Triassic high temperature and low pressure metamorphic event has been widely debated. Some authors propose a late-Variscan post-orogenic collapse (e.g., Desmons, Compagnoni, Cortesogno, Frey, & Gaggero, 1999; Gardien, Reusser, & Marquer, 1994; Neubauer, Frisch, Schmerold, & Schläöser, 1989; Picazo et al., 2016), active in a mainly strike-slip dominated tectonics (Arthaud & Matte, 1977; Cassinis & Perotti, 1994; von Raumer et al., 2013). From ca. 260 Ma, a thermal relaxation of the lithosphere would have characterized the entire Triassic until ca. 200-180 Ma, when the Mesozoic continental rifting started (Picazo et al., 2016).

Alternatively, numerical models (Marotta & Spalla, 2007; Marotta et al., 2009; Spalla et al., 2014) support the interpretation of a long-lasting lithospheric thinning and extensional tectonics leading to the Mesozoic continental rifting (e.g., Bertotti, Siletto, & Spalla, 1993; Diella, Spalla, & Tunesi, 1992; Lardeaux & Spalla, 1991). Long-lasting extension process has already been proposed for the opening of the Central Atlantic Ocean (from the Permian to Early Jurassic, Piqué & Laville, 1996) and for the Northern Atlantic region (Doré & Stewart, 2002). In the latter, a sequence of rift basins, from Permian to Cretaceous in age, has been described occurring before the opening of the ocean (e.g., Doré & Stewart, 2002), making the rifting of the North Atlantic Ocean a long-lasting process characterized by successive extensional stages associated with a migration of eulerian poles, as testified by the anti-clockwise and successive clockwise rotation of superposed rift axes (Doré & Stewart, 2002).

In this work, we revise the Permo-Jurassic geological data from the Alps and Northern Apennines (Figures 1, 2, 3) and the predictions of existing models of post-collisional extension of continental lithosphere (Marotta et al., 2009) and of rifting and oceanization (Marotta, Roda, Conte, & Spalla, 2018). The goal is to test whether a protracted extension active on a lithosphere thermally and mechanically perturbed by the Variscan subduction and collision, can lead the Mesozoic rifting and the oceanization of the Alpine Tethys. We quantify the impact of extensional tectonics on the thermal regime of the continental lithosphere of Pangea, and we also calculate the melting fraction of the lithospheric mantle to be compared with the different magmatic suits of continental gabbros and ophiolite

2 TECTONIC EVOLUTION AND METAMORPHIC AND MAGMATIC RECORDS

The Alps and Apennines belts developed during the closure of the Mesozoic Tethys along two opposite subduction zones active during Cretaceous-Oligocene time for the Alpine system and from Eocene time to the present day for the Apennines system (e.g., Carminati & Doglioni, 2012; Dal Piaz, 2010; Handy, Schmid, Bousquet, Kissling, & Bernoulli, 2010). Pre-Alpine metamorphic and magmatic records from the Alps occur within the four main tectonic domains, namely Southalpine, Austroalpine, Penninic and Helvetic. The Southalpine domain consists of a south-verging thrust system that has been active since Cretaceous time, involving Palaeozoic continental basement and Permo-Cenozoic cover units, both locally affected by very low-grade Alpine metamorphism (Brack, 1981; Zanoni & Spalla, 2018; Zanoni, Spalla, & Gosso, 2010, and refs therein). The Penninic domain consists of a *mélange* of crustal slices deriving from both pre-Alpine continental and Mesozoic oceanic lithosphere, the latter tectonically sampled from the subducted Alpine Tethys Ocean (e.g., Malatesta, Crispini, Federico, Capponi, & Scambelluri, 2012; Malatesta, Gerya, Crispini, Federico, & Capponi, 2013; Platt, 1986; Polino, Dal Piaz, & Gosso, 1990; Roda, Marotta, & Spalla, 2010; Roda, Spalla, & Marotta, 2012; Spalla, Gosso, Marotta, Zucali, & Salvi, 2010; Spalla, Lardeaux, Dal Piaz, Gosso, & Messiga, 1996; Stöckhert & Gerya, 2005). In contrast, ophiolites do not occur in the Austroalpine domain although they are tectonically coupled together with Mesozoic sediments along the external boundary of the domain, under high-pressure (HP) metamorphic conditions (Lardeaux, 2014; Roda et al., 2012; Spalla et al., 1996). Finally, the Helvetic domain consists of an Europe-verging thrust system that includes basement and cover slices stacked during the late stages of the Alpine continental collision since Tertiary (Bousquet et al., 2004; Dal Piaz, 2010; Polino et al., 1990).

Due to the weak Alpine metamorphism and deformation, the Helvetic and Southalpine domains broadly preserve pre-Alpine metamorphic, structural and stratigraphic imprints. Pre-Alpine imprints and igneous relicts are also preserved in Penninic and Austroalpine domains, even if pervasive Alpine structural and metamorphic reworking shapes them into small-sized relicts, inhibiting the correlation of pre-Alpine structures over a regional scale. These records consist of high-temperature low-pressure (HT-LP) metamorphic relicts and Permo-Triassic gabbro bodies, and postdate structures and metamorphic imprints developed during the Variscan convergence.

Variscan metamorphic imprints are widely recorded in the Alpine continental crust (Figure 1) and consist of eclogite, granulite and amphibolite facies rocks (Spalla & Marotta, 2007).

We integrate the geological data of Alpine ophiolites and age data and tectono-magmatic characterization of ophiolites from the Northern Apennines. The Northern Apennines are characterized by oceanic and continental units (Figure 3) that widely escaped subduction-related metamorphism. Oceanic rocks occur in the Internal and External Ligurian units that are deformed under low-pressure and low-temperature metamorphic conditions (Donatio, Marroni, & Rocchi, 2013; Marroni & Pandolfi, 2007). The Internal Ligurian units consist of ophiolites of Jurassic age overlain by a sedimentary sequence of Late Jurassic to Paleocene age, whereas the External Ligurian units comprise Late-Cretaceous sedimentary *mélanges* including slide-blocks of ophiolites at the base (Marroni, Molli, Ottri, & Pandolfi, 2001; Marroni & Pandolfi, 2007).

2.1 Permo-Triassic metamorphic records

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3 144
4 145 HT-LP metamorphic Permo-Triassic imprints are widely recorded in the Austroalpine and
5 146 Southalpine continental crust. They are rarely detected in the Penninic crust of Western
6 147 Alps, and never in the Helvetic domain (Figure 1). These relicts occur in continental crust
7 148 and mantle rocks and show T-climax mineral assemblages developed under amphibolite
8 149 and granulite facies conditions (Figure 4a and Table 1 of Supporting Information). T_{\max}/P
9 150 ratio is generally high and suggests thermal conditions characteristic of Barrovian and
10 151 Abukuma metamorphic field gradients (Figure 4a). Most of the pressure (P) and
11 152 temperature (T) data lie between the fully relaxed (England & Thompson, 1984) and the
12 153 near-spreading ridge geotherms (Cloos, 1993), suggesting the occurrence of an extra heat
13 154 supply with respect to the normal heat generated by radioactive decay after crustal
14 155 thickening consequent to continental collision (Figure 4a). T_{\max}/depth ratios vary from 25 to
15 156 50°C/km and, in particular, the maximum ratios rapidly increase from 300 to 270 Ma and
16 157 gradually decrease up to 220 Ma (Figure 4b). This distribution has been interpreted as the
17 158 ending of the Permian extensional regime at ca. 260–240 Ma, followed by sag-stage
18 159 subsidence during slow lithospheric cooling from the Early Triassic (Picazo et al., 2016;
19 160 Schuster & Stüwe, 2008). No relation occurs between T_{\max} ages, metamorphic field
20 161 gradients (Barrovian and Abukuma) and their present-day distribution through the Alps
21 162 (Figures 5, 6a).
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2.2 Permo-Triassic continental gabbros

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27 167 Permo-Triassic continental gabbros are located in the Austroalpine and Southalpine
28 168 domains of the whole Alpine realm (Figure 2) and are associated with subcontinental
29 169 peridotites emplaced at different structural levels (Figure 4c and Table 2 of Supporting
30 170 Information) (Bonin et al., 1993; McCarthy & Muntener, 2015; Rottura et al., 1998; Spalla
31 171 et al., 2014; Spiess, Cesare, Mazzoli, Sassi, & Sassi, 2010; Staehle et al., 2001). Although
32 172 the gabbros are generally considered to be generated from variably contaminated mantle
33 173 sources in an extensional tectonic regime under a high thermal state, associated with
34 174 lithospheric thinning and rifting (Spalla et al., 2014, and refs therein), from a geochemical
35 175 point of view, most of the gabbros have a tholeiitic signature (Spalla et al., 2014, and refs
36 176 therein). The peridotites contain clino-pyroxene and spinel with Cr content between 20 and
37 177 40% (McCarthy & Muntener, 2015; Muntener, Manatschal, Desmurs, & Pettke, 2010;
38 178 Nicot, 1977). Such a high Cr content of spinel suggests that Permo-Triassic gabbros result
39 179 from >10% near-fractional melting of peridotites (Hellebrand, Snow, Dick, & Hofmann,
40 180 2001; McCarthy & Muntener, 2015). Experimental results on melt compositions and mantle
41 181 melting percentage at different PT conditions, allow to estimate a maximum pressure of
42 182 partial melting during Permo-Triassic at 1.5 GPa (Jaques & Green, 1980; Winter, 2003).
43 183 Continental gabbros show a broad range of radiometric ages, from 300 to 220 Ma, and no
44 184 relation between ages and the present-day geographic location occurs (Figure 6b).
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2.3 Sedimentary basins and volcanics

