

1 **Extreme reduction: Mantle-derived oxide xenoliths from a hydrogen-rich**
2 **environment**

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4 Griffin, W.L.^{1*}, Gain, S.E.M.^{1,2}, Cámara, F.³, Bindi, L.⁴, Shaw, J.², Alard, O.¹, Saunders, M.², Huang, J-
5 X.¹, Toledo, V.⁵ and O'Reilly, S.Y.¹

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7 ¹ARC Centre of Excellence for Core to Crust Fluid Systems (CCFS) and GEMOC, Earth and
8 Planetary Sciences, Macquarie University, NSW 2109, Australia; bill.griffin@mq.edu.au

9 ² Centre for Microscopy, Characterisation and Analysis, The University of Western Australia, WA
10 6009, Australia

11 ³ Dipartimento di Scienze della Terra "Ardito Desio", Università degli Studi di Milano, Via Mangiagalli
12 34, I-20133 Milano, Italy

13 ⁴ Dipartimento di Scienze della Terra, Università di Firenze, Via La Pira 4, I-50121 Florence, Italy

14 ⁵ Shefa Gems Ltd., Netanya 4210602, Israel

15

16 **Abstract**

17 Coarse-grained xenoliths of hibonite + grossite + Mg-Al-V spinel from Cretaceous
18 pyroclastic rocks on Mt Carmel, N. Israel, and from Sierra de Comechingones,
19 Argentina, include spherules, rods and dense branching structures of native
20 vanadium and V-Al alloys. Microstructures suggest that vanadium melts became
21 immiscible with the host Ca-Al-Mg-Si-O melt, and nucleated as droplets on the
22 surfaces of the oxide phases, principally hibonite. Many extended outward as rods or
23 branching structures as the host oxide crystal grew. The stability of V⁰ implies
24 oxygen fugacities ≥ 9 log units below the Iron-Wustite buffer, suggesting a hydrogen-
25 dominated atmosphere. This is supported by wt%-levels of hydrogen in gasses
26 released by crushing, by Raman spectroscopy, and by the presence of VH₂ among
27 the vanadium balls. The oxide assemblage formed at 1400-1200 °C; the solution of
28 hydrogen in the metal could lower the melting point of vanadium to these
29 temperatures. These assemblages probably resulted from reaction between
30 differentiated mafic melts and mantle-derived CH₄+H₂ fluids near the crust-mantle

31 boundary, and they record the most reducing magmatic conditions yet documented
32 on Earth.

33

34 **Introduction**

35 Vanadium is a mildly incompatible, refractory but lithophile transition element; it also
36 can be mildly siderophile, and may have partitioned into Earth's core at high
37 pressure and low oxygen fugacity (fO_2 ; Huang et al., 2015). Its valence states
38 normally range from +5 to +2. At 1500°C the V^{5+}/V^{4+} transition occurs at fO_2 5 log
39 units above the Iron-Wustite buffer ($\Delta IW+5$), and the V^{3+}/V^{2+} transition at $\Delta IW-6$;
40 native vanadium (V^0) becomes stable at ca $\Delta IW-9$, an fO_2 similar to the early solar
41 nebula (Grossman et al., 2008). Natural V^0 has been reported previously only from
42 sublimates in volcanic fumaroles (Ostrooumov and Taran, 2016), and in ophiolitic
43 chromitites (Ifandi et al., 2018).

44 Here we describe V^0 and V-Al alloys trapped as melts in hibonite ($CaAl_{12}O_{19}$)
45 and grossite ($CaAl_4O_7$) that crystallized in Cretaceous super-reduced melt-fluid
46 systems in northern Israel (Griffin et al., 2019) and Argentina. Sampling and
47 analytical methods are described in the Supplementary Data.

48

49 **Background**

50 Most of the material described here comprises xenoliths from the pyroclastic
51 ejecta of Cretaceous (99-85 Ma) intraplate basaltic volcanoes exposed on Mt
52 Carmel, in northern Israel (Fig.1; Griffin et al., 2016, 2018a, 2019a, references
53 therein). At least eight of these volcanoes erupted over an area of ca 150 km², and
54 their xenoliths provide snapshots of similar melt-fluid systems, sampled at different
55 stages of their evolution. Aggregates of skeletal corundum crystals (Carmel

56 Sapphire™) occur as xenoliths in the tuffs, or re-deposited in nearby alluvial placers,
57 mainly in the Kishon River. Melt pockets trapped within and between the corundum
58 crystals contain mineral assemblages requiring high T (>1450-1200 °C), moderate P
59 (ca 1 GPa) and extremely low fO_2 (see below). Detailed evidence for the natural
60 magmatic (i.e. non-anthropogenic, non-electrostatic) origin of the samples has been
61 presented by Griffin et al. (2019b).

62 Paragenetic studies (Griffin et al., 2018a, 2019; Xiong et al., 2017) suggest
63 that both the low fO_2 and the crystallization of skeletal corundum reflect the
64 interaction of differentiated mafic magmas with CH_4+H_2 at high fluid/melt ratios.
65 These fluid-melt systems were sampled by eruption of the host basalts, but were not
66 directly genetically related to them.

67 The earliest paragenesis in the melt pockets of these xenoliths is corundum +
68 tistarite (Ti_2O_3) + carmeltazite ($ZrAl_2Ti_4O_{11}$; Griffin et al., 2018b) + Mg-Al-Ti spinel in
69 a Ca-Mg-Al-Si-O glass. The crystallization of these Ti^{3+} -bearing phases requires fO_2
70 at least 6 log units below the Iron-Wustite buffer ($\Delta IW-6$; Griffin et al., 2016). The
71 original silicate melts were depleted in Fe and Si by the exsolution of immiscible Fe-
72 Ti-C-Si melts, which also occur abundantly as inclusions in the corundum; these
73 melts crystallized moissanite (SiC), TiB_2 , TiN and khamrabaevite (TiC) (Griffin et al.,
74 2016). The metal-silicide melts efficiently scavenged Fe and heavier transition
75 elements, so that no Fe-bearing oxides or silicates are found in the melt pockets.
76 The assemblages described here thus formed from residual melts after extensive
77 crystal fractionation and sequential immiscibility, which enriched the residua in minor
78 elements (REE, Zr, Ti, V, Mn, Sc), Ca and Al. Hibonite first appears late in the
79 crystallization sequence, together with corundum, carmeltazite, Mg-Al spinel, TiN,
80 Fe-Ti silicides, TiC and glass.

81 Coarser-grained (to cm-sized crystals; Fig. 2) intergrowths of hibonite,
82 grossite and spinel, as grains up to 2.5 cm across, also occur in placers of the
83 Kishon and Yoqneam Rivers that drain Mt Carmel. They suggest that hibonite-
84 bearing melts like those trapped in the corundum aggregates evolved even further
85 beneath some volcanic centres. In these xenoliths, V^0 and V-Al alloys occur as
86 spherical, rod-like and branching inclusions, mainly in hibonite.

87 At Mt Carmel, the late crystallization of dmisteinbergite (a polymorph of
88 anorthite) in the melt pockets reflects the peritectic reaction $Liq + Crn \rightarrow An$,
89 equivalent to the incongruent melting of anorthite ($An \rightarrow Liq + Crn$). This well-studied
90 reaction is experimentally constrained to pressures (P) >0.9 GPa and temperatures
91 (T) of ca 1450 °C (Goldsmith, 1980). Thermodynamic modelling in the $CaO-Al_2O_3-$
92 SiO_2 system (Ottonello et al., 2013) suggests that grossite may not be stable at $P >1$
93 GPa. The same modelling shows that the melt-pocket glasses, which clearly were
94 crystallizing corundum (Griffin et al., 2016, 2018a), cannot be in equilibrium with
95 corundum at $P <1$ GPa. These P - T estimates place the proposed site of fluid-melt
96 interaction in the uppermost mantle; the crust-mantle boundary in the area has been
97 geophysically defined at 20-25 km depth (Segev and Rybakov, 2011). Comparisons
98 of mineral parageneses in the melt pockets with experimental studies suggest that T
99 decreased from $>1450^\circ C$ to ca $1200^\circ C$ as fO_2 dropped from $\Delta IW-6$ to $\Delta IW \leq -9$ (Griffin
100 et al., 2018a, 2019a; references therein).

101 We also describe very similar assemblages have been found in an as-yet
102 undocumented locality in the Sierra de Comechingones, Argentina. These samples,
103 obtained from mineral collectors, probably have been taken from similar basaltic
104 cones known to exist in the area.

105

106 **Petrography**

107 The petrography and parageneses of the coarse intergrowths from Mt Carmel
108 are described in detail elsewhere (Griffin et al., 2019a). Resorbed remnants of
109 corundum in hibonite (Mt Carmel) document the peritectic reaction $\text{Crn} + \text{Liq} \rightarrow \text{Hbn}$.
110 The latest paragenesis (Fig. 3a) is grossite + spinel + krotite (CaAl_2O_4) + fluorite \pm
111 the undescribed oxyfluoride phase $\text{Ca}_4\text{Al}_6\text{F}_2\text{O}_{12} \pm$ fluorine-rich perovskite
112 ($\text{CaTi}(\text{O},\text{F})_3$). The Argentinian samples comprise platy cm-sized crystals of V-rich
113 hibonite, in a matrix of grossite, dellagiustaite ($\text{Al}_2\text{V}^{2+}\text{O}_4$ spinel, Cámara et al. 2018),
114 and other phases including $\text{Ca}_4\text{Al}_6\text{F}_2\text{O}_{12}$ (Figs 3b,c).

115 In the Mt Carmel xenoliths, V^0 occurs mainly as spheroidal to amoeboid balls
116 up to mm size in hibonite crystals; some balls show exsolution of Al-rich vs Mn-rich
117 vanadium alloys. More commonly, V^0 forms droplets on (0001) planes, or rods
118 normal to (0001); these rods commonly are necked down to produce linear trains of
119 droplets normal to (0001) (Fig 3a). The inclusions are zonally distributed in some
120 crystals (Fig. 4). In the Sierra de Comechingones samples, V^0 forms elongate
121 inclusions in hibonite, normal to (0001), necked down at irregular intervals; some are
122 partially empty (Fig. 5; Fig. SD1). Spheroidal to amoeboid rods up to 0.1-0.2 mm
123 across are mantled by dellagiustaite or $\text{V}^{3+}/\text{V}^{2+}$ -dominant spinel (Fig. 6). Some
124 examples contain tiny precipitates of Cu, and larger inclusions (Figs 6b,c) can be
125 vesicular, attesting to the presence of volatiles.

