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Corresponding Author	Family Name	Borghini
	Particle	
	Given Name	Giulio
	Suffix	
	Division	Dipartimento di Scienze della Terra "Ardito Desio"
	Organization	Università degli Studi di Milano
	Address	Via Botticelli 23, 20133, Milan, Italy
	Phone	
	Fax	
	Email	giulio.borghini@unimi.it
	URL	
	ORCID	http://orcid.org/0000-0002-0246-7478
Author	Family Name	Fumagalli
	Particle	
	Given Name	P.
	Suffix	
	Division	Dipartimento di Scienze della Terra "Ardito Desio"
	Organization	Università degli Studi di Milano
	Address	Via Botticelli 23, 20133, Milan, Italy
	Phone	
	Fax	
	Email	
	URL	
	ORCID	
Author	Family Name	Rampone
	Particle	
	Given Name	E.
	Suffix	
	Division	
	Organization	DISTAV, Università di Genova
	Address	16132, Genoa, Italy
	Phone	
	Fax	
	Email	
	URL	

ORCID

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Abstract	<p>We performed partial melting experiments at 1 and 1.5 GPa, and 1180–1400 °C, to investigate the melting under mantle conditions of an olivine-websterite (GV10), which represents a natural proxy of secondary (or stage 2) pyroxenite. Its subsolidus mineralogy consists of clinopyroxene, orthopyroxene, olivine and spinel (+garnet at 1.5 GPa). Solidus temperature is located between 1180 and 1200 °C at 1 GPa, and between 1230 and 1250 °C at 1.5 GPa. Orthopyroxene (\pmgarnet), spinel and clinopyroxene are progressively consumed by melting reactions to produce olivine and melt. High coefficient of orthopyroxene in the melting reaction results in relatively high SiO₂ content of low melt fractions. After orthopyroxene exhaustion, melt composition is controlled by the composition of coexisting clinopyroxene. At increasing melt fraction, CaO content of melt increases, whereas Na₂O, Al₂O₃ and TiO₂ behave as incompatible elements. Low Na₂O contents reflect high partition coefficient of Na between clinopyroxene and melt ($D_{\text{Na}_2\text{O}}^{\text{cpx/liquid}}$). Melting of GV10 produces Quartz- to Hyperstene-normative basaltic melts that differ from peridotitic melts only in terms of lower Na₂O and higher CaO contents. We model the partial melting of mantle sources made of different mixing of secondary pyroxenite and fertile lherzolite in the context of adiabatic oceanic mantle upwelling. At low potential temperatures ($T_P < 1310$ °C), low-degree melt fractions from secondary pyroxenite react with surrounding peridotite producing orthopyroxene-rich reaction zones (or refertilized peridotite) and refractory clinopyroxene-rich residues. At higher T_P (1310–1430 °C), simultaneous melting of pyroxenite and peridotite produces mixed melts with major element compositions matching those of primitive MORBs. This reinforces the notion that secondary pyroxenite may be potential hidden components in MORB mantle source.</p>	
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2 **Partial melting of secondary pyroxenite at 1 and 1.5 GPa,**
3 **and its role in upwelling heterogeneous mantle**

4 **Giulio Borghini¹ · P. Fumagalli¹ · E. Rampone²**

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7 **Abstract** We performed partial melting experiments at 1
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9 under mantle conditions of an olivine-websterite (GV10),
10 which represents a natural proxy of secondary (or stage 2)
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23 TiO₂ behave as incompatible elements. Low Na₂O contents
24 reflect high partition coefficient of Na between clinopyroxene
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potential hidden components in MORB mantle source.

Keywords Pyroxenites · Experimental petrology · Partial
melting · Basalt composition · Mantle heterogeneity · Melt
productivity

Introduction

Several geochemical and petrological observations support
the idea that the source of oceanic basalts consists of heterogeneous
mantle including peridotite mixed to pyroxene-rich olivine-poor
mafic to ultramafic lithologies, such as eclogite and pyroxenite
(e.g., Allègre and Turcotte 1986; Hirschmann and Stolper 1996;
Kogiso et al. 2004a; Hofmann 2007; Sobolev et al. 2007; Lambart
et al. 2013 and references therein). In particular, the contribution
of pyroxene-rich lithologies has been invoked to explain the large
isotopic and trace elements variability of oceanic basalts,
although uncertainties still exist on how they affect their major
element chemistry. Pyroxenites can either represent recycled
crustal material or be introduced in the mantle by various
magmatic processes, thus resulting in pyroxene-rich mantle
components with a wide range of compositions

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A5 ✉ Giulio Borghini
A6 giulio.borghini@unimi.it

A7 ¹ Dipartimento di Scienze della Terra “Ardito Desio”,
A8 Università degli Studi di Milano, Via Botticelli 23,
A9 20133 Milan, Italy

A10 ² DISTAV, Università di Genova, 16132 Genoa, Italy



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60 (e.g., Kogiso et al. 2004a; Herzberg 2011; Lambart et al.
61 2013).