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51 189 Three cycles of sedimentary successions and volcanics emplacement have been
52 190 described from the lower Permian to Jurassic (see extended review in Marotta et al., 2009;
53 191 Spalla et al., 2014). The first volcano-sedimentary cycle (lower Permian) is characterized
54 192 by a widespread volcanic activity in the whole Southalpine domain and part of the Helvetic
55 193 and Penninic domains (Bussy, Sartori, & Thélin, 1996; Cannic, Lapierre, Monié, Briquet, &
56 194 Basile, 2001; Maino, Dallagiovanna, Gaggero, Seno, & Tiepolo, 2012), associated with
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3 195 continental clastic deposits (from conglomerates to siltites, Bargossi, Rottura, Vernia,
4 196 Visonà, & Tranne, 1998; Bussien, Bussy, Masson, Magna, & Rodionov, 2008; Cadel, Cosi,
5 197 Pennacchioni, & Spalla, 1996; Cassinis & Perotti, 1994, 2007; Dallagiovanna, Gaggero,
6 198 Maino, Seno, & Tiepolo, 2009; Gretter, Ronchi, Langone, & Perotti, 2013; Maino,
7 199 Dallagiovanna, Gaggero, Seno, & Tiepolo, 2012; Quick et al., 2009; Visonà, Fioretti, Poli,
8 200 Zanferrari, & Fanning, 2007). Oldest volcanic products are 291 Ma-old andesites in the
9 201 northern Dolomites (Visonà et al., 2007), 285-282 Ma-old volcanics in the Southalpine
10 202 basins of the Central Alps (Zanoni & Spalla, 2018), and 291-285 Ma-old relict lamprophyre
11 203 in the Antigorio nappe rocks of the Penninic domain of the western Alps (Bussien et al.,
12 204 2008). The Athesian Platform (Adige/Etsch valley and Dolomites region) is the largest
13 205 calc-alkaline volcanic and subvolcanic complex of the Southern Alps dated at 280-277 Ma
14 206 (Schaltegger & Brack, 2007). Locally, two volcanic events are distinguished within the
15 207 lower Permian volcano-clastic cycle in the Central Alps (Cassinis & Perotti, 1994, 2007;
16 208 Gretter et al., 2013). The Permian sedimentation is found south and southeast of the
17 209 Argentera massif in the Helvetic domain (Faure-Muret, 1955). Magmatism and clastic
18 210 sedimentation in the intracontinental basins have supposed to be originated during large-
19 211 scale strike-slip transtensional tectonics associated with crustal thinning, upwelling and
20 212 partial melting of mantle, and advection of melts and heat into the crust (Schaltegger &
21 213 Brack, 2007) during the transition from the lower Permian Pangea-B to the Late Permian
22 214 Pangea-A (Muttoni et al., 2003; Muttoni, Erba, Kent, & Bachtadse, 2005). Dextral strike-
23 215 slip tectonics related to the subduction of Paleotethys ridge has been proposed by (Gretter
24 216 et al., 2013) to explain the developed of continental basins and intense magmatism during
25 217 the lower Permian. The origin of the Permian sedimentary troughs in the Southalpine
26 218 domain may be interpreted as the result of the initial stages of the Tethyan rifting (Siletto,
27 219 Spalla, Tunesi, Lardeaux, & Colombo, 1993; Winterer & Bosellini, 1981; Zanoni & Spalla,
28 220 2018).

29 221
30 222 A second cycle of volcanics emplacement and shallow-water to marine deposits occurred
31 223 from the Anisian to Norian in the Southalpine domain (Bernoulli, Bertotti, & Froitzheim,
32 224 1990; Bertotti, Picotti, Bernoulli, & Castellarin, 1993; Doglioni, 1987; Gillcrist, Coward, &
33 225 Mugnier, 1987; Lemoine & Trümpy, 1987; Venturini, 1983). The easternmost sector of the
34 226 Southalpine domain displays the most continuous tectono-sedimentary record, from the
35 227 Carboniferous to Triassic. Three changes in the sedimentary sequence are observed and
36 228 they are interpreted as derived from alternating transpressional and transtensional regime
37 229 and fault-controlled ground-level fluctuations (Venturini, 1991). Triassic basaltic
38 230 shoshonites documented in the Dolomites were interpreted as derived from a mantle
39 231 contaminated by a previous subduction (Sloman, 1989) or transpressional to compressive
40 232 tectonics (Castellarin, Lucchini, Rossi, Selli, & Simboli, 1988; Doglioni, 1984). However,
41 233 the generalized subsidence, the lateral variations in thickness of sedimentary successions,
42 234 and the deepening upward of the facies support the occurrence of extensional tectonics in
43 235 the Southalpine domain during the Triassic (Carminati et al., 2010). Furthermore, syn-
44 236 sedimentary normal faults of Triassic age can be detected in the Dolomites (e.g., Doglioni
45 237 & Carminati, 2008). In the Helvetic domain of the Western Alps, east-dipping Triassic listric
46 238 faults are responsible for the deposition of syn-rift sediments with interbedded basalts in
47 239 half-graben structures (Gillcrist et al., 1987; Lemoine & Trümpy, 1987).

48 240
49 241 The third cycle of sedimentation started from early Jurassic (ca. 185 Ma, Berra, Galli,
50 242 Reghellin, Torricelli, & Fantoni, 2009), and is associated with the major rifting stage that
51 243 evolved in the opening of the Alpine Tethys Ocean (Berra et al., 2009).
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2.4 Jurassic ophiolites

In the Alps, the transition from rifting to oceanic spreading is marked by the deposition of post-rift sediments (172-165 Ma, Baumgartner et al., 1995; Bill, O'Dogherty, Guex, Baumgartner, & Masson, 2001; Handy et al., 2010; Stampfli et al., 1998) and by the exhumation and serpentinization of subcontinental and ocean-continent transition zone mantle (Figure 3; e.g., Desmurs, Manatschal, & Bernoulli, 2001; Manatschal, 2004; Manatschal & Müntener, 2009; Picazo et al., 2016; Rampone et al., 2014; Tribuzio, Garzetti, Corfu, Tiepolo, & Renna, 2016). The radiometric ages of the oceanic gabbros and peridotites (Figure 6a and Table 3 of Supporting Information) cluster around approximately 160 Ma (Li, Faure, Lin, & Manatschal, 2013; Mevel, Caby, & Kienast, 1978; Riccardo Tribuzio et al., 2016), with older values of 166-183 Ma from the Apennines, Corsica and Erro-Tobbio ophiolitic units (Li, Faure, Rossi, Lin, & Lahondère, 2015; Rampone et al., 2014; R. Tribuzio, 2004). Oldest ages of 198 ± 22 (Sm-Nd on gabbro) have been obtained by Costa & Caby (2001) in ophiolites from the Western Alps (Chenaillet) and interpreted by the authors as the signature of lithospheric extension announcing the oceanic spreading. However, Li et al. (2013) presented new age for the Chenaillet at ca. 165 Ma.

3 NUMERICAL MODELING

We here review the results obtained from two numerical models that simulate the evolution of the Pangea lithosphere from late-collisional evolution of the Variscan chain (Marotta et al., 2009) to the Jurassic opening of the Alpine Tethys (Marotta et al., 2018). The first model (Permo-Triassic extension - MOD1, Marotta et al., 2009) accounts for the evolution of a thermally and mechanically perturbed lithosphere after 70 Myr the end of the Variscan collision (ca. 360 Ma, Franke, 2000; Lardeaux et al., 2014; Matte, 2001; von Raumer et al., 2013), i.e. from 290 to 220 Ma (Marotta et al., 2009). The second model (Jurassic rifting and oceanization - MOD2, Marotta et al., 2018) accounts for the rifting of a mechanically unperturbed continental lithosphere until the crustal break-up and the formation of the oceanic crust (Marotta et al., 2018). The two models have independent initial conditions, and the initial lithospheric structure of MOD2 is not affected by previous events. Therefore, it predicts the evolution of the rifting only considering the thermal perturbation of the lithosphere, and not the presence of mechanical heterogeneities inherited by MOD1. We compare model predictions with geological data from Permo-Triassic metamorphic and intrusive rocks from the Alps and Jurassic ophiolites from the Alps and the Northern Apennines.

3.1 Model setup

The numerical models were performed by using SubMar code (Marotta, Spelta, & Rizzetto, 2006) and successive versions (Marotta et al., 2018; Regorda, Roda, Marotta, & Spalla, 2017; M. Roda et al., 2012). The dynamics of the crust-mantle system were investigated by numerical integration of the three fundamental equations of conservation of mass, momentum and energy, which include the Extended Boussinesq approximation (e.g., Christensen & Yuen, 1985) for incompressible fluids. The numerical code uses the penalty function formulation to integrate the equation for the conservation of momentum and the streamline upwind Petrov-Galerkin method (used to integrate the equation for the conservation of energy). The marker in-cell technique was used to compositionally

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3 297 differentiate crust and mantle rocks. A viscous-plastic behavior was assumed for all
4 298 materials (Marotta et al., 2006). The material parameters are listed in Table 1. In both
5 299 models, boundary conditions are defined in terms of velocity and temperature. A
6 300 temperature of 300 K (ca. 27°C) is fixed at the top of the crust and 1600 K (ca. 1330°C) is
7 301 fixed at the base of the model (Figure 7). Zero flux is assumed through the lateral sides of
8 302 the model. Velocity boundary conditions and initial thermal structures vary in the two
9 303 models and are presented in the following sections.
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11 305 12 306 **3.2 Permo-Triassic extension - MOD1** 13 307