126 In the Mt Carmel material, some spherical V^0 inclusions develop into “dense
127 branching structures” (DBS; Goldenfeld, 1989) that grew roughly normal to (0001) of
128 the host hibonite crystal (Fig. 7; Fig. SD2); each filament terminates in a single
129 euhedral crystal, showing that growth took place in contact with melt. Clusters of
130 fine branches terminate at planes parallel to (0001), suggesting pauses in the growth

131 of the hibonite crystal. Some branches continued to grow when crystal growth
132 resumed, but new droplets of V^0 also nucleated on the new crystal plane and grew
133 independently of the older branches. In the example shown in Fig. 7, this stop-start
134 process occurred at least four times. The 3D- μ CT images (Fig. SD2) show that the
135 outer parts of many filaments contain no solid phases, even where they terminate
136 below the present crystal face. These “empty” portions are continuous with the
137 vanadium-filled branches, and make up approximately 3% of the total volume of the
138 branching structures. In the Mt Carmel material, small bulbous protuberances on the
139 sides of some spheroidal V^0 inclusions enclosed in grossite consist of VH_2 , the first
140 reported natural metal hydride (Bindi et al., 2019).

141

142 **Mineral Chemistry**

143 The smallest V^0 inclusions in hibonite from Mt Carmel contain 0.5-2.0% Si and
144 1.5-2.5% Mn (the oxide matrix contains no Si). Most larger inclusions have less Si,
145 1-2% Cr, 0.5-3.5% Mn, and up to *ca* 4 wt% Al; irregular balls included in grossite are
146 the latest to form, and contain up to 15 wt% Al (Table 1). Many of the EMP analyses
147 total <100%, which may reflect H in solid solution (see below). The discovery of
148 vanadium dihydride (VH_2 ; Bindi et al. 2019) among these balls suggests that the
149 deficit could be due to hydrogen in solid solution.

150 EBSD imaging confirms that V^0 in the larger inclusions is cubic, and a TEM
151 diffraction pattern matches the $\langle 110 \rangle$ zone axis of V (James and Straumanis, 1960).
152 Three inclusions of Al-rich vanadium analysed by single-crystal XRD have space
153 group *Im-3m*, consistent with the Al-V phase diagram (Kroupa et al., 2017). The unit
154 cell increases from 3.032 Å to 3.068 Å as Al increases from *ca* 2 at.% to >20 at.%
155 (Fig. 8).

156 V^0 in the Sierra de Comechingones hibonite contains almost no Al, but has
157 more Cr than the Mt Carmel examples, and similar Si contents (0.5-2.2 wt%);
158 vanadium-oxide phases analysed by single-crystal XRD yielded either a tetragonal
159 or a monoclinic cell, corresponding to two polymorphs of synthetic $V_{14}O_{68}$: a β -
160 phase (space group $I4/mmm$), and a γ -phase (space group $C2/m$) with ordered
161 vacancies. The coexistence of these two phases constrains the temperature to
162 <1350 °C (Davydov and Rempel, 2009). These oxides suggest an fO_2 somewhat
163 higher than in the Mt. Carmel rocks, but the presence of V^{2+} in spinel (dellagiustaitite)
164 requires $fO_2 = \Delta IW$ between -6 and -9. Because dellagiustaitite rims are found at the
165 interface between V^0 and hibonite and grossite, this texture might represent a later
166 subsolidus high- T diffusion corona formed at higher fO_2 .

167

168 **Analysis of included fluids**

169 To evaluate the composition of fluids that might be enclosed in the observed
170 voids associated with the vanadium rods and branching structures, four 2-3 mm
171 grains of the hibonite-grossite aggregates and four corundum aggregates from Mt
172 Carmel were analysed for H, C, N, and S contents by ignition in an O_2 -He
173 atmosphere at ca 1150 °C, followed by chromatographic analysis (see
174 Supplementary Data). Contents of H, C, S and N are given in Table 2. The hibonite
175 samples contain high concentrations of gas, averaging 1120 ppm C and 2800 ppm
176 H_2 by weight. This is 4-5 times the concentrations found in the corundum
177 aggregates, even though corundum aggregates appear to have more void space
178 than the hibonite crystals. However, in both cases the included fluids are comprised
179 of >95 at.% hydrogen, assuming that oxygen is at most a minor constituent, as
180 consistent with the required low fO_2 .

181

182 **Raman Spectroscopy**

183 Raman spectra of hibonite from both localities show sharp peaks in the range
184 0-1000 cm^{-1} , indicating that the hibonite is well-crystallized, in contrast to most
185 meteoritic hibonite (e.g. Ma, 2010; Fig. 9). Peak positions agree very well with those
186 observed by Hofmeister et al. (2004) for synthetic hibonite (Table 3). We observe a
187 strong polarization effect on particular Raman shifts. Changes in the Raman shift
188 and intensity of the bands can be ascribed to the small amount of V in the hibonite
189 structure. A broad band at ca 1470 cm^{-1} is present only in N-S orientation of the
190 crystals, along with two shoulders at 1256 and 1588 cm^{-1} . These are difficult to
191 attribute, as plausible groups are ν_3 modes of CO_3 groups, but these are usually
192 very weak in carbonates. Another plausible assignment is to $-\text{CH}_2$ and $-\text{CH}_3$
193 deformations. In the 3000-4150 cm^{-1} spectral range, we observe (also polarized)
194 Raman shifts corresponding to the stretching of O-H bonds. This implies the
195 presence of H in a nominally anhydrous phase. However, the bonding is not well
196 ordered and the bands are rather broad. Bands are also observed in hibonite from
197 Mt. Carmel but at different Raman shifts. The only mode observed in hibonite from
198 both localities is the Raman shift at ca 3950 cm^{-1} , which corresponds to one OH
199 bond with the hydrogen far from any donor oxygen. In the Mt. Carmel hibonite, the
200 band at 3819 cm^{-1} is very strong and this points to a higher content of H diffused
201 within the lattice and defects in the hibonite. This is consistent with the
202 chromatography measurements reported above. The intensity of the bands changes
203 across the crystal, suggesting that the H is related to the degassing of V^0 inclusions,
204 producing voids by volume reduction, as seen in the 3D μ -CT images (Figs SD-1,
205 SD-2), and diffusing the H across the hibonite host.

206

207 **Discussion and Conclusions**

208 The structures of the V^0 inclusions (Figs 2-7) clearly suggest the coexistence
209 of an immiscible metallic melt and a low- α_{SiO_2} silicate melt that was crystallizing Ca-
210 Al oxides and Mg-Al-V spinel. They are best explained by the nucleation of drops of
211 molten vanadium on the faces of oxide crystals exposed to the silicate melt. Many of
212 these drops apparently attracted more V^0 as the crystal grew, producing larger and
213 elongated inclusions, continuously in contact with the surrounding mixture of
214 immiscible melts. These inclusions are thus analogous in many ways to the CO_2 -
215 filled cylindrical inclusions of sulphide melts in clinopyroxene megacrysts from some
216 alkali basalts (Andersen et al., 1987).

217 The dense branching structures (DBS) shown in Figure 7 probably evolved in
218 a different way. They are not crystallographically controlled, but radiate outwards
219 from initial melt droplets (Fig. 7a; Fig. SD2). The formation of such DBS typically
220 reflects diffusion-controlled interface motion such as the growth of a crystal into a
221 fluid under conditions of mild supercooling (Goldenfeld, 1989). Since each branch of
222 the DBS would need to be growing in contact with a fluid, these structures probably
223 grew from the initial ball on a crystal surface, outward into the melt surrounding the
224 hibonite crystal, only to be encased in hibonite as the crystal surface advanced
225 further out into the melt. The obvious planes in Figures 3 and 7b thus could
226 represent stages where the hibonite overgrew and terminated some parts of the
227 expanding DBS, but presented a stable surface for the nucleation of new V^0 droplets.
228 These structures thus reflect an interplay between the growth rates of the V^0 DBS
229 and the hibonite crystal, involving diffusion of several chemical species.

230 Very similar structures, described as “trees” or “cones” have been observed
231 in some high-quality synthetic (HPHT) diamonds grown in metal fluxes (Palynov et
232 al., 1997; Shiryayev et al., 2018); the material in the structures corresponds to Fe
233 alloys or Fe oxides, related to the growth media and/or the sample container. The
234 main differences between these inclusions and the V^0 DBS described here is that
235 many of the branches apparently consist of individual inclusions, and none of the
236 high-quality images provided by Shiryayev et al. (2018) appear to grow from individual
237 larger balls of the metallic material, such as those illustrated in Fig. 7. Both Palynov
238 et al. (1997) and Shiryayev et al. (2018) interpret the structures as bunches of defects
239 related to dislocations in the structure of the host diamond, implying that the metallic
240 phases have nucleated on the defects in a semi-continuous manner as the host
241 crystal grew. In contrast the spherical morphology of most V^0 inclusions described
242 here argues for separation of an immiscible melt, nucleation on the crystal surface
243 and its trapping by the host oxide,

244 However, a defect-related control is plausible in the case of the V^0 inclusions
245 described here. Tiny balls of V^0 appear to have nucleated in abundance on the
246 faces of growing hibonite (and grossite) crystals, but most of those that grow outward
247 form single rods (Fig. 3-5), with a tendency to neck down as seen in fluid-inclusion
248 trails in many minerals. The sites of such nucleation may have represented single
249 dislocations or other imperfections. Only a few inclusions have developed into the
250 spectacular DBS shown in Fig. 7; these may have been the sites of more complex
251 dislocation structures, such as subgrain boundaries (Shiryayev et al., 2018).

252 Inclusions of V^0 occur mainly in hibonite, to a lesser extent in grossite and
253 spinel, and rarely in corundum mantled by hibonite; the metallic melt apparently
254 began separating from the parental oxide melt near the temperature of the

255 corundum-hibonite peritectic, driven by a continuing decrease in fO_2 ($V^{2+} \rightarrow V^0$).
256 Vanadium melts at $>1900^\circ\text{C}$, and experimental data in binary systems suggest that
257 the observed levels of Al, Si, Mn and Cr might lower the melting temperature only by
258 $<100^\circ\text{C}$ (Table 1a, b). Since the estimated temperature of the Cor +L \rightarrow Hib
259 peritectic is close to 1300°C in the CaO-Al₂O₃-MgO-SiO₂ system (de Aza et al.,
260 2000; Jung, 2010), any immiscible V-dominated melt would be expected to freeze
261 upon unmixing, which apparently did not happen.