14 308 MOD1 lasts 70 Myr, from 290 to 220 Ma. The initial configuration accounts for the
15 309 temperatures and material distribution obtained after 2500 km of oceanic subduction from
16 310 425 to 362.5 Ma (e.g., Tait, Bachtadse, Franke, & Soffel, 1997; von Raumer, Stampfli, &
17 311 Bussy, 2003), followed by continental collision and a post-collisional gravitational evolution
18 312 that lasts 72.5 Ma (Marotta et al., 2009). Four different boundary conditions are tested for
19 313 this model to quantify the influence of the extension rate on the thermo-mechanical setting.
20 314 An extension rate of 0 cm/yr is used to simulate a late-orogenic gravitational collapse
21 315 without any far-field forcing (mode 1 collapse, Rey, Vanderhaeghe, & Teyssier, 2001). In
22 316 the other cases, extension rates of 0.5, 1 and 2 cm/yr, respectively, are applied at the right
23 317 boundary (mode 2 collapse, Rey et al., 2001) through the crustal thickness of the upper
24 318 plate (from 0 to 30 km depth). Here, the largest amount of subducted crustal material
25 319 occurs and therefore it represents the favored location for the occurrence of the Permo-
26 320 Triassic extension and metamorphism (Marotta et al., 2009) and zero normal stress is
27 321 assumed from 30 to 80 km depth (Figure 7a,b).
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29 323 We compare the thermal state predicted by the four simulations with the thermal state
30 324 suggested by Permo-Triassic PT estimates documented in the Alps. In particular, we plot
31 325 T_{\max}/depth estimates inferred from metamorphic rocks of the Alpine continental crust
32 326 against the geotherms predicted from the simulations for 70 Myr of evolution (Figure 8).
33 327 For each output time, we extract the geotherms along the vertical section characterized by
34 328 highest thermal gradient. For the simulation with an extension rate of 0 cm/yr, the fitting
35 329 between predictions and geological data is obtained for few T_{\max} only (Figure 8a). The
36 330 fitting for simulations with forced extension increases with the extension rate (Figure
37 331 8b,c,e). For an extension rate of 2 cm/yr most of the T_{\max}/depth estimates overlap the
38 332 simulated geotherms (Figure 8d), suggesting a good agreement with the thermal state of
39 333 Alpine continental lithosphere during the Permo-Triassic. However, some T_{\max}/depth
40 334 estimates remain higher than the simulated geotherms. This simulation shows an increase
41 335 in the thermal state (warmer geotherms) for 30-35 Myr (i.e., from 290 to 260 Ma) and a
42 336 subsequent gradual thermal relaxation until 70 Myr (i.e., from 260 to 220 Ma), albeit the
43 337 forced extension remains constant during the entire simulation.
44 338

45 339 The most peculiar character of the Permo-Triassic igneous activity is the widespread
46 340 emplacement of tholeiitic and olivine-tholeiitic gabbro stocks at different structural levels in
47 341 the continental crust (Figure 4c), and the occurrence of basaltic products (see Table 4 in
48 342 Marotta et al., 2009). Therefore, we verify whether the thermal state of the lithospheric
49 343 mantle predicted by different V_e configurations allows the partial melting of peridotite.
50 344 When the partial melting occurs, we infer the melting percentage. For each time step, the
51 345 maximum melt fraction (Mf) of the mantle is calculated on the basis of the parametrized
52 346 melting model from Katz, Spiegelman, & Langmuir (2003):
53 347

$$Mf = \left(\frac{T - T_{solidus}}{T_{liquidus}^{lherz} - T_{solidus}} \right)^{1.5} \rightarrow T > T_{solidus}$$

where T is the temperature of an element belonging to the lithospheric mantle, $T_{liquidus}^{lherz}$ is the temperature of the lherzolite liquidus and $T_{solidus}$ is the temperature of a wet solidus of an enriched peridotite with 0.2% bulk water contents. The latter is compatible with a post-collisional mantle variably contaminated by continental subduction and partially hydrated by Variscan oceanic subduction. Although predictions from the four Ve configurations satisfy the thermal state for mantle partial melting, only the simulation with extension rate of 2 cm/yr produces mantle partial melting >10% (Figure 9). This amount of partial melting is reached after 10 Myr of forced extension and persists until the end of the simulation (70 Myr), reaching a maximum value of 15-17%.

The final thermo-mechanical setting is very different between the four Ve configurations. In the gravitational simulation (Ve=0 cm/yr) both the crustal thickness and the lithospheric thermal state are similar to the initial conditions, while in the forced extension simulations (Ve=0.5, 1 and 2 cm/yr) a strong lithospheric thinning associated with a hot thermal state occurs (Marotta et al., 2009). Localization of maximum shear strain in the gravitational simulation occurs at the crust-mantle interface and mainly under a compressional regime (Figure 10a,c,e). On the contrary, in the forced extension simulations shear strain localization occurs through the entire crust and under an extensional regime (Figure 10b,d,f). The localization of strain occurs cyclically with a periodicity between 10 and 5 Myr (Marotta et al., 2009).

3.3 Jurassic rifting and oceanization - MOD2

MOD2 accounts for the transition from rifting to oceanization and lasts 45 Myr. The model includes the hydration of the uprising mantle peridotite (Table 1) and a constant extension rate fixed to 1.25 cm/yr on both sides of the domain through the crustal thickness, resulting in total extension rate of 2.5 cm/yr, compatible with the magma-poor nature of the rift (Manatschal & Müntener, 2009). Shear-free conditions are prescribed along the top and the bottom boundaries and a sticky air material is included above the crust in order to model vertical movement of the topographic surface. Both crust and lithospheric mantle markers are allowed to exit the model domain since zero normal stress is assumed along vertical boundaries. A weak seed is placed at crust-mantle interface to focus the rifting at the center of the computational domain. Two different thermal settings of the lithosphere represent the initial conditions: a hot and thin lithosphere, characterized by the 1600 K (ca. 1330°C) isotherm located at 80 km depth, and a cold and thick lithosphere, with the 1600 K isotherm located at 220 km depth (Figure 7c). The two configurations result from two different interpretations of the pre-Alpine evolution of the lithosphere. The first one accounts for a thermally perturbed lithosphere consequent to the Permo-Triassic extension (e.g., Marotta & Spalla, 2007; Marotta et al., 2009; Spalla et al., 2014), while the second one is consequent to a Triassic phase of thermal and mechanical requilibration (Picazo et al., 2016; Schuster & Stüwe, 2008).

On the basis of the setup, the model results in a symmetric rifting of the continental lithosphere and shows the exhumation of a serpentized lithospheric mantle (ocean-continent transition zone - OCTZ) for both simulations (Figure 11). However, the onset of the lithospheric thinning strongly depends on the initial lithospheric thermal state (see

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3 396 Figure 3 in Marotta et al., 2018): for a cold and thick (i.e., strong) lithosphere, the thinning
4 397 is very rapid (4.4 Myr) with respect to a hot and thin (i.e., weak) lithosphere (15.4 Myr).
5 398 Similarly, the occurrence of the crustal break-up is faster for a cold lithosphere (7.4 Myr)
6 399 than for a hot lithosphere (approximately 31.4 Myr). For both initial thermal configurations
7 400 of the lithosphere, the exhumation of the serpentinized mantle starts before the oceanic
8 401 spreading and mantle partial melting, making the model compatible with a magma-poor
9 402 rifting, as suggested for the Alpine case (e.g., Manatschal, Lavier, & Chenin, 2015). For
10 403 the hot configuration the thickness of the continental crust sensibly decreases during the
11 404 extension from 30 km to approximately 5 km close to the OCTZ. For the cold configuration
12 405 instead, the crustal thickness decreases from 30 km to approximately 20 km (see Figure 3
13 406 in Marotta et al., 2018).
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15 408 In order to verify whether the thermal state predicted by the two simulations can result in
16 409 the partial melting of the mantle and the formation of the oceanic lithosphere, we estimate
17 410 the time and melting percentage of the lithospheric mantle during the rifting. As for MOD1,
18 411 the melt fraction is still calculated based on the parametrized melting model from (Katz et
19 412 al., 2003), but taking into account a dry solidus of a depleted peridotite. We assume that
20 413 mantle partial melting and oceanic lithosphere start simultaneously. Both simulations result
21 414 in a melting fraction >20%, typical for basaltic melts (Winter, 2003), but at two different
22 415 time steps. For a hot and thin lithosphere it occurs after 38 Myr, while for a cold and thick
23 416 lithosphere after 24 Myr (Figure 12). Both simulations match the emplacement PT
24 417 conditions and paleogeographic setting of oceanic gabbros exposed in the Alps (Table 3
25 418 of the Supporting Information), as demonstrated by (Marotta et al., 2018).

26 419 In addition, considering that a hyperextended system was assumed to the Alpine Tethys
27 420 rifting (e.g., Manatschal et al., 2015) and that the emplacement of oceanic gabbros (ca.
28 421 160-170 Ma, see review in Marotta et al., 2018, 2009) occurred approximately 35-40 Myr
29 422 after the first extensional structures related to the rifting event (200 Ma, Mohn, Manatschal,
30 423 Beltrando, Masini, & Kuszniir, 2012), we indicate that a rifting developing on thermally
31 424 perturbed lithosphere better agrees the natural data.
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35 426 36 427 **4 DISCUSSION** 37 428

38 429 The thermal state predicted by a model of forced lithospheric extension (MOD 1) fits well
39 430 with the thermal state extrapolated from the review of metamorphic and intrusives rocks of
40 431 Permo-Triassic age (Figure 8). This correspondence enforces the interpretation of a
41 432 regional extensional tectonics that characterized the lithosphere of Pangea during the
42 433 Permo-Triassic, causing the asthenospheric upwelling and lithospheric thinning (Diella et
43 434 al., 1992; Lardeaux & Spalla, 1991; Marotta et al., 2009; Muntener & Hermann, 2001;
44 435 Schuster & Stüwe, 2008; Spalla et al., 2014; Spalla & Marotta, 2007). In particular, an
45 436 extension rate of at least 2 cm/yr generates a thermal state in agreement with most
46 437 T_{\max}/depth estimates obtained from Permo-Triassic metamorphic rocks (Figure 8d).
47 438 Notably, in this simulation (Ve 2 cm/yr) the thermal state (T_{\max}/depth ratio) increases within
48 439 the first 30-35 Myr (from 290 to 260 Ma) and then gradually decreases up to the end of the
49 440 simulation (70 Myr, 220 Ma). A similar trend is shown by the T_{\max}/depth ratio extrapolated
50 441 from metamorphic rocks. T_{\max}/depth ratio shows a rapid increase from 300 to 270-260 Ma
51 442 and a successive gradual decrease up to 220 Ma (Figure 4b). This distribution was
52 443 previously interpreted as the result of a Permian extensional regime active until ca. 260-
53 444 240 Ma, followed by sag-stage subsidence active during slow lithospheric cooling from the
54 445 Early Triassic (Picazo et al., 2016; Schuster & Stüwe, 2008). In this review we
55 446 demonstrate that the variation in the thermal state of the lithosphere from 290 to 220 Ma
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3 447 can be explained even by a model of continuous extension from the Permian to Triassic. In
4 448 this model, the long-lasting extension does not result in the continental break-up but rather
5 449 in a uniform thinning of the lithosphere.