262 However, the fO_2 required for the presence of a V^0 melt ($\leq \Delta IW -9$) implies a
263 hydrogen-rich environment, similar to that of the early solar nebula in which hibonite-
264 bearing Calcium-Aluminum Inclusions formed (Grossman et al., 2008). The included
265 gasses (Table 2) are dominated by hydrogen, consistent with the inferred low fO_2 ,
266 the evidence of hydrogen defects in the Raman spectra of hibonite, and the
267 presence of VH_2 among the V^0 balls (Bindi et al., 2019).

268 In a hydrogen atmosphere, the melting point of vanadium can be lowered by
269 $>1000^\circ\text{C}$ as several wt % H_2 dissolves in the melt (Fukai, 2012); solubility increases
270 dramatically with increasing pressure. H_2 is expelled during solidification of the melt,
271 but the solid V^0 phase can retain significant levels of hydrogen in solution to
272 temperatures below 1000°C . The consistently low analytical totals of the V^0 balls
273 from both localities may reflect this solid solution.

274 The outer portions of many DBS (Mt Carmel), and some rods (Argentina),
275 contain no solid phases, even though the empty branches clearly are continuous
276 with the vanadium-filled ones. Furthermore, some of the larger droplets in the DBS
277 now comprise an irregular lump of V^0 within a subspherical void (Figures 7, SD2).
278 We suggest that these void spaces reflect the exsolution of H_2 from the solidifying V^0
279 (and contracting) melts in the inner parts of the structures. This could lead to the

280 remelting and expulsion of V^0 from the outer parts of the branches, closer to the front
281 of hibonite nucleation, and perhaps to production of the observed VH_2 phase.
282 Similarly, the vesicles observed in some V^0 grains in the hibonites from Sierra de
283 Comechingones (Fig. 5b) may represent the exsolution of H_2 from a cooling melt.
284 The rimming observed around V^0 inclusions in the Argentinian samples could be
285 related to solid-state diffusion of O and Al into V^0 , forming both the non-
286 stoichiometric vanadium oxides and the rims of V^{3+}/V^{2+} -dominant spinels, while H_2
287 from former VH_2 inclusions diffused into the hibonite host.

288 The petrographic and paragenetic similarities between the assemblages from
289 Mt Carmel and Sierra de Comechingones strongly suggest similar conditions of
290 formation. At present more constraints are available on the Mt Carmel xenoliths,
291 which apparently crystallized late in the pre-eruption evolution of fO_2 in magma-fluid
292 systems. This evolution ultimately led to the most reducing conditions yet found on
293 Earth, with free hydrogen as the fluid phase. The observed abundance of carbon in
294 the assemblages suggests an important role for CH_4 in this process. At Mt Carmel,
295 such conditions existed for at least 10 m.y. over an area of ca 150 km², in the
296 uppermost part of a thin mantle lithosphere, and high fluid/rock ratios would be
297 required to maintain such reduced volumes within the mantle. This suggests the
298 derivation of abundant $CH_4 \pm H_2$ fluids from the deeper mantle, which in turn seems to
299 imply that the sublithospheric mantle beneath this area was metal-saturated (i.e. fO_2
300 = IW), such that any C-O-H fluid issuing from it was dominated by $CH_4 \pm H_2$. This
301 conclusion has implications for understanding melting and metasomatic processes in
302 the mantle (Griffin et al., 2018).

303 As discussed elsewhere (Xiong et al., 2017; Griffin et al., 2018) there is a
304 significant, if scattered, body of evidence in the published literature that similar

305 reducing conditions may exist elsewhere on Earth, in connection with several types
306 of explosive volcanism. Finally, the Al-rich late-forming V⁰ melts reported here
307 represent the first *in situ* observation of Al alloys on Earth, and may add further
308 credence to reports of Al⁰ in many other localities (cf. Dekov et al., 2009).

309

310 **Acknowledgements**

311 We thank Luca Velentini (U. Padova) for the 3D- μ CT images of hibonite from
312 Sierra de Comechingones. We also thank Prof. Andrey Shiryayev and an anonymous
313 referee for useful comments and suggestions, and Prof. Shiryayev for providing high-
314 resolution images of structures in diamonds. Instruments used at Macquarie
315 University are funded by DEST Systemic Infrastructure Grants, ARC LIEF,
316 NCRIS/AuScope, industry partners and Macquarie University. LB thanks MIUR-
317 PRIN2017, project “TEOREM deciphering geological processes using Terrestrial and
318 Extraterrestrial ORE Minerals”, prot. 2017AK8C32. We acknowledge the scientific
319 and technical assistance of the Microscopy Australia at the Centre for Microscopy,
320 Characterization and Analysis, a facility funded by The University of Western
321 Australia, and State and Commonwealth Governments. This is contribution XXXX
322 from the ARC Centre of Excellence for Core to Crust Fluid Systems and XXXX from
323 the GEMOC Key Centre.

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430

431

432 **Figure Captions**

433

434 Fig. 1. Location maps showing the location of the Mt Carmel volcanic field. (a) Regional
435 setting; box shows Mt Carmel area; (b) Geological map indicating location of Cretaceous
436 volcanic centers.

437

438 Fig. 2. Combined phase map and backscattered electron (BSE) image of typical hibonite-
439 grossite-spinel assemblage from Mt Carmel, with drop-like inclusions of native vanadium
440 (V^0).

441

442 Fig. 3. (a) Transmitted-light photo (plane polars) of hibonite-grossite-spinel aggregate from
443 Mt Carmel, showing inclusions of V^0 in platy hibonite crystals (ca 5 mm across). Note
444 alternating layers with abundant or scarce inclusions, elongation of some inclusions parallel
445 to c-axis, and necking down of inclusions. (b) Hibonite from Sierra de Comechingones, with
446 vanadium rods normal to (0001) in (c) a grossite- $Ca_4Al_6F_2O_{12}$ matrix with skeletal F-rich

447 perovskite and dellagiustaite (V^{3+}/V^{2+} spinel), some rimming $V_{14}O_6$. Each figure is 1.7 mm
448 wide.

449

450 Fig. 4. BSE image of polished section of a hibonite crystal from Mt Carmel (blue prism shows
451 orientation derived from EBSD), with zonal distribution of V^0 inclusions.

452

453 Figure 5. (a) 3D- μ CT image of hibonite fragment (length 600 μ m) from Sierra de
454 Comechingones, showing necking-down of cylindrical rods of V^0 . (b) false-coloured image of
455 (a) showing that some V^0 rods (red) are partially empty (green) and some are rimmed by
456 dellagiustaite.

457

458 Fig. 6. (a) Reflected-light image of a polished rock chip from Sierra de Comechingones,
459 showing a $V_{14}O_6$ inclusion rimmed by dellagiustaite. Red arrow points to a very thin grain of
460 $V^{3+}_{1.78}Al_{0.21}V^{2+}_1O_4$; the composition of the dellagiustaite rim is $Al_{1.04}V^{3+}_{0.96}V^{2+}_{0.92}Mg_{0.08}O_4$. (b)
461 BSE image of V^0 inclusion with abundant fluid inclusions (Table 1b, analysis b). (c) BSE image
462 of V^0 inclusion (Table 1b; points 2 and 3 correspond to analyses c and d) with Cu
463 precipitates, rimmed by euhedral Ti-rich dellagiustaite.

464

465 Figure 7. (a) Transmitted-light and BSE images of a spiral dense branching structure of V^0 in
466 hibonite. (b) 3D- μ CT image of vanadium DBS in a hibonite crystal from Mt Carmel; each
467 cluster is ca 1 mm high and grows normal to the (0001) face of the hibonite. Red-orange
468 colours show V^0 , while green colours represent voids.

469

470 Figure 8. Al content vs unit cell dimension of inclusions of Al-bearing vanadium in grossite.

471

472 Figure 9. Raman spectra of hibonite from Sierra de Comechingones and Mt Carmel.

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484 **Table 1.** Mean major-element compositions of vanadium phases, Mt Carmel.

wt%	vanadium		V ₉ (Mn,Al)		V ₄ Al	
	EMP		EMP		EMP	
	n=5	stdev	n=5	stdev	n=7	stdev
Si	0.42	0.22	0.52	0.11	0.38	0.04
Zr			0.01	0.01		
Ti	0.01	0.02	0.15	0.11	0.20	0.03
Al	0.32	0.5	3.94	1.27	11.58	0.25
Cr	1.6*		0.71*		0.44	
V	95.2	3.23	79.83	6.86	83.37	0.76
Fe	0.03	0.01	0.01	0.01		
Mn	1.21	0.63	12.20	6.55	2.09	0.07
Ca	0.09	0.06	0.41	0.35	0.21	0.17
	98.9		97.78		98.28	

* EDS data

485

486 **Table 2.** Gas analyses of hibonite and corundum aggregates from Mt Carmel.

487

	sample	N [ppm]	C [ppm]	H [ppm]	S [ppm]
hibonite xl	982	60	1041	1572	73
hibonite aggregate	982	41	1165	2360	64.3

hibonite	1124B	122	1165	3293	113.1
	1124B-				
hibonite	dup.	43	1196	3925	228.7
corundum semi-transparent	1124	156	329	554	166
corundum-brecciated	1124	21	152	63	22.6
corundum with glass	1124	26	415	677	115.7
corundum	1124	226	347	667	222.5

488

489

490 **Table 3.** Observed Raman shifts in hibonite from Sierra di Comechingones (Argentina) and

491 Mt. Carmel (Israel)

492

Synthetic (Hofmeister et al. 2004)	Sierra de Comechingones	Sierra de Comechingones	Mt Carmel
<i>Orientation</i> ν (cm ⁻¹)	lath E-W ν (cm ⁻¹)	lath N-S ν (cm ⁻¹)	lath N-S ν (cm ⁻¹)
		85	84.3
			102.1
	130.8	130.8	127
193.7			
209.4p	205.4	205.4	208.9
250.5p			
274.0	268.6	275.8	273.9
331.9	318.9	323.7	
	358.3	345.7	349.6
399.4			
449.5		438	
457.8	457	453.8	451.9
489.1	485.4	485.4	485.4
529.5	522	531	534.7
564.5	576.3		
		593.6	
624.5			
640.4		643.7	643.3
683.8	685.6	681.6	678.9
	726.3		
741.0	744.9		
773.9	767		762.6
795.8		804.3	808
837.3		838.	837.4
872.9	870	871.7	875.7
910.0	905.6	909.9	902.4
		3314	
			3437
			3497
	3640	3623	
	3682	3681	

493 *bold = stronger bands
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495
496
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498

499 **Appendix.** Sampling and analytical methods

500 *Sampling*

501 The Shefa Yamim exploration project is aimed at the discovery of economically viable
502 placer deposits of gemstones (mainly sapphire, ruby, hibonite, moissanite) and other
503 commodities within the drainage basin of the Kishon River. Over >10 years, the project has
504 sampled the Cretaceous pyroclastic centers on Mt Carmel (Fig. SD1: Rakefet, Har Alon, Bat
505 Shelomo, Muhraka and Beit Oren complexes) and adjacent areas (Ein Ha-Shofet), and minor
506 and major drainages in the Yaz'rael Valley. Samples range in size from several kg to >1000
507 tonnes. All samples were run through a static grizzly screen to remove pieces larger than
508 100mm in diameter. Rock samples from the vents were coarsely crushed and then treated
509 in the same way as alluvial samples. The <100mm fraction was washed in a scrubber that
510 breaks up any clods. The <0.5mm component is suspended in the wash water and pumped
511 to settling ponds; fractions larger than 25 mm are used to backfill exploration pits. Samples
512 in the +8mm-16mm and +16mm-24mm size fractions are sorted by hand on a picking belt.
513 The +0.5 -8mm component of the sample is washed and classified into 5 fractions: 0.5-
514 0.7mm, 0.7-1mm, 1-2mm, 2-4mm, 4-6mm, 6-8mm. These fractions are transferred to a
515 pulsating jig plant for gravity separation. Samples in the 2mm-8mm size fractions are
516 visually inspected after the jiggling process and sorted in the recovery laboratory. The three
517 smallest size fractions are jiggled separately. The heavy concentrate in the center of the jig
518 pan is collected and dried; material on the outer part of the jig pan is discarded. The sorters
519 in the laboratory have demonstrated their efficiency in identifying and recovering a wide
520 range of mineral species, including garnet (pyrope), ilmenite, spinel, chrome-diopside,
521 diamond, moissanite, sapphire, ruby, Carmel Sapphire™, hibonite, rutile and zircon. The
522 material described here is dominated by samples from the Rakefet Magmatic Complex, and

523 the alluvial deposits of the Kishon River mid-reach, because these localities have been most
524 thoroughly bulk-sampled, providing abundant material.

525 The unprocessed heavy mineral concentrates of several samples were hand-picked
526 under a binocular microscope in the CCFS laboratories. Several rock samples also have been
527 processed by SelfFrag (electrostatic disaggregation) techniques at CCFS, sieved and hand-
528 picked after magnetic and heavy-liquid separation.

529

530

531 *Analytical Techniques*

532 *SEM, EMP*

533 Samples were mounted in epoxy blocks, polished and coated with carbon. A Zeiss
534 EVO MA15 scanning electron microscope (SEM) at the Geochemical Analysis Unit (GAU),
535 Macquarie University, Sydney, Australia was used to capture Backscattered Electron (BSE)
536 images and Energy Dispersive X-ray spectrometry (EDS) was used to map the elemental
537 composition of the samples, and to analyze phases of interest. An accelerating voltage of 15
538 keV and a beam current of 1 nA was used.

539 Major and minor elements were determined by electron microprobe (EMP) using a
540 CAMECA SX100 equipped with five wavelength-dispersive spectrometers at the Macquarie
541 University GeoAnalytical (MQGA; formerly GAU), Macquarie University, Sydney, Australia.
542 Analyses were performed using a focused beam (1-2 μm) with an accelerating voltage of 15
543 keV and a beam current of 20 nA. Standards were a suite of natural and synthetic minerals.
544 Peak counting varied between 10 – 20 s, and background was counted for 5 – 10 s on either
545 side of the peak. Oxygen and carbon were measured directly; matrix corrections were
546 carried out by the ZAF software.

547 Major and minor elements also were determined by electron microprobe (EMP)
548 using a JEOL 8530F electron microprobe equipped with five wavelength dispersive
549 spectrometers at the Centre for Microscopy Characterisation and Analysis (CMCA), The
550 University of Western Australia, Perth, Australia. Analyses were performed using a fully
551 focussed beam with an accelerating voltage of 15 keV, a beam current of 15 nA. Standards
552 were a suite of natural and synthetic minerals. On-peak counting times were 30 seconds for
553 all elements. X-ray intensity data was corrected for Time Dependent Intensity (TDI) loss (or
554 gain) using a self-calibrated correction for Si $k\alpha$, Ti $k\alpha$, Mn $k\alpha$, Na $k\alpha$, K $k\alpha$ and mean atomic

555 number (MAN) background corrections were used throughout (Donovan & Tingle, 1996;
556 Donovan et al., 2016). Unknown and standard intensities were corrected for deadtime.
557 Oxygen was measured, matrix corrections and ZAF correction were applied throughout.

558

559 *Transmission Electron Microscopy*

560 FIB foils from the region of interest were prepared for TEM using a dual-beam FIB
561 system (FEI Nova NanoLab 200). High Angle Annular Dark Field Scanning Transmission
562 Electron Microscopy (HAADF-STEM) imaging and element mapping were carried out using
563 an FEI Titan G2 80-200 TEM/STEM with ChemiSTEM Technology operating at 200 kV at the
564 Centre for Microscopy Characterisation and Analysis (CMCA), The University of Western
565 Australia, Perth, Australia. The element maps were obtained by energy dispersive X-ray
566 spectroscopy using the Super-X detector on the Titan with a probe size ~1 nm and a probe
567 current of ~0.25 nA. Total acquisition times of 20-30 minutes were used to obtain good
568 signal-to-noise ratios. Electron diffraction was carried out using a field limiting aperture that
569 selected an area approximately 400nm in diameter.

570

571 *3D- μ CT*

572 X-ray micro-computed tomography (μ CT) of Mt Carmel samples was done at the
573 CMCA, Perth. The grain was mounted on a finger of rigid PVC plastic with double-sided tape
574 and a separate tomogram was collected for each grain using a μ CT system (Versa 520 XRM,
575 Zeiss) running Scout and Scan software (v11.1.5707.17179, Zeiss). Samples were scanned at
576 50 kV and 77 μ A with the source detector positions set to -17.7 and 117 mm, respectively.
577 An LE1 source filter was applied to screen out low energy X-rays. A 0.4X objective lens was
578 used together with 2x camera binning to achieve an isotropic voxel resolution of 9 μ m. Each
579 scan comprised 2501 projections through 360 $^\circ$ with an exposure of 3s for each projection.
580 Raw data were reconstructed using XMReconstructor software (v11.1.5707.17179, Zeiss)
581 following a standard centre shift and beam hardening correction. The standard 0.7 kernel
582 size recon filter setting was also used. The visualization and analysis of data generated from
583 μ CT scans were performed using Avizo (v8.1.1, FEI) software.

584 Micro-CT images of hibonite from Sierra de Comechingones were obtained with a
585 Skyscan1172, Bruker micro CT, with a pixel size of 2.90 μ m and a X-ray source at 74 KV and
586 133 μ A and an Al-foil of 0.125 mm as a filter. Object to source distance was 46.2 mm and

587 camera to source distance was 276.56 mm. Raw data were reconstructed using CTVOx
 588 volume rendering software from Bluescientific

589

590 *Released-Gas Analysis*

591 H, C, N, and S contents were obtained using a vario EL cube elemental analyzer
 592 (Elementar, Langensfeld, Germany). Aliquots of finely powdered sample, weighing
 593 between 80 and 120 mg, were mixed with WO₃ flux (sample:flux ratio between 1/1.2 and
 594 1/1.5), packed in Sn-foils, hand pressed, and then ignited in an O₂–He gas atmosphere
 595 furnace at around 1150 °C. The produced gases were trapped and sequentially released into
 596 a set of chromatographic columns for the analysis of N (no trapping), then C, H, and S. Each
 597 sample was measured for 9 min, and released gases were sequentially analyzed with a
 598 thermal conductivity detector. Reference materials JP-1, BE-N, PM-S and BAM-U110, were
 599 repeatedly measured in the same analytical conditions and yielded average values (Table
 600 XX) well within error of accepted or published values as tabulated in the GeoReM database
 601 (<http://georem.mpch-mainz.gwdg.de/>).

602 **Table SD-1.** Reference materials for CHNS analyzer.

603

	H TCD	C TCD	N TCD	S TCD	S IR
JP-1 Peridotite massif (JGS)					
<i>n</i>	4	12	14	14	14
Average ($\mu\text{g g}^{-1}$) \pm SD	3195 \pm 170	763 \pm 82	91 \pm 23	27 \pm 14	26 \pm 7
<i>RSD</i> %	5	11	26	51	27
BHVO-2 Basalt, Hawaiian Volcanic Observatory (USGS)					
<i>n</i>	5	16	12	12	8
Average ($\mu\text{g g}^{-1}$)	181 \pm 45	316 \pm 60	219 \pm 34	191 \pm 41	179 \pm 42
<i>RSD</i> %	25	19	15	21	23
PM-S Microgabbros (SARM)					
<i>n</i>	25	26	24	26	22
Average ($\mu\text{g g}^{-1}$)	689 \pm 70	992 \pm 55	154 \pm 38	1128 \pm 51	1165 \pm 119
<i>RSD</i> %	10	6	15	5	10
BE-N Altered Basalts (SARM)					
<i>n</i>	14	20	17	21	8
Average ($\mu\text{g g}^{-1}$)	2771 \pm 534	2301 \pm 147	197 \pm 42	301 \pm 37	298 \pm 23
<i>RSD</i> %	19	6	21	12	8
BAM-U110					
<i>n</i>	13	18	18	17	–
Average ($\mu\text{g g}^{-1}$)	12,258 \pm 1758	72,340 \pm 2640	4237 \pm 165	9114 \pm 1082	–
<i>RSD</i> %	14	4	4	12	–

604 *n* denotes the number of measurements performed; average is the arithmetic means of the *n* values
 605 measured and SD is the related standard deviation; *RSD* % is relative standard deviation expressed in
 606 %; TCD: Thermal Conductivity Detector; IR: Infra Red detector.