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7 451 Within a continuous extensional regime, cyclic variation in the localization of shear strain
8 452 may occur through the entire crust with a periodicity of about 10 Myr (Marotta et al., 2009).
9 453 This periodicity can be interpreted as a local variation in the tectonic regime that affects
10 454 the crust or can be related to the free-slip condition imposed for the top boundary of the
11 455 domain that precludes vertical movement of the crust.

12 456 In the whole Southalpine domain widespread volcanic activity characterized the Permian
13 457 time. In the Dolomites and Carnia, Early to Middle Triassic sediments host dykes and
14 458 pillow basaltic lavas. The sedimentary sequences record periodic continental-marine
15 459 transition during the lower Permian and at the Permo-Triassic boundary in the. Dextral
16 460 strike-slip faulting with an extensional component is compatible with Permian tectonic
17 461 history, although minor sinistral effects are recorded in the Dolomites. In any cases, there
18 462 is general consensus in connecting the Permian calc-alkaline magmatism with the
19 463 lithospheric thinning and strike-slip tectonics (Bergomi, Dal Piaz, Malusà, Monopoli, &
20 464 Tunesi, 2017; Dallagiovanna et al., 2009; Schaltegger & Brack, 2007).

21 465 Extension starts prevailing on strike-slip displacement since the Triassic. The Triassic
22 466 volcanics products in the Dolomites have a mainly calc-alkaline affinity and have been
23 467 interpreted as deriving from the melting of a mantle contaminated by a previous
24 468 subduction (Sloman, 1989) or as the result of crustal contamination of mafic melts.
25 469 Alternatively, the calc-alkaline volcanics are interpreted as produced in a back-arc region
26 470 during the final stages of the Paleotethys subduction (Muttoni et al., 2015). However, the
27 471 lack of Triassic high-pressure rocks in the Alpine region does not support the presence of
28 472 an oceanic subduction.

29 473 Numerical modeling indicates that these cyclic periods of active extension and tectonic
30 474 stasis leading the formation of basins in the upper crust and volcanics emplacement at the
31 475 surface, can be explained by the cyclic variation in the localization of shear strain during a
32 476 continuous extension. The localization of the shear strain in discrete shear zones could
33 477 accommodate the subsidence of upper crust basins and allow the ascent of magma (and
34 478 related volcanics products) from deep-seated mafic intrusions and their consequent
35 479 contamination with crustal material.

36 480

37 481 In the Alpine region, the Permo-Triassic period is also characterized by a widespread
38 482 igneous activity with intrusion of large gabbro bodies through the continental crust.
39 483 Radiometric ages show a broad range of intrusion time, from ca. 300 to ca. 220 Ma, and
40 484 no relation between emplacement age and present-day geographic location can be
41 485 inferred (Figure 6b). Generally, basic intrusions have a tholeiitic and olivine-tholeiitic
42 486 geochemical signature (Spalla et al., 2014) and derives from 10-20% partial melting of
43 487 peridotites (Hellebrand et al., 2001; McCarthy & Muntener, 2015) occurred at depth <50km
44 488 (Jaques & Green, 1980; Winter, 2003). This means that the lithosphere must have been
45 489 hot enough to produce tholeiitic gabbros during the entire Permian and Triassic period. A
46 490 simulation with continuous extension of 2 cm/yr satisfies this condition, resulting in a
47 491 potential mantle melting >10% for almost 60 Myr, from 280 to 220 Ma (Figure 9).

48 492

49 493 The model of rifting (MOD2) of a thermally perturbed lithosphere (i.e., hot, thin and weak,
50 494 Marotta et al., 2018) results in a hyperextended system with the first emplacement of
51 495 oceanic gabbros after 35 Myr from the start of the extension (Figure 11). This scenario fits
52 496 with the hyperextended nature of the Alpine Tethys rifting (Manatschal et al., 2015) as well
53 497 as the time span existing between the first extensional structures related to the rifting (ca.

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3 498 200 Ma, Mohn et al., 2012) and the emplacement of oceanic gabbros (ca. 170-160 Ma,
4 499 Marotta et al., 2018, 2009). Therefore, a thermally perturbed lithosphere can better
5 500 represent the lithosphere of Pangea at ca. 200 Ma. In addition, the good fitting existing
6 501 between geological data and model predictions of post-collisional extension and the ages
7 502 of HT-LP metamorphic rocks and continental gabbros emplacement, clearly indicate the
8 503 persistence of a high thermal state of the lithosphere from the Permian until 230-220 Ma
9 504 (Figures 4, 5, 6). The duration of igneous activity in the Ivrea-Verbanò Zone further
10 505 supports this interpretation (Klötzli et al., 2014; Langone et al., 2017; Peressini, Quick,
11 506 Sinigoi, Hofmann, & Fanning, 2007; Schaltegger et al., 2015; Zanetti et al., 2016).
12 507 Furthermore, a time span of 20-30 Myr (from 230 to 200 Ma) is a too short period to
13 508 complete the thermo-mechanical requilibration of the lithosphere. In contrast with the rifting
14 509 model (MOD2), the post-collisional model (MOD1) does not result in the continental break-
15 510 up and this is mainly due to the heterogeneity that characterize the lithosphere of the
16 511 MOD1 which would represent the heritage of the Variscan subduction and collision. This
17 512 thermo-mechanical setting results in a slower lithosphere thinning. Furthermore, lack of a
18 513 weak seed in the model setup leads a rather uniform extension of the lithosphere. The
19 514 ophiolites of the Alps and Apennines have an age younger than 175 Ma (Fig. 6a), that is in
20 515 agreement with the emplacement time of the oceanic gabbros in the model with a
21 516 thermally perturbed lithosphere (>40 Myr from the start of the extension, i.e. <180 Ma).
22 517

23 518 The Permo-Triassic HT-LP metamorphic rocks and the gabbro bodies are concentrated in
24 519 the Austroalpine and Southalpine domains and lack in the Helvetic domain (Figures 1, 2),
25 520 supporting the interpretation of an asymmetric rifting (e.g., Lardeaux & Spalla, 1991;
26 521 Marotta et al., 2009). On the contrary, in the presented review, the model (MOD2)
27 522 simulates the extension of a homogeneous lithosphere resulting in a symmetric rifting.
28 523 However, a rifting developed on heterogeneous lithosphere and upper mantle
29 524 characterized by structures inherited from previous tectonics events (i.e., Variscan
30 525 subduction and collision), can be strongly asymmetric (e.g., Manatschal et al., 2015;
31 526 Marotta et al., 2009; Petersen & Schiffer, 2016; Spalla et al., 2014) even if the far field
32 527 extension is symmetrically applied on both margins.
33 528

34 529 We propose that after the Variscan orogeny, a long-lasting extensional tectonics, starting
35 530 in the Permian (ca. 290-300 Ma), resulted in the Mesozoic rifting and breaking of Pangea
36 531 supercontinent (Figure 13). This tectonic context can explain the Permo-Triassic high
37 532 thermal regime and the opening of the Alpine Tethys as a hyperextended magma-poor
38 533 margin. This process can be characterized by alternated periods of active extension and
39 534 tectonic stasis in the upper crust until the formation of the ocean, as proposed for the
40 535 Northern Atlantic rifting (Doré & Stewart, 2002) or as envisaged for the Ivrea-Verbanò
41 536 Zone (as documented by three metamorphic ages, Permian, Triassic and Jurassic;
42 537 Langone & Tiepolo, 2015). The strong structural asymmetry resulting from the thermal and
43 538 mechanical perturbation induced by the convergence-related processes related to the
44 539 Variscan orogeny in the whole upper mantle, can lead the occurrence of an asymmetric
45 540 rifting of Pangea in the Alpine region.
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48 543 **5 CONCLUSIONS**

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50 545 The review of Permo-Triassic to Jurassic geological data from the Alps and the Northern
51 546 Apennines and the predictions of two existing numerical models that simulate the evolution
52 547 of the lithosphere from the late collisional events of the Variscan orogeny to the Jurassic
53 548 opening of the Alpine Tethys supports the interpretation of an Alpine Tethys rifting and

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3 549 oceanization developed on a lithosphere characterized by a thermo-mechanical
4 550 configuration consequent to a post-Variscan extension, which affected Pangea during the
5 551 Permian and Triassic. Therefore, a long-lasting period of continuous active extension (2
6 552 cm/yr) can be envisaged for the breaking of Pangea, starting from the unrooting of the
7 553 Variscan belts (ca. 300 Ma), followed by the Permo-Triassic thermal peak (as suggested
8 554 by HT-LP metamorphism and emplacement of gabbros), and ending with the crustal
9 555 break-up and the formation of the Alpine Tethys Ocean (ca. 170-160 Ma). This continuous
10 556 process of lithospheric extension led to thinning and break-up of Pangea through the
11 557 occurrence of cyclic periods of active extension and tectonic stasis in the upper continental
12 558 crust, until the formation of the oceanic crust. On the other hand, PT estimates from the
13 559 rocks exhumed from the pre-Alpine lower crust indicate the persistence of a high T/P ratio
14 560 at deep structural level from the Permian to Jurassic.
15 561

16 562 The occurrence of a heterogeneous lithosphere and a strongly thermally and mechanically
17 563 perturbed upper mantle inherited from Variscan convergence promoted the development
18 564 of an asymmetric rifting, as suggested by the concentration of Permo-Triassic HT
19 565 metamorphic rocks and gabbro bodies in the southern margin of the Alpine Tethys.
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	Continental crust	Mantle	Serpentiniz. mantle Weak seed	Sticky air
Composition	66% gneiss, 33% granite	100% dunite	100% serpentine	-
Ref. Density [kg/m ³]	2640	3200	3000	1180
Heat product. [10 ⁻⁶ W/m ³]	2.5	0.002	0.002	-
Thermal conduct. [W/(mK)]	3.06	4.15	4.15	0.026
Rheology	dry granite	dry dunite	serpentinite	sticky air
Activation energy [kJ/mol]	123	444	-	-
Ref. Viscosity [Pa s]	3.47e ²¹	5.01e ²⁰	1e ¹⁹	1e ²⁰
MOD	1-2	1-2	2	2
References	a,b,c	a,b,d	e,b	a,b

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Table 1: Material parameters used for the numerical modeling after and Marotta et al. (2018). References: (a) Best & Christiansen (2001); Dubois & Diament (1997); (b) Rybach (1988); (c) Ranalli & Murphy (1987); (d) Chopra & Paterson (1981); (e) Arcay, Tric, & Doin (2005); Honda & Saito (2003); Regorda et al. (2017); Roda, Marotta, & Spalla (2010).