607

608 *Raman Spectroscopy*

609 Raman spectra were collected with and Horiba LabRAM HR Evolution microscope with a
610 Sincerity OE detector, using 532-nm laser the Department of Earth Sciences, University of
611 Milan. Reflected and transmitted light optics were used to select analysis locations. Hibonite
612 from Sierra di Comechingones was studied from a polished thin section. Hibonite from Mt.
613 Carmel was studied using a raw fragment. The analysis was performed using a 100x
614 objective and 600 (500 nm) grating. Laser power was attenuated to 50% of the power. To
615 maximize the signal-to-noise ratio, two spectra were collected using an exposure time of 30
616 s, and then merged together at the end of the acquisition. Spectra were taken at two
617 polarizations at right angles to check for changes in intensity of the raman shifts. The system
618 was calibrated using the 520 cm^{-1} Raman line of metallic silicon. Spectra were processed
619 using LabSpec v.6 (Horiba[®]). Peak positions were obtained by fitting Gaussian peaks.
620 Observed peaks are reported in Table 3 and compared with Raman peaks reported by
621 Hofmeister et al. (2004) for synthetic hibonite. Two spectral ranges are reported in Figures
622 3a (50-1000 cm^{-1}) and 3b (3000-4150 cm^{-1}). Raw Raman spectra have been deposited as
623 supplementary material.

624

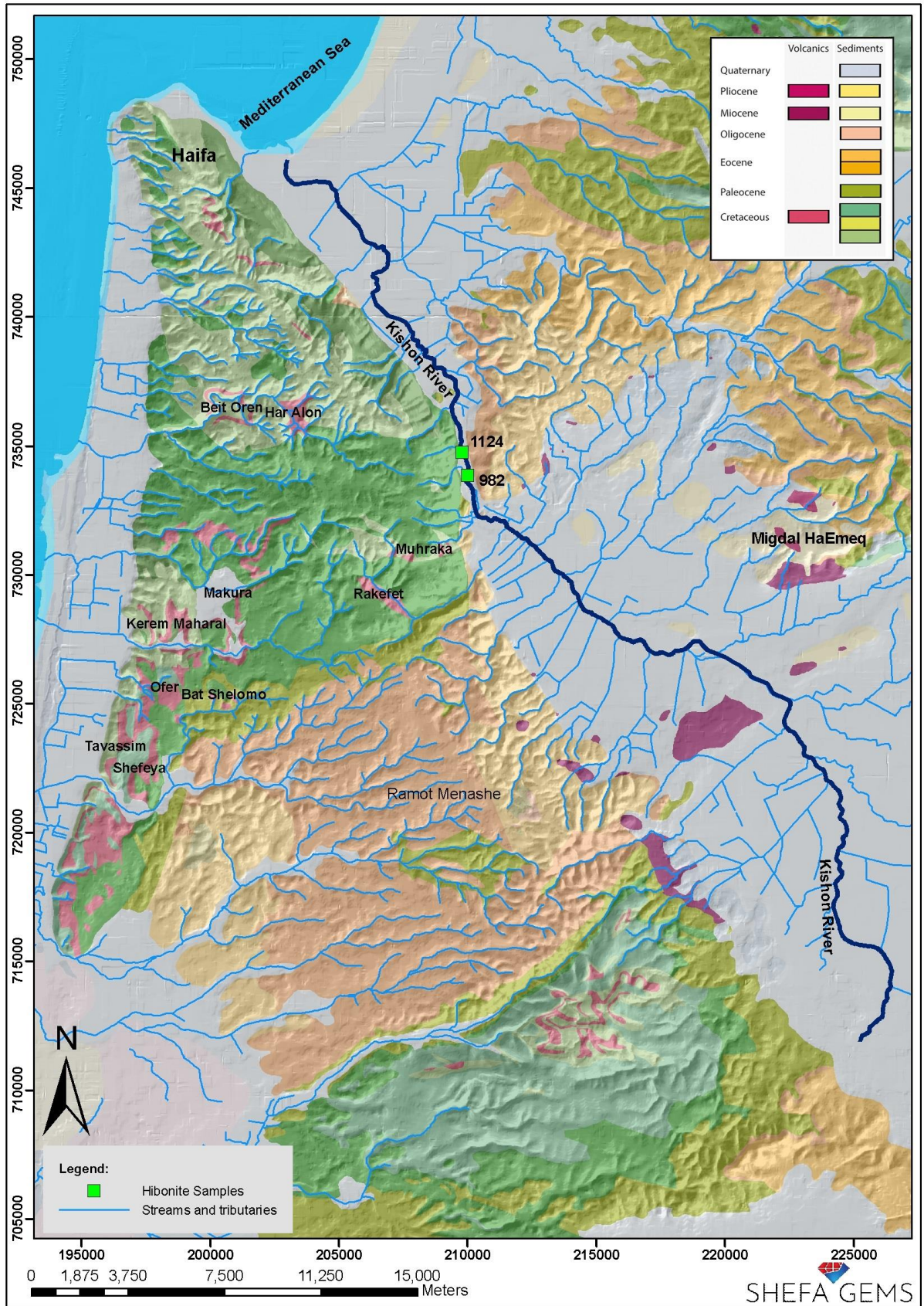
625 **Supplementary Figures**

626 Figure SD1. Animated 3D- μ CT image of 7b.

627

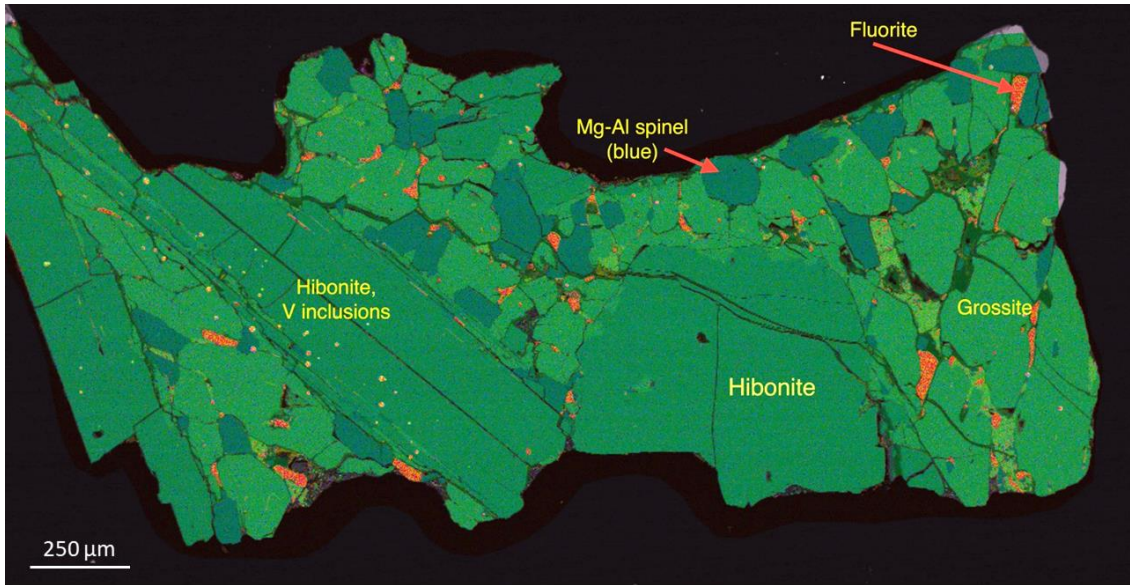
628 Figure SD2. 3D animation showing successive slices through the structures shown in Figure
629 7b (rotated 90°). Left, grey-scale image; V^0 appears as bright tones, and voids in black.
630 Right, image from Fig. 7b, to show correlation of voids with green colours. Note that some
631 large vanadium “drops” in the middle parts of the structures actually are comprised of voids
632 (green) surrounding an irregular fragment of V^0 (yellow). This is interpreted as the result of
633 exsolution of hydrogen from the crystallizing vanadium melt.

634



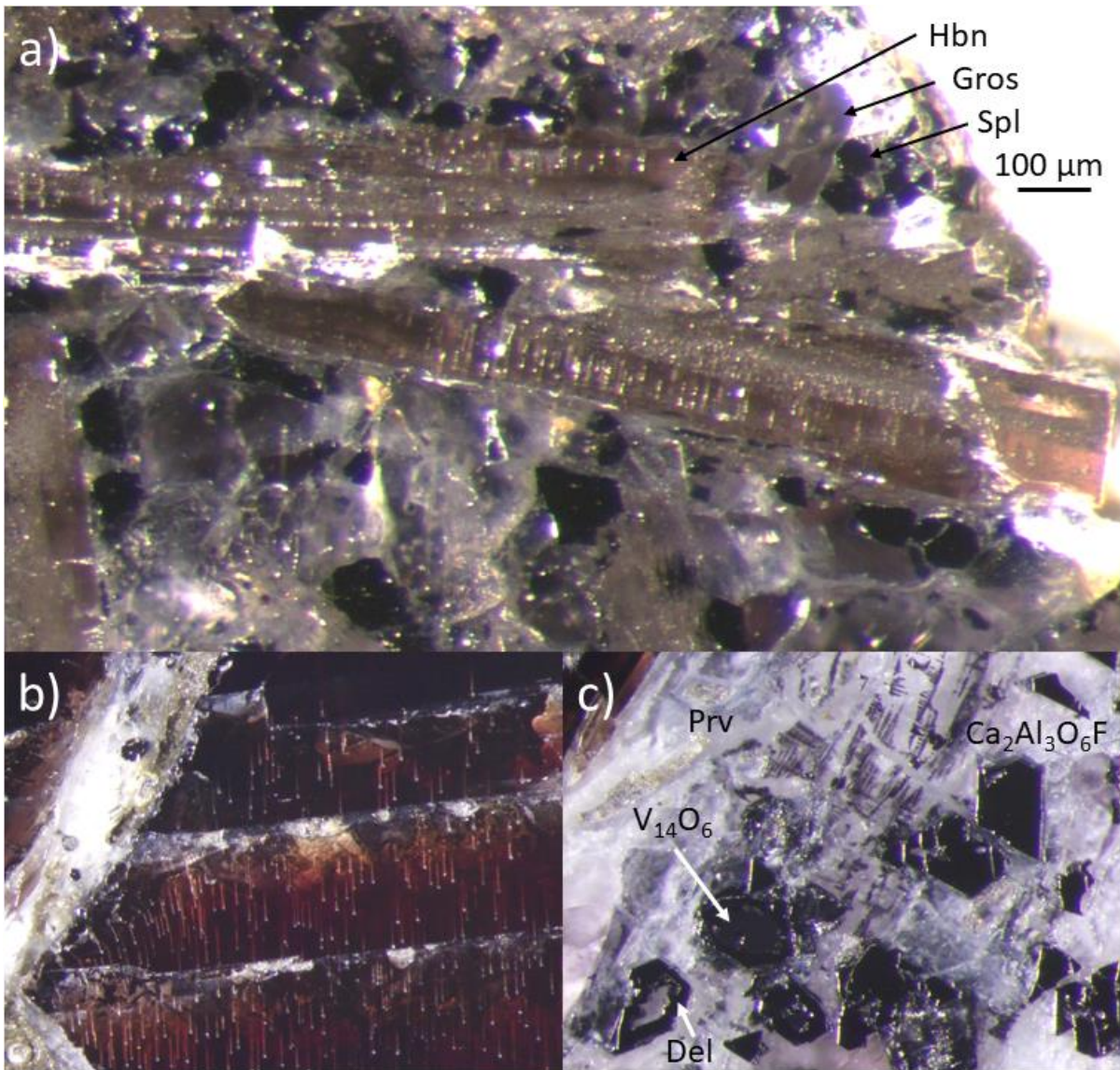
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Figure 1