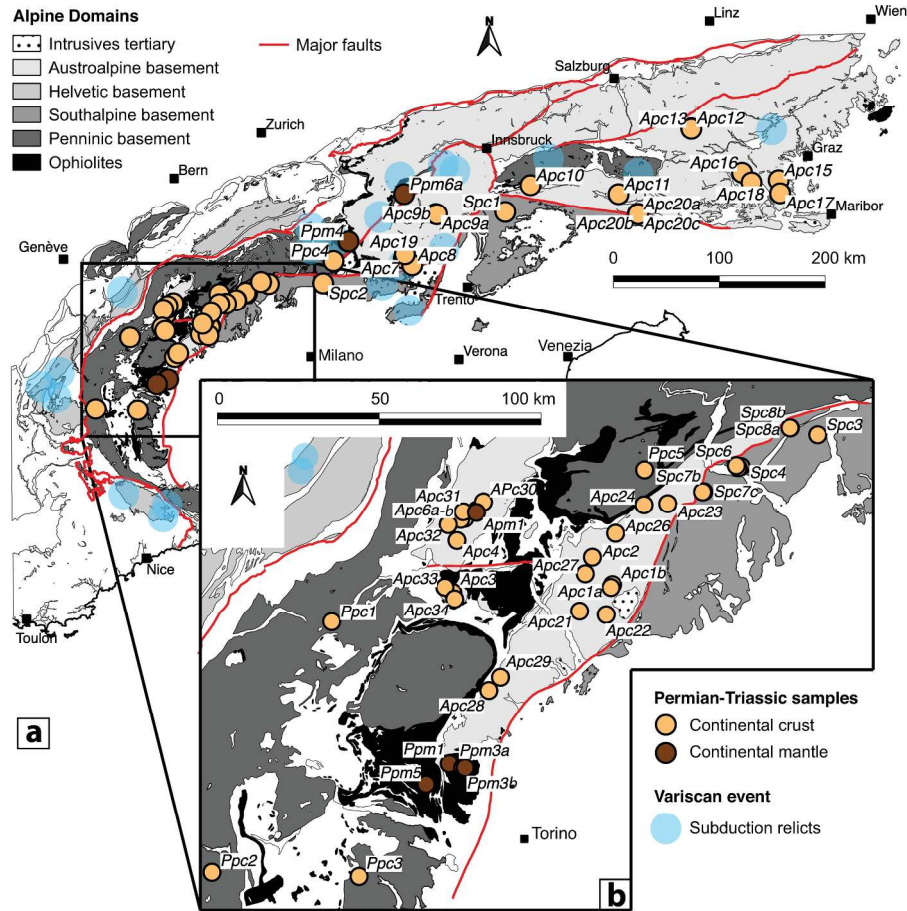


Figure 1 (a) Location of Permo-Triassic metamorphic samples in continental crust and mantle in the Alps. (b) Focus on the Western Alps. Sample codes, geological locations, lithologies, PT estimates, ages (where present) and references are present in Table 1 of Supporting Information.

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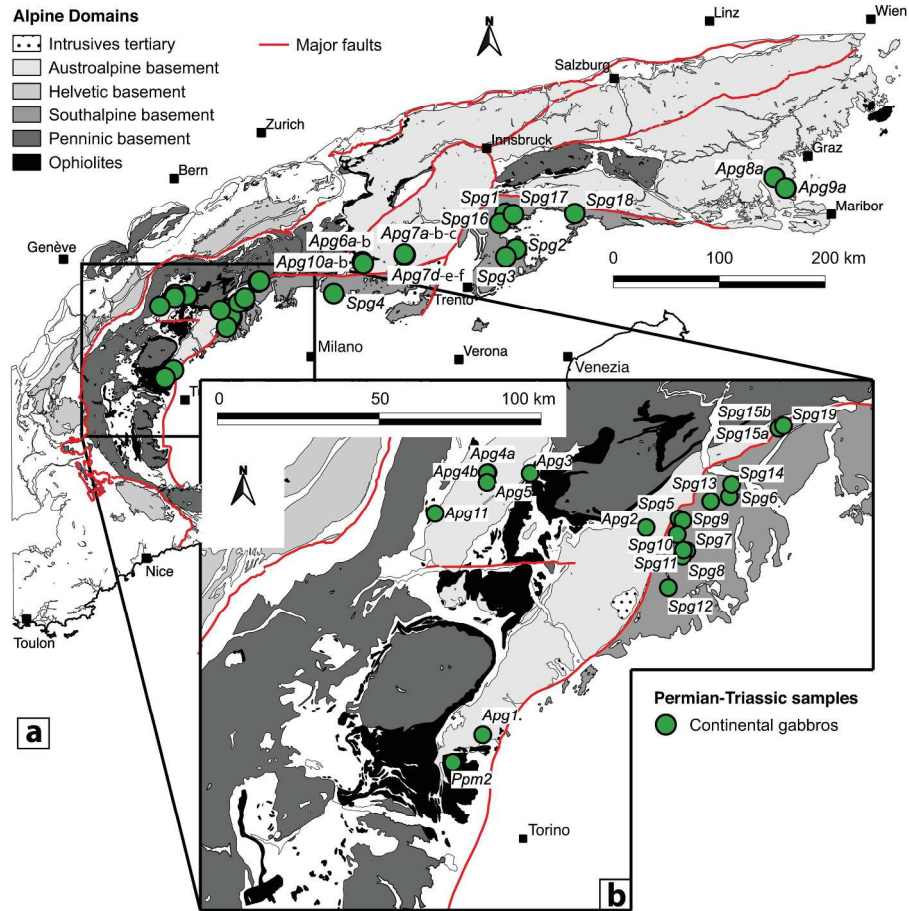
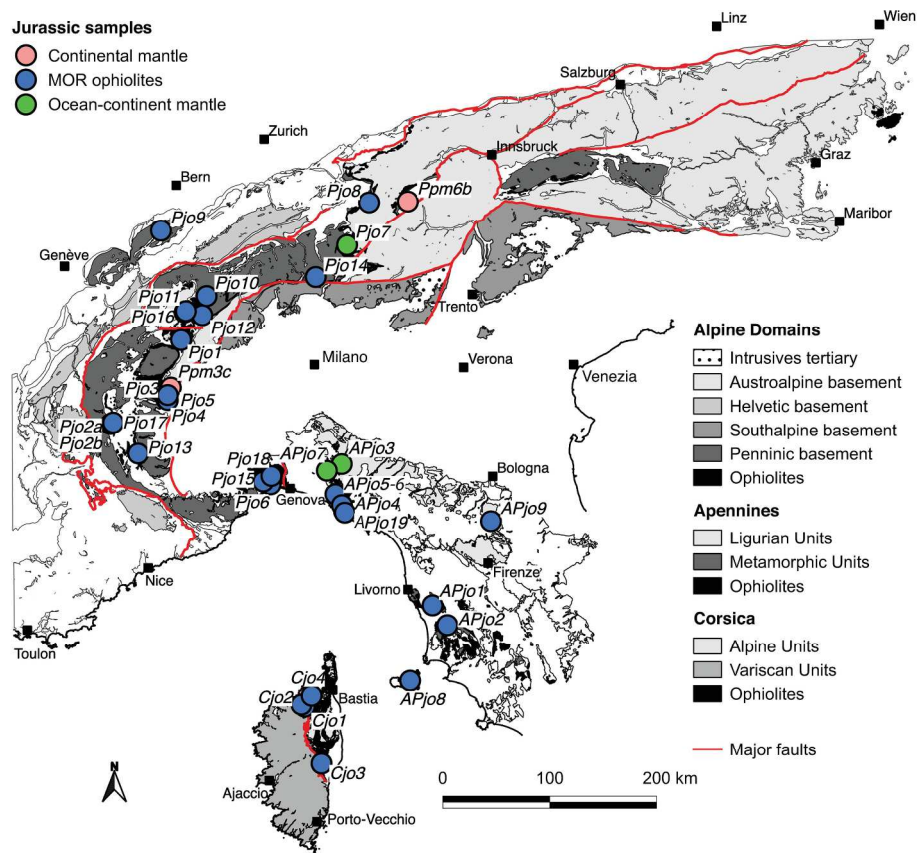


Figure 2 (a) Location of Permo-Triassic continental gabbros in the Alps. (b) Focus on the Western Alps. Sample codes, geological locations, lithologies, PT estimates, ages (where present) and references are present in Table 2 of Supporting Information.

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34 Figure 3 Location and geodynamics setting of Jurassic ophiolites and continental mantle samples in the Alps,
 35 Corsica and Northern Apennines. Sample codes, geological locations, lithologies, PT estimates, ages (where
 36 present) and references are present in Table 3 of Supporting Information.