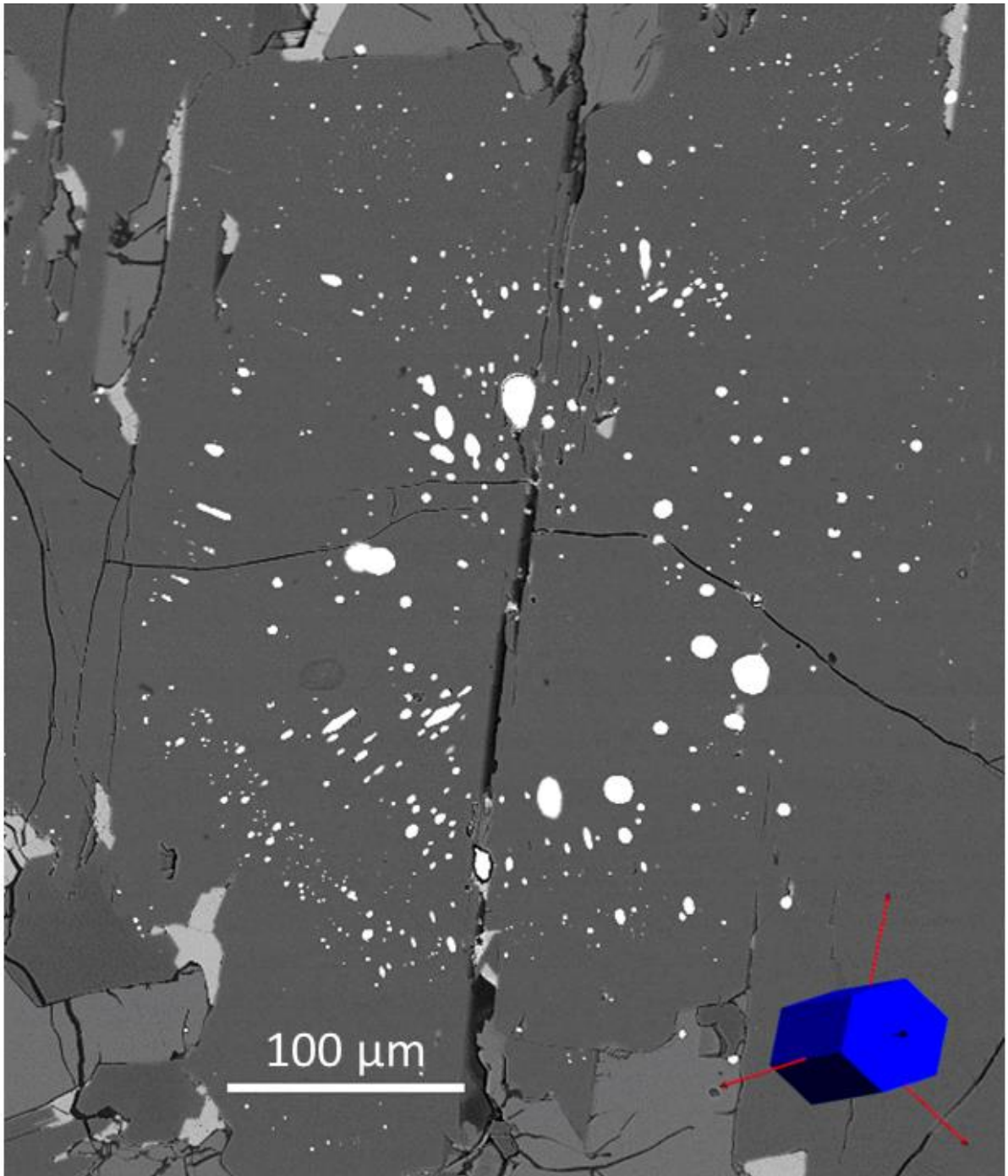
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Figure 2



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Figure 3



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Figure 4

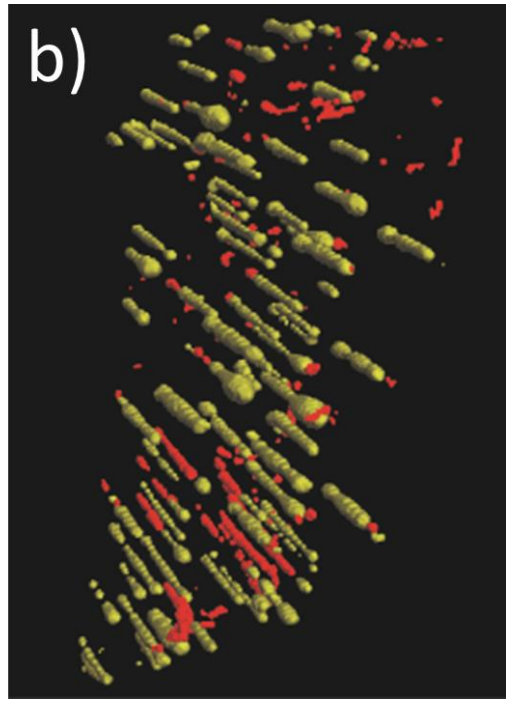
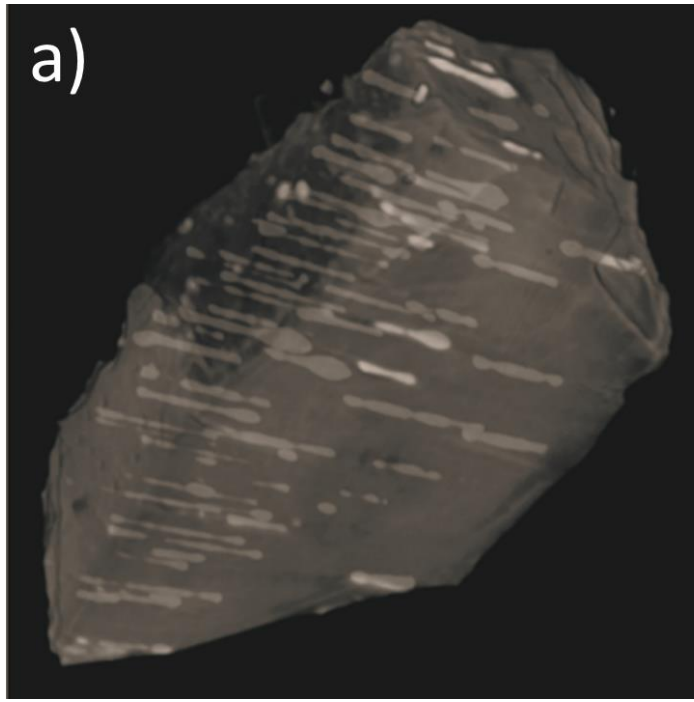


Figure 5

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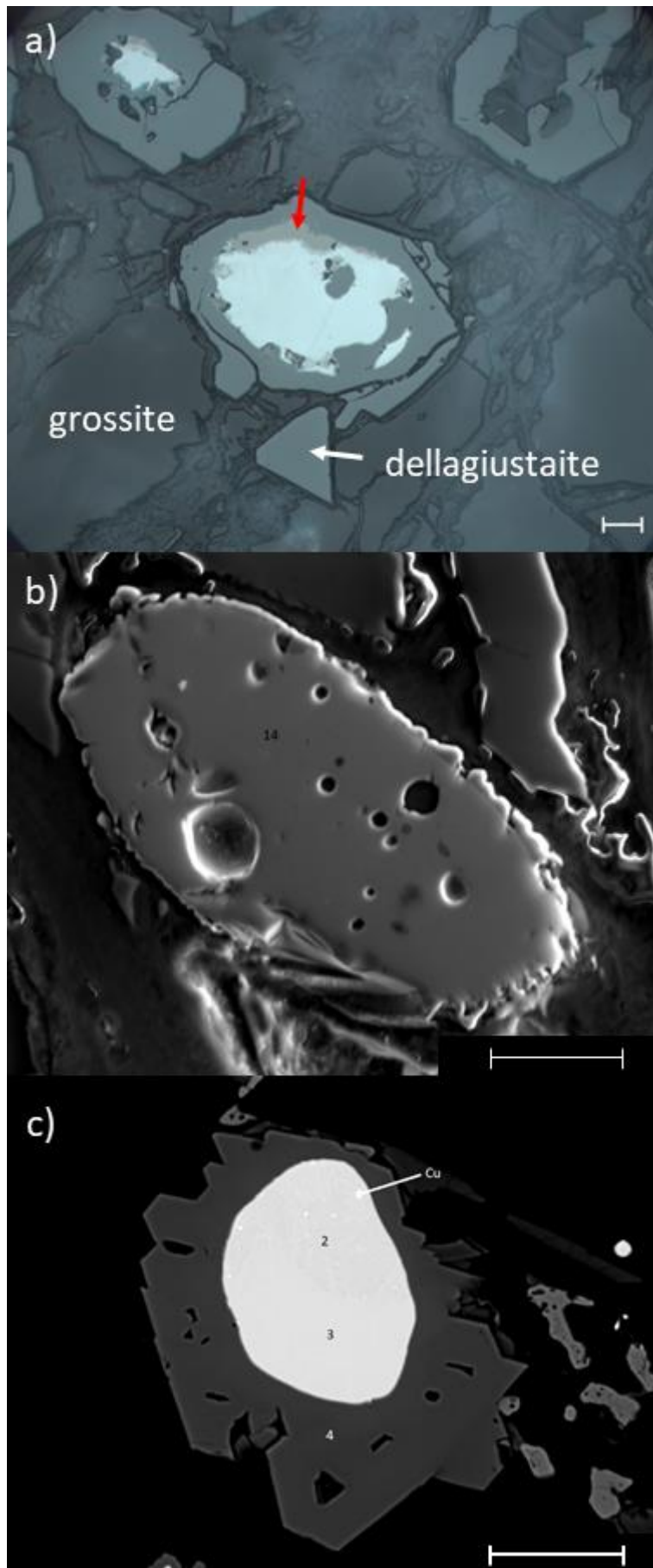


Figure 6

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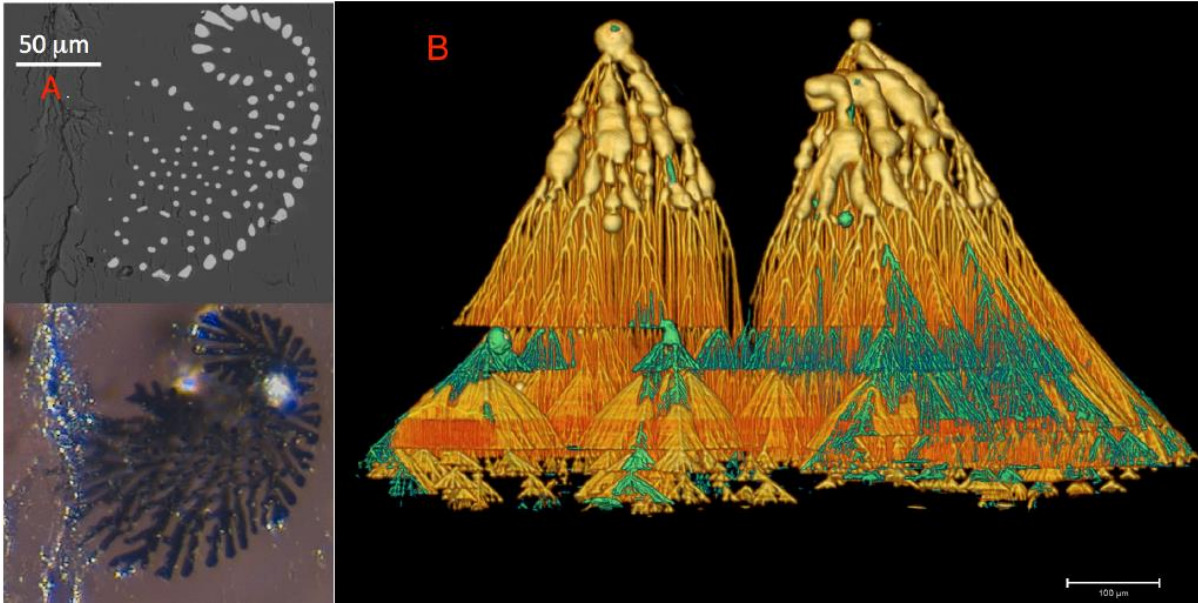


Figure 7

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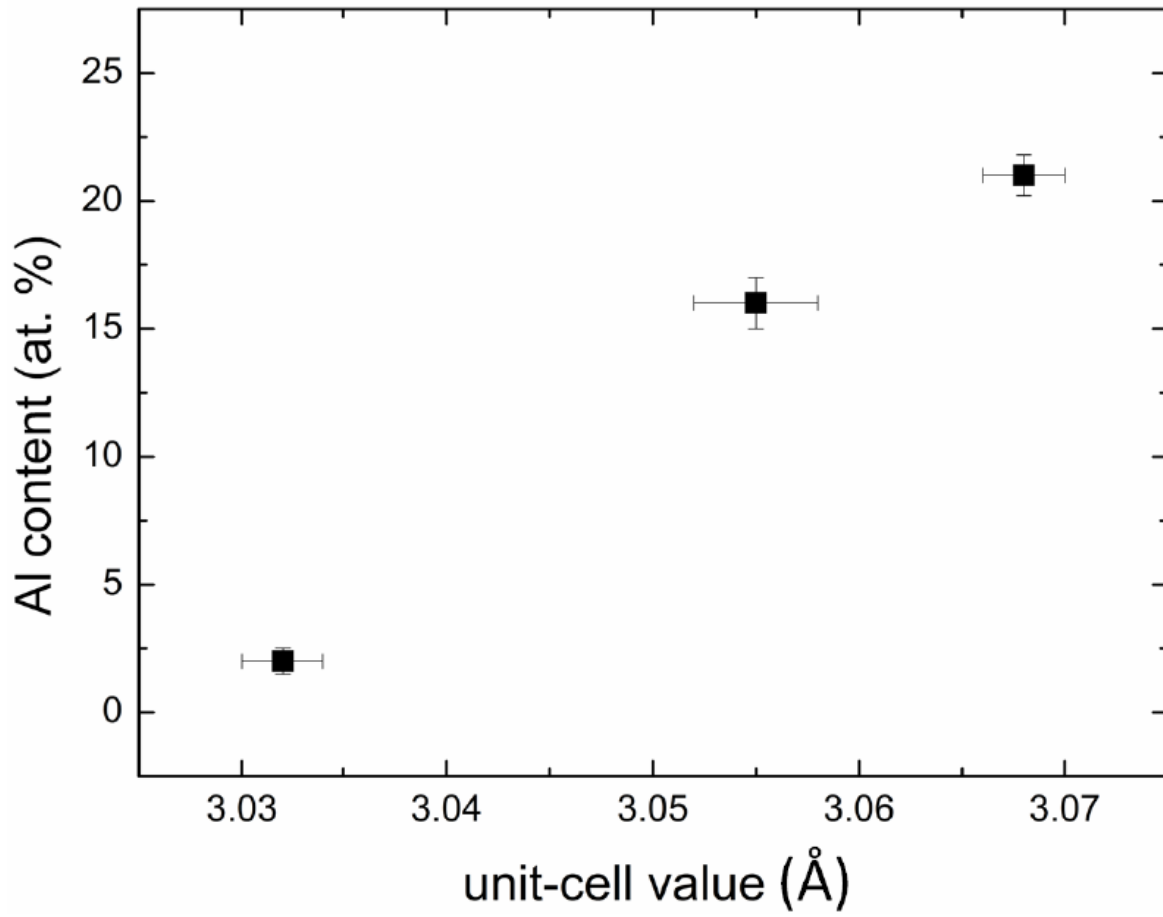
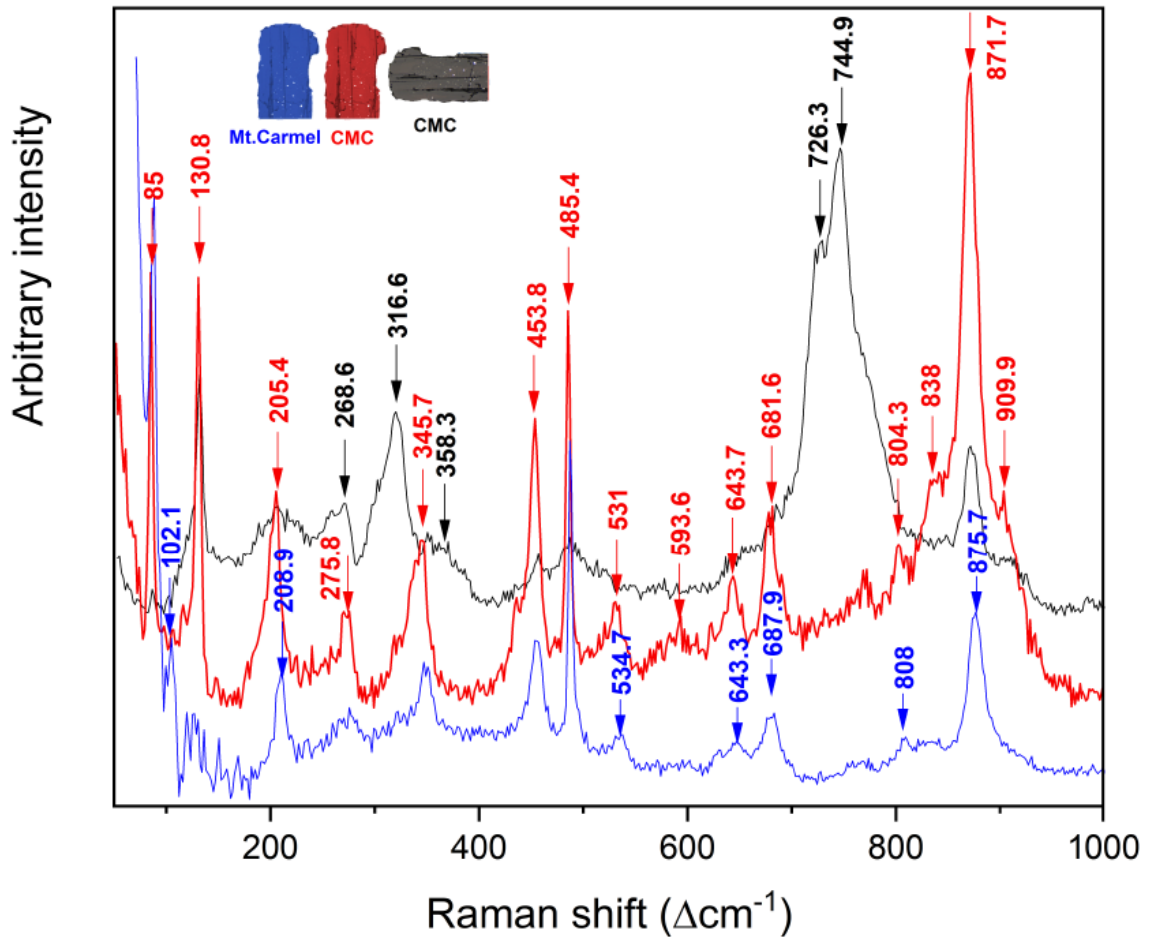