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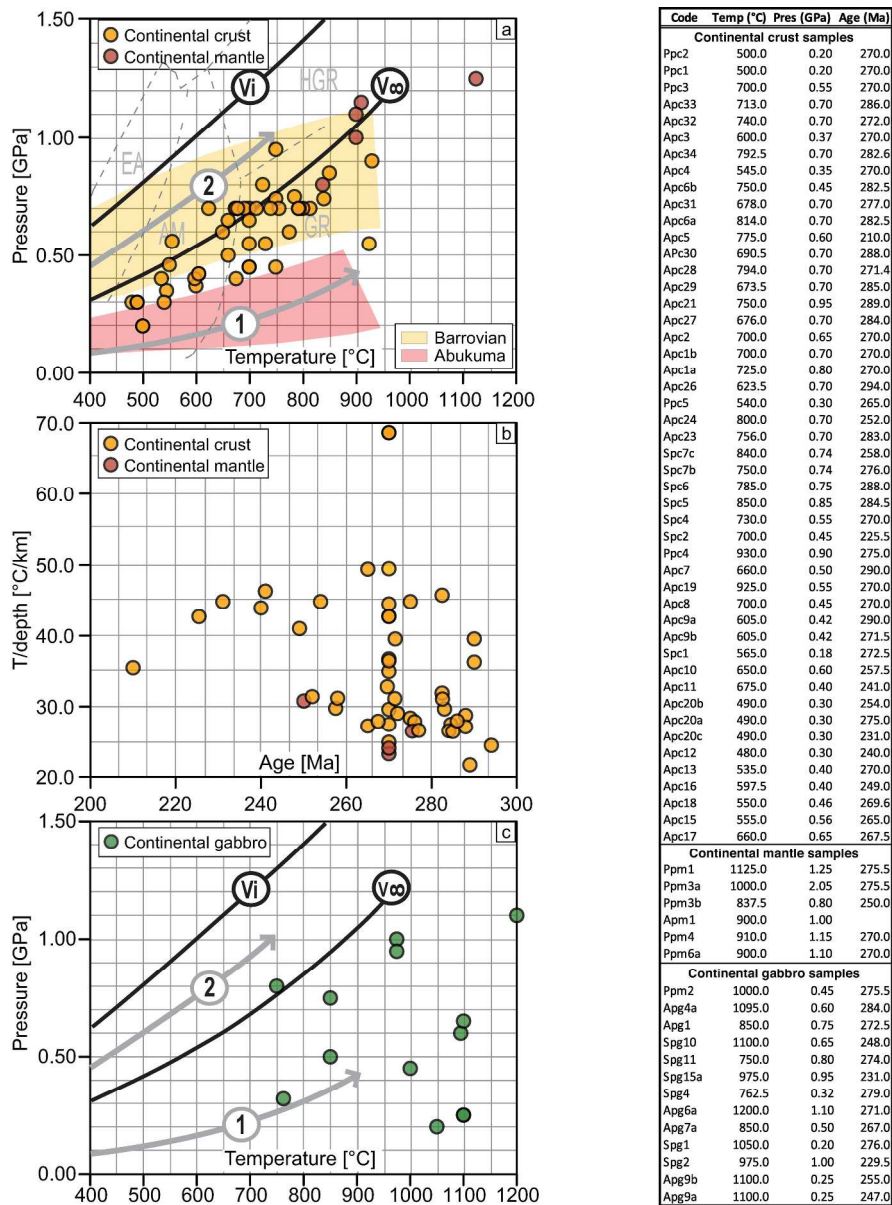


Figure 4 (a) PT conditions of Permo-Triassic metamorphic rocks from the Alps compared to Barrovian and Abukuma metamorphic field gradients and facies (after Ernst and Liou, 2008). Most of the PT data lie between the fully relaxed (2) and the near-spreading ridge (1) geotherms. EA = epidote amphibolite; AM = amphibolite; HGR = high pressure granulite; GR = granulite. Gray lines refer to: (1) near spreading ridge or volcanic arc and (2) normal gradient of old plate interior (Cloos, 1993); Black lines refer to Vi = stable and V∞ = relaxed geotherms for the continental crust (England & Thompson, 1984). (b) T/depth ratio vs age of Permo-Triassic metamorphic crust and mantle samples. The ratios vary from 25 to 50°C/km and rapidly increase from 300 to 270 Ma and gradually decrease up to 220 Ma. Depths are obtained using a constant density of 2800 kg/m³ for the crust and 3000 kg/m³ for the mantle, which represent the average values on the basis of the simulated lithostratigraphy. (c) PT conditions of Permo-Triassic gabbros from the Alps. EA = epidote amphibolite; AM = amphibolite; HGR = high pressure granulite; GR = granulite. Gray lines refer to: (1) near spreading ridge or volcanic arc and (2) normal gradient of old plate interior (Cloos, 1993); Black lines refer to Vi = stable and V∞ = relaxed geotherms for the continental crust (England & Thompson,

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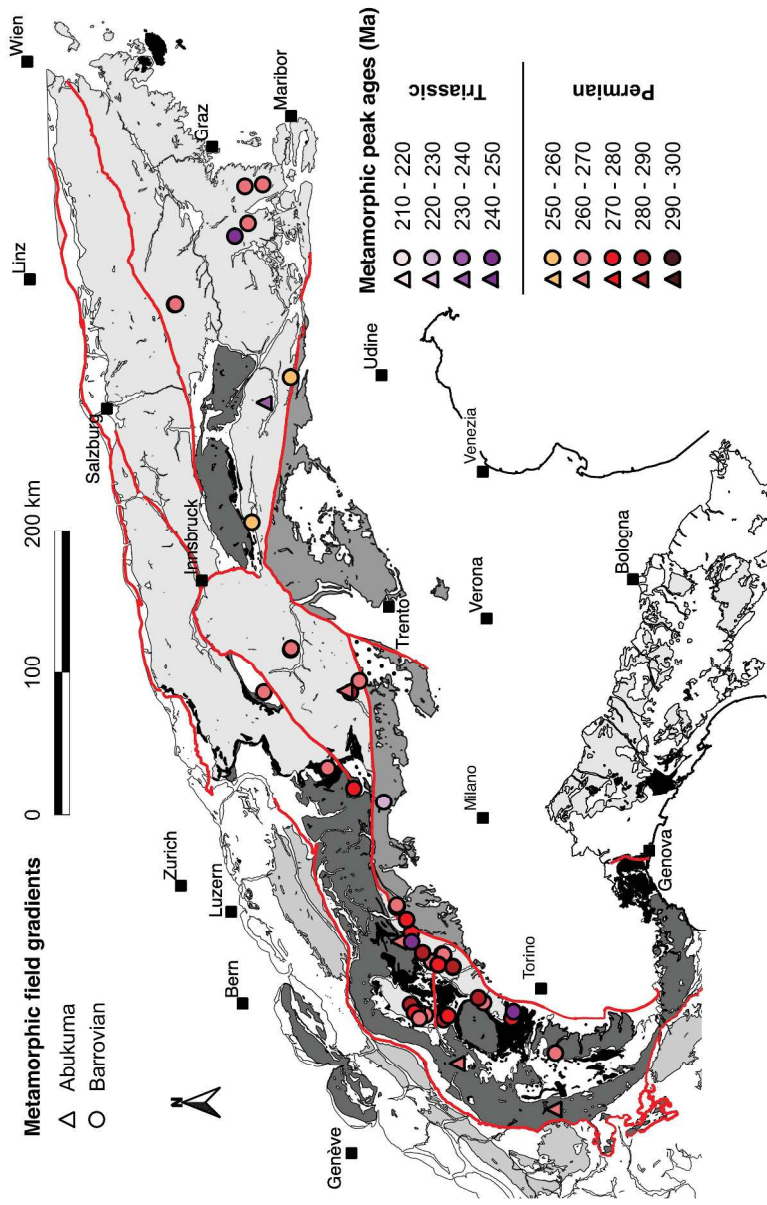


Figure 5 Geographic distribution of metamorphic field gradients (Barrovian and Abukuma) and their Permo-Triassic peak ages in the Alps. No relation occurs between metamorphic field gradients, ages, and their present-day distribution through the Alps.

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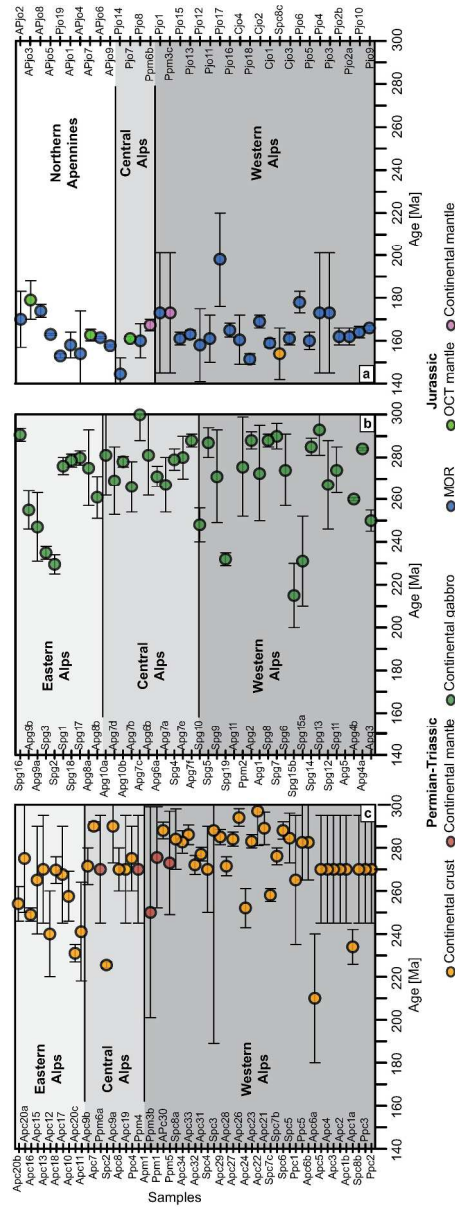


Figure 6 Ages of metamorphic and igneous events recorded in Permo-Triassic metamorphic rocks (a), continental gabbros (b) and ophiolites (c) from the Alps and Northern Apennines. Sample codes, geological locations, lithologies, PT estimates, ages (where present) and references are present in the tables of Supporting Information.