Figure 8

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Figure 9

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Appendix Sampling + Methods

Methods

Sampling

The Shefa Yamim exploration project is aimed at the discovery of economically viable placer deposits of gemstones (mainly sapphire, ruby, hibonite, moissanite) and other commodities within the drainage basin of the Kishon River. Over >10 years, the project has sampled the Cretaceous pyroclastic centers on Mt Carmel (Fig. SD1: Rakefet, Har Alon, Bat Shelomo, Muhraka and Beit Oren complexes) and adjacent areas (Ein Ha-Shofet), and minor and major drainages in the Yaz'rael Valley. Samples range in size from several kg to >1000 tonnes. All samples were run through a static grizzly screen to remove pieces larger than 100mm in diameter. Rock samples from the vents were coarsely crushed and then treated in the same way as alluvial samples. The <100mm fraction was washed in a scrubber that breaks up any clods. The <0.5mm component is suspended in the wash water and pumped to settling ponds; fractions larger than 25 mm are used to backfill exploration pits. Samples in the +8mm-16mm and +16mm-24mm size fractions are sorted by hand on a picking belt. The +0.5 -8mm component of the sample is washed and classified into 5 fractions: 0.5-0.7mm, 0.7-1mm, 1-2mm, 2-4mm, 4-6mm, 6-8mm. These fractions are transferred to a pulsating jig plant for gravity separation. Samples in the 2mm-8mm size fractions are visually inspected after the jiggling process and sorted in the recovery laboratory. The three smallest size fractions are jigged separately. The heavy concentrate in the center of the jig pan is collected and dried; material on the outer part of the jig pan is discarded. The sorters in the laboratory have demonstrated their efficiency in identifying and recovering a wide range of mineral species, including garnet (pyrope), ilmenite, spinel, chrome-diopside, diamond, moissanite, sapphire, ruby, Carmel Sapphire™, hibonite, rutile and zircon. The material described here is dominated by samples from the Rakefet Magmatic Complex, and the alluvial deposits of the Kishon River mid-reach, because these localities have been most thoroughly bulk-sampled, providing abundant material.

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693 *Analytical Techniques*

694 *SEM, EMP*

695 Samples were mounted in epoxy blocks, polished and coated with carbon. A Zeiss
696 EVO MA15 scanning electron microscope (SEM) at the Geochemical Analysis Unit (GAU),
697 Macquarie University, Sydney, Australia was used to capture Backscattered Electron (BSE)
698 images and Energy Dispersive X-ray spectrometry (EDS) was used to map the elemental
699 composition of the samples, and to analyze phases of interest. An accelerating voltage of 15
700 keV and a beam current of 1 nA was used.

701 Major and minor elements were determined by electron microprobe (EMP) using a
702 CAMECA SX100 equipped with five wavelength-dispersive spectrometers at the Macquarie
703 University GeoAnalytical (MQGA; formerly GAU), Macquarie University, Sydney, Australia.
704 Analyses were performed using a focused beam (1-2 μm) with an accelerating voltage of 15
705 keV and a beam current of 20 nA. Standards were a suite of natural and synthetic minerals.
706 Peak counting varied between 10 – 20 s, and background was counted for 5 – 10 s on either
707 side of the peak. Oxygen and carbon were measured directly; matrix corrections were
708 carried out by the ZAF software.

709 Major and minor elements also were determined by electron microprobe (EMP)
710 using a JEOL 8530F electron microprobe equipped with five wavelength dispersive
711 spectrometers at the Centre for Microscopy Characterisation and Analysis (CMCA), The
712 University of Western Australia, Perth, Australia. Analyses were performed using a fully
713 focussed beam with an accelerating voltage of 15 keV, a beam current of 15 nA. Standards
714 were a suite of natural and synthetic minerals. On-peak counting times were 30 seconds for
715 all elements. X-ray intensity data was corrected for Time Dependent Intensity (TDI) loss (or
716 gain) using a self-calibrated correction for Si $k\alpha$, Ti $k\alpha$, Mn $k\alpha$, Na $k\alpha$, K $k\alpha$ and mean atomic
717 number (MAN) background corrections were used throughout (Donovan & Tingle, 1996;
718 Donovan et al., 2016). Unknown and standard intensities were corrected for deadtime.
719 Oxygen was measured, matrix corrections and ZAF correction were applied throughout.

720

721 *Transmission Electron Microscopy*

722 FIB foils from the region of interest were prepared for TEM using a dual-beam FIB
723 system (FEI Nova NanoLab 200). High Angle Annular Dark Field Scanning Transmission
724 Electron Microscopy (HAADF-STEM) imaging and element mapping were carried out using
725 an FEI Titan G2 80-200 TEM/STEM with ChemiSTEM Technology operating at 200 kV at the
726 Centre for Microscopy Characterisation and Analysis (CMCA), The University of Western
727 Australia, Perth, Australia. The element maps were obtained by energy dispersive X-ray
728 spectroscopy using the Super-X detector on the Titan with a probe size ~ 1 nm and a probe
729 current of ~ 0.25 nA. Total acquisition times of 20-30 minutes were used to obtain good
730 signal-to-noise ratios. Electron diffraction was carried out using a field limiting aperture that
731 selected an area approximately 400nm in diameter.

732

733 *3D- μ CT* (add method for SdC material)

734 For X-ray micro-computed tomography (μ CT) at the CMCA, the grain was mounted
735 on a finger of rigid PVC plastic with double-sided tape and a separate tomogram was
736 collected for each grain using a μ CT system (Versa 520 XRM, Zeiss) running Scout and Scan
737 software (v11.1.5707.17179, Zeiss). Samples were scanned at 50 kV and 77 μ A with the
738 source detector positions set to -17.7 and 117 mm, respectively. An LE1 source filter was
739 applied to screen out low energy X-rays. A 0.4X objective lens was used together with 2x
740 camera binning to achieve an isotropic voxel resolution of 9 μ m. Each scan comprised 2501
741 projections through 360 $^\circ$ with an exposure of 3s for each projection. Raw data were
742 reconstructed using XMReconstructor software (v11.1.5707.17179, Zeiss) following a
743 standard centre shift and beam hardening correction. The standard 0.7 kernel size recon
744 filter setting was also used. The visualization and analysis of data generated from μ CT scans
745 were performed using Avizo (v8.1.1, FEI) software.

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747 *Released-Gas Analysis*

748 H, C, N, and S contents were obtained using a vario EL cube elemental analyzer
749 (Elementar, Langensfeld, Germany). Aliquots of finely powdered sample, weighing
750 between 80 and 120 mg, were mixed with WO₃ flux (sample:flux ratio between 1/1.2 and
751 1/1.5), packed in Sn-foils, hand pressed, and then ignited in an O₂-He gas atmosphere
752 furnace at around 1150 $^\circ$ C. The produced gases were trapped and sequentially released into
753 a set of chromatographic columns for the analysis of N (no trapping), then C, H, and S. Each

754 sample was measured for 9 min, and released gases were sequentially analyzed with a
 755 thermal conductivity detector. Reference materials JP-1, BE-N, PM-S and BAM-U110, were
 756 repeatedly measured in the same analytical conditions and yielded average values (Table
 757 XX) well within error of accepted or published values as tabulated in the GeoReM database
 758 (<http://georem.mpch-mainz.gwdg.de/>).

759 **Table SD-1.** Reference materials for CHNS analyzer.

760

	H TCD	C TCD	N TCD	S TCD	S IR
JP-1 Peridotite massif (JGS)					
<i>n</i>	4	12	14	14	14
Average ($\mu\text{g g}^{-1}$) \pm SD	3195 \pm 170	763 \pm 82	91 \pm 23	27 \pm 14	26 \pm 7
RSD %	5	11	26	51	27
BHVO-2 Basalt, Hawaiian Volcanic Observatory (USGS)					
<i>n</i>	5	16	12	12	8
Average ($\mu\text{g g}^{-1}$)	181 \pm 45	316 \pm 60	219 \pm 34	191 \pm 41	179 \pm 42
RSD %	25	19	15	21	23
PM-S Microgabbros (SARM)					
<i>n</i>	25	26	24	26	22
Average ($\mu\text{g g}^{-1}$)	689 \pm 70	992 \pm 55	154 \pm 38	1128 \pm 51	1165 \pm 119
RSD %	10	6	15	5	10
BE-N Altered Basalts (SARM)					
<i>n</i>	14	20	17	21	8
Average ($\mu\text{g g}^{-1}$)	2771 \pm 534	2301 \pm 147	197 \pm 42	301 \pm 37	298 \pm 23
RSD %	19	6	21	12	8
BAM-U110					
<i>n</i>	13	18	18	17	–
Average ($\mu\text{g g}^{-1}$)	12,258 \pm 1758	72,340 \pm 2640	4237 \pm 165	9114 \pm 1082	–
RSD %	14	4	4	12	–

761 *n* denotes the number of measurements performed; average is the arithmetic means of the *n* values
 762 measured and SD is the related standard deviation; RSD % is relative standard deviation expressed in
 763 %; TCD: Thermal Conductivity Detector; IR: Infra Red detector.

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765 *Raman Spectroscopy*

766 Raman spectra were collected with and Horiba LabRAM HR Evolution microscope with a
 767 Sincerity OE detector, using 532-nm laser the Department of Earth Sciences, University of
 768 Milan. Reflected and transmitted light optics were used to select analysis locations. Hibonite
 769 from Sierra di Comechingones was studied from a polished thin section. Hibonite from Mt.
 770 Carmel was studied using a raw fragment. The analysis was performed using a 100x
 771 objective and 600 (500 nm) grating. Laser power was attenuated to 50% of the power To
 772 maximize the signal-to-noise ratio, two spectra were collected using an exposure time of 30
 773 s, and then merged together at the end of the acquisition. Spectra were taken at two

774 polarizations at right angles to check for changes in intensity of the raman shifts. The system
775 was calibrated using the 520 cm^{-1} Raman line of metallic silicon. Spectra were processed
776 using LabSpec v.6 (Horiba[®]). Peak positions were obtained by fitting Gaussian peaks.
777 Observed peaks are reported in Table 3 and compared with Raman peaks reported by
778 Hofmeister et al. (2004) for synthetic hibonite. Two spectral ranges are reported in Figures
779 3a ($50\text{-}1000\text{ cm}^{-1}$) and 3b ($3000\text{-}4150\text{ cm}^{-1}$). Raw Raman spectra have been deposited as
780 supplementary material.

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