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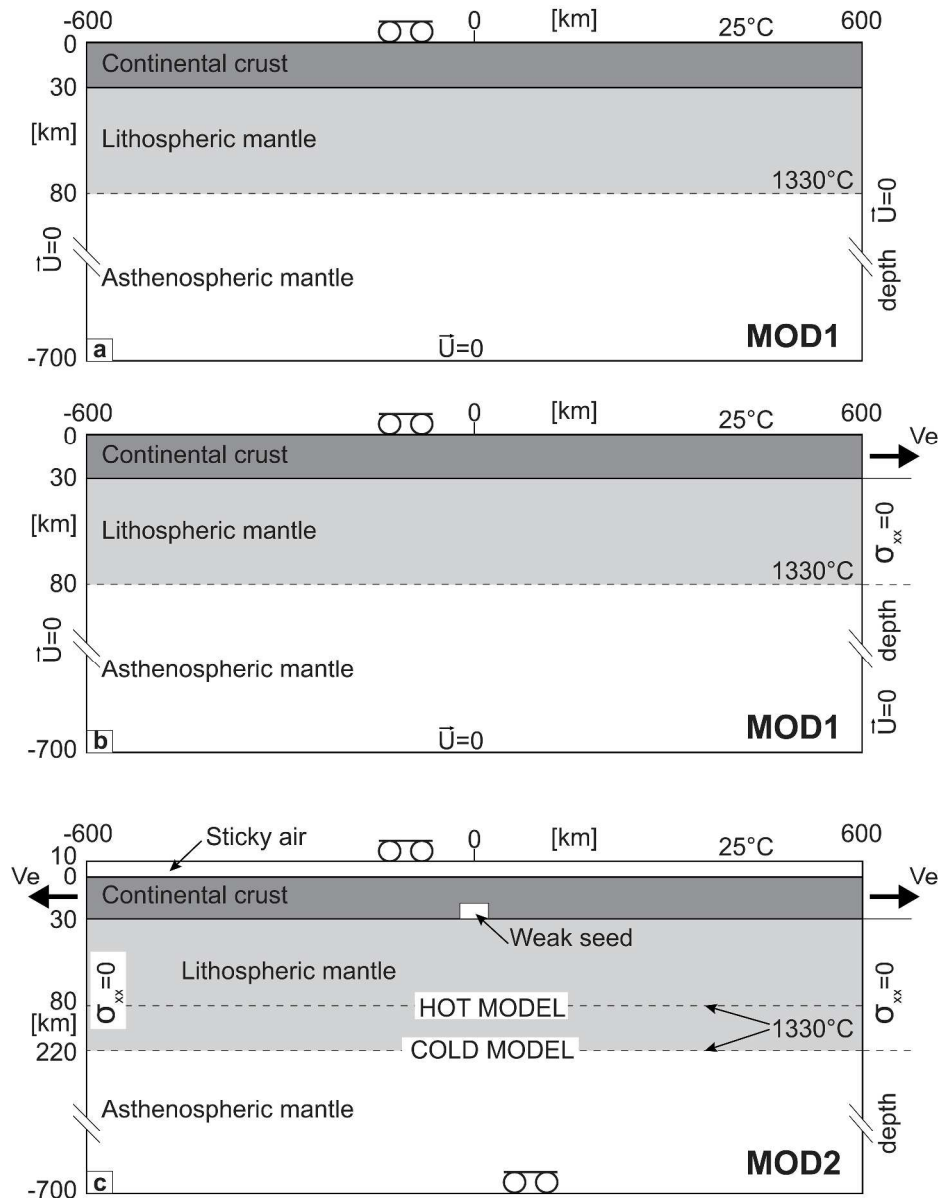


Figure 7 Model setups. (a,b) MOD1 lasts 70 Myr, from 290 to 220 Ma. Different extension rates (V_e) of 0 (a - gravitational collapse), 0.5, 1 and 2 cm/yr (b - forced extension) are applied through the crustal thickness of the upper plate at the right boundary. Zero normal stress (σ_{xx}) is assumed from 30 to 80 km depth along the right boundary while zero velocity (U) is assumed along the left and bottom boundary. A shear-free condition is applied at the top boundary. Initial thermo-mechanical conditions are obtained after 70 Myr of gravitational re-equilibration at the end of the Variscan subduction and collision. (c) MOD2 lasts 45 Myr. The extension rate (V_e) is constant and fixed to 1.25 cm/yr on both sides of the domain (total extension rate of 2.5 cm/yr) and shear-free conditions are prescribed at the top and the bottom boundaries. Two different initial thermal settings of the lithosphere are considered: a hot and thin lithosphere characterized by 1600 K (ca. 1330°C) isotherm located at 80 km depth, and a cold and thick lithosphere with 1600 K isotherm located at 220 km depth.

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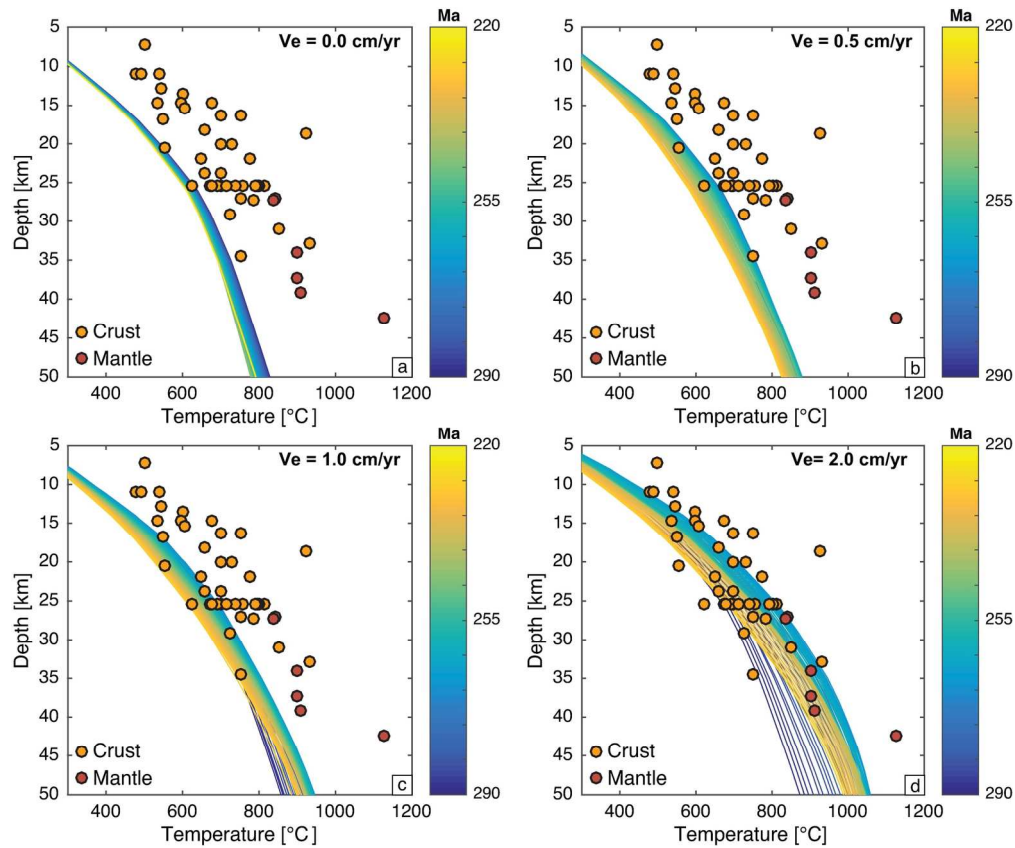


Figure 8 Geotherm curves predicted by MOD1 for 0, 0.5, 1.0 and 2.0 cm/yr extension rates during 290-220 Ma interval, compared to T-depth estimates from Permo-Triassic metamorphic rocks (same samples of Figure 4a,b). We extracted the geotherms along the vertical section characterized by the highest thermal gradient. Depths are obtained using a constant density of 2800 kg/m³ for the crust and 3000 kg/m³ for the mantle (i.e. average values on the basis of the simulated lithostratigraphy). For an extension rate of 0 cm/yr, the fitting between predictions and geological data is poor. The best fit occurs for an extension rate of 2 cm/yr.

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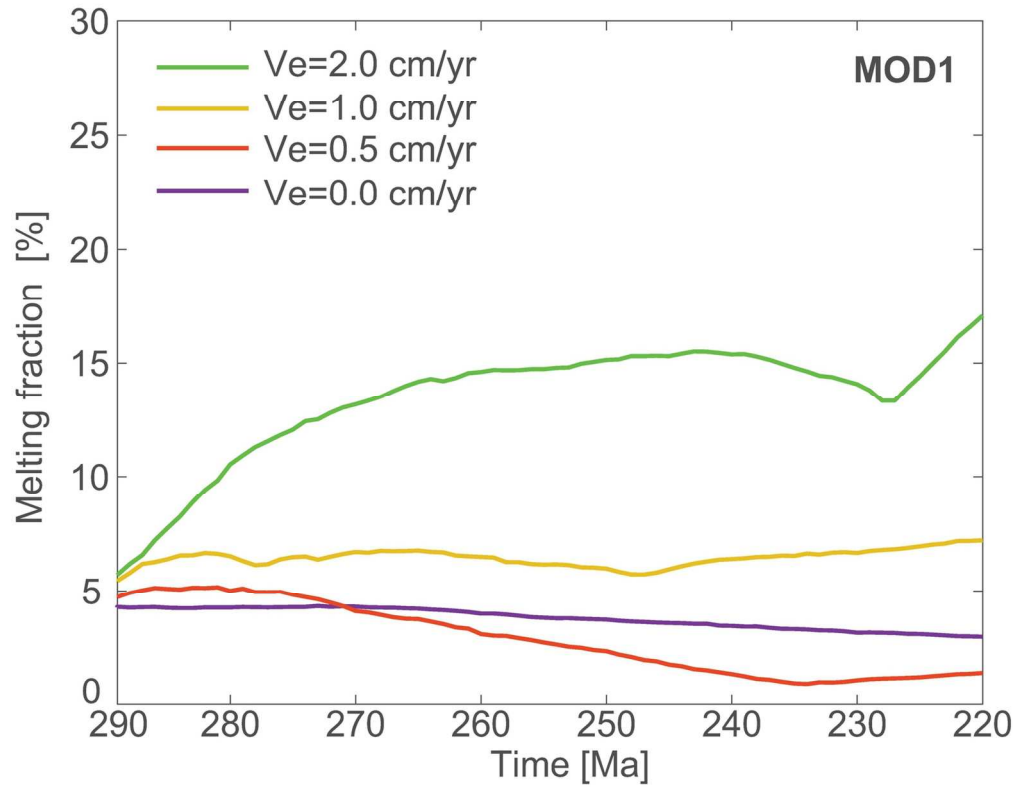


Figure 9 Melting fraction predicted from 290 to 220 Ma by MOD1 at four different extension rates. Although predictions from all configurations satisfy the thermal state for mantle partial melting, only the simulation with an extension rate of 2 cm/yr produces mantle partial melting >10%, reaching a maximum value of 15-17%.

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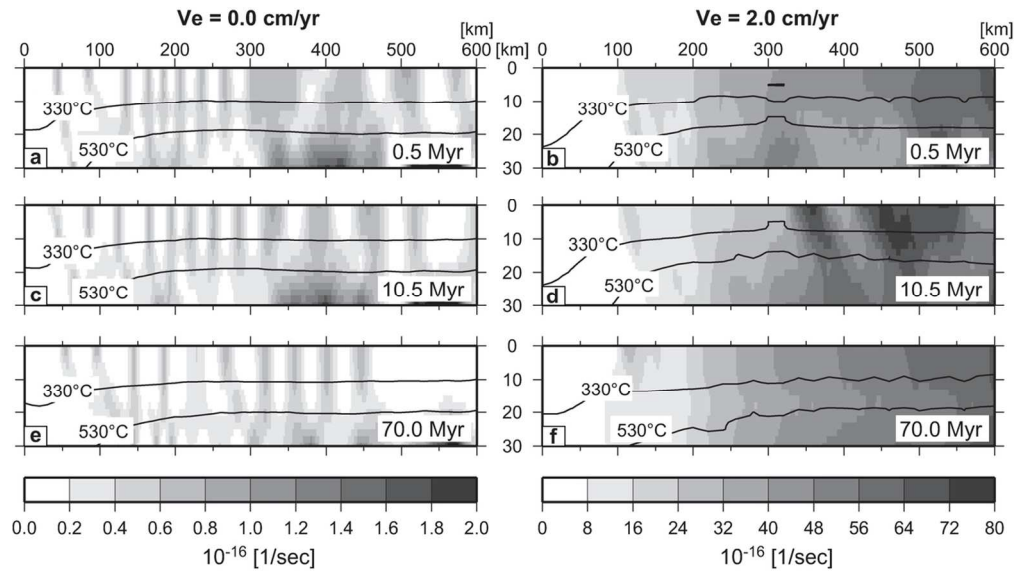


Figure 10 Variation in time (0.5, 10.5 and 70.0 Myr) of maximum strain rate at the lithospheric level predicted by model MOD1 with an extension rate of 0.0 cm/yr (a, c and e) and 2.0 cm/yr (b, d and f). Solid black lines are isotherms (330 and 530°C). Localization of maximum shear strain in the gravitational simulation occurs at the crust-mantle interface and mainly under a compressional regime. On the contrary, in the forced extension simulations shear strain localization occurs through the entire crust and under an extensional regime. Redrawn after Marotta et al. (2009).

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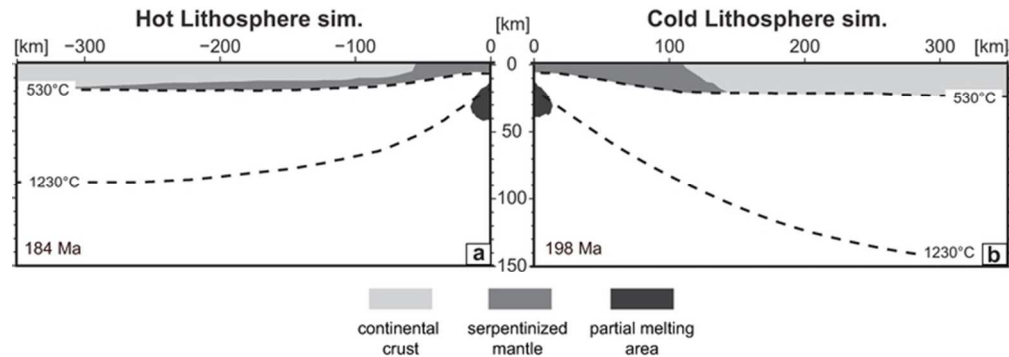
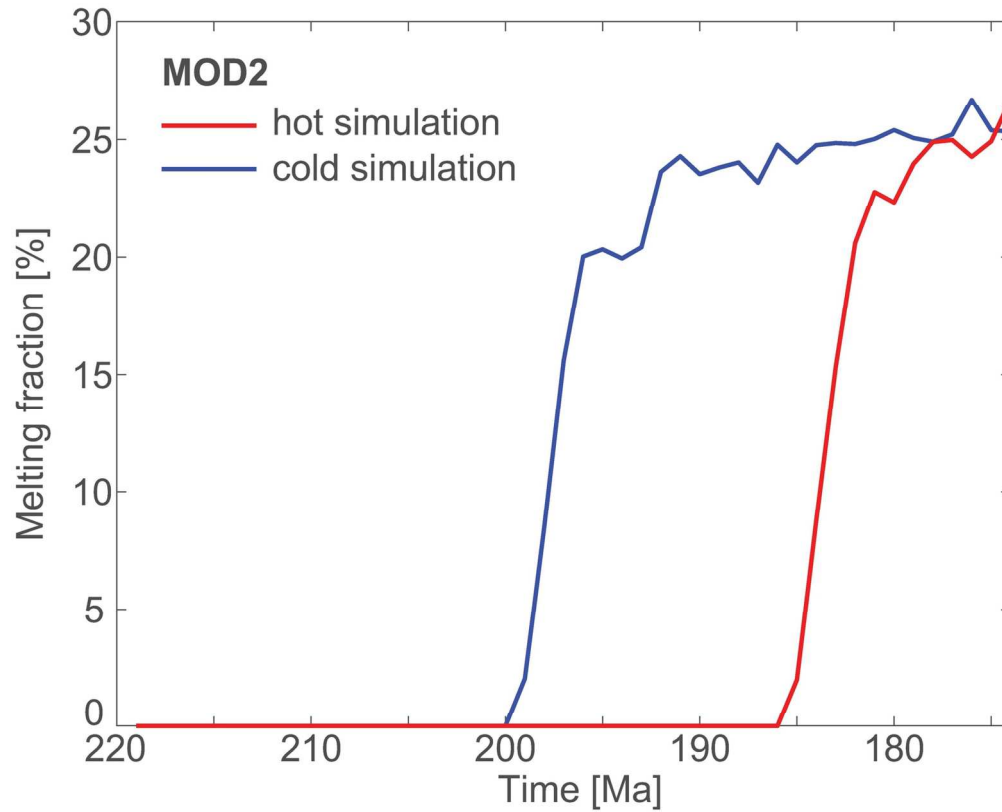


Figure 11 Thermo-mechanical setting predicted by MOD2 model for hot (a) and cold (b) lithosphere simulations at time of mantle partial melting. Dashed lines refer to 800 K (ca. 530°C) and 1500 K (ca. 1230°C) isotherms. In the hot configuration, the thickness of the continental crust decreases up to 5 km close to the OCTZ. In the cold model instead, the crustal thickness decreases up to approximately 20 km. Redrawn after Marotta et al. (2018).

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Figure 12 Predictions of melting fraction in time of the lithospheric mantle for rifting (MOD2) of two different lithosphere thicknesses and thermal states (hot and cold lithosphere). Both configurations reach a melting fraction >20%, typical for basaltic melts, but at two different time steps. For a hot and thin lithosphere it occurs after 40 Myr, while for a cold and thick lithosphere after 22 Myr.

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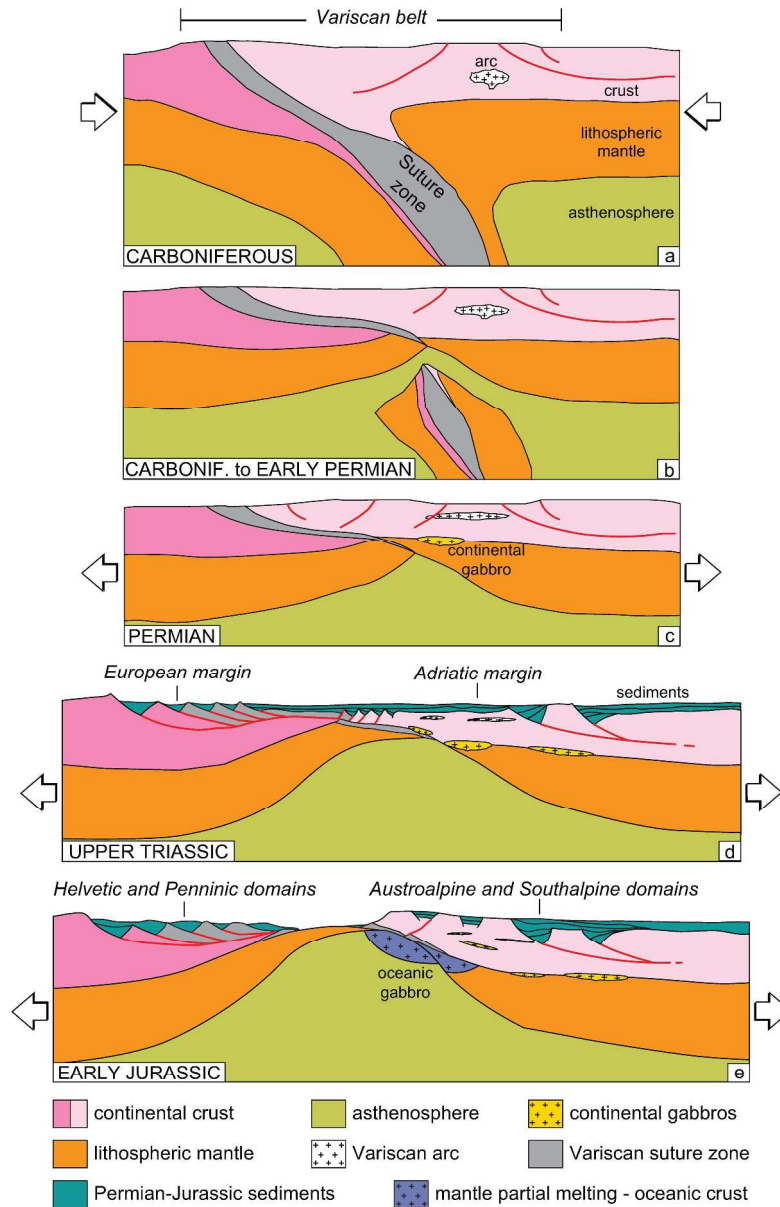


Figure 13 Geodynamic cartoon illustrating some stages of the proposed transition from late collisional slab break-off after the Variscan subduction to Jurassic ocean opening. Variscan magmatic arc after Lardeaux et al. (2014) and Delleani, Rebay, Zucali, Tiepolo & Spalla (2018). Variscan suture zone after Regorda et al. (2017). Continental crust is subdivided as belonging to the European (pale pink) or Adriatic (light pink) margin.

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