62 A large number of experimental studies has been dedi-
63 cated to partial melting of pyroxenites, and their phase rela-
64 tions and melting behavior, including solidus temperature
65 and melt productivity, have been defined for a wide range
66 of temperature, pressure and chemical compositions (e.g.,
67 Yaxley and Green 1998; Kogiso and Hirschmann 2001;
68 Hirschmann et al. 2003; Pertermann and Hirschmann
69 2003a; Kogiso et al. 2004a; Keshav et al. 2004; Médard
70 et al. 2006; Sobolev et al. 2007; Lambart et al. 2009a,
71 2013). Profiting of available experimental data, increas-
AQ1 72 ingly sophisticated geochemical models have investigated
73 the role of mixed pyroxenite–peridotite mantle sources
74 in the genesis of mid-ocean ridge basalts (MORBs) and
75 ocean island basalts (OIBs) (e.g., Hirschmann and Stolper
76 1996; Phipps Morgan 2001; Sobolev et al. 2005, 2007;
77 Stracke and Bourdon 2009; Herzberg 2006, 2011; Short-
78 tle and Maclennan 2011; Guerenko et al. 2013; Lambart
79 et al. 2012, 2016; Shorttle et al. 2014; Delavault et al. 2015;
80 Brown and Leshner 2016).

81 An important issue concerning the melting of pyroxen-
82 ite-bearing mantle sources is the role of interaction between
83 pyroxenite-derived melts and adjacent peridotite. Low-
84 solidus mantle lithologies, such as crustal-derived rocks
85 (eclogite or low-MgO pyroxenite), start to melt at higher
86 pressure than “dry” peridotites (e.g., Yasuda et al. 1994;
87 Pertermann and Hirschmann 2003a,b; Kogiso et al. 2004a;
88 Kogiso and Hirschmann 2006; Yaxley and Sobolev 2007;
89 Spandler et al. 2008). Their partial melts are expected to
90 react to some extent with the surrounding subsolidus
91 peridotite either modifying the composition of the rising
92 melts (Pilet et al. 2008; Mallik and Dasgupta 2012, 2013,
93 2014), or creating new hybrid rock-types, called second-
94 ary or stage 2 pyroxenites (e.g., Yaxley and Green 1998;
95 Sobolev et al. 2005, 2007; Herzberg 2006, 2011; Lambart
96 et al. 2012). Moreover, heterogeneous upwelling mantle is
97 subject to continuous events of partial melting and melt–
98 rock reaction that potentially generate a large range of new
99 lithological heterogeneities, including metasomatized peri-
100 dotites and variably residual pyroxenites (e.g., Yaxley and
101 Green 1998; Rosenthal et al. 2014). On this basis, recent
102 models of basalts petrogenesis have proposed that second-
103 ary pyroxenites could represent diffuse mantle components
104 in the MORB mantle sources (e.g., Sobolev et al. 2007;
105 Lambart et al. 2012). Remarkably, such hybrid components
106 are expected to produce MgO-rich partial melts that are
107 able to percolate through peridotite with minimum interac-
108 tions, thus preserving the pyroxenite signature of the origi-
109 nal melts (Lambart et al. 2009a, 2012). These pyroxenites
110 represent good candidates to transmit the trace element and
111 isotopic signature of deep mantle heterogeneities without
112 impacting the major element composition of basalts.

113 In spite of their inferred potential role in oceanic basalt
114 generation, the origin of such hybrid mantle components,
115 as well as the composition of their partial melts, is still
116 not well defined, mostly due to the scarceness of available
117 experimental data (Lambart et al. 2013). Lambart et al.
118 (2012) investigated the interaction between melts produced
119 by silica-deficient pyroxenites and peridotite, conclud-
120 ing that this process leads to the formation of secondary
121 high-MgO pyroxenites, compositionally similar to those
122 documented in the Ronda peridotite (Garrido and Bodi-
123 nier 1999). Sobolev et al. (2007) investigated the melting
124 behavior at 3.5 GPa of an olivine-free secondary pyroxenite
125 originated by eclogite-derived silica-rich melts reacted with
126 peridotite. However, experimental knowledge of the com-
127 positions of melts generated by secondary-type pyroxen-
128 ites in the pressure range of basalt generation at mid-ocean
129 ridges (i.e. 1–1.5 GPa; e.g., Presnall et al. 2002; Falloon
130 et al. 2008) is currently not available.

131 Natural examples of secondary pyroxenites, originated
132 by melt-peridotite reaction processes, have been docu-
133 mented in a number of studies on ultramafic massifs (i.e. AQ2
134 Garrido and Bodinier 1999; Bodinier et al. 2008; Gysi et al.
135 2012; Marchesi et al. 2013; Borghini et al. 2013, 2016;
136 Montanini and Tribuzio 2015). Specifically, pyroxenite
137 layers in External Liguride (EL) mantle sequences (North-
138 ern Apennines, Italy) have been recently interpreted as
139 the product of the reaction between deep infiltrated melts,
140 likely derived from a hybrid eclogite-peridotite source, and
141 the host peridotite (Borghini et al. 2016). Hence, they rep-
142 resent an excellent natural example of secondary (or stage
143 2) pyroxenite, which forms together with their host perido-
144 tite a potential source of E-MORBs (Borghini et al. 2013).

145 In this paper, we present the results of partial melting
146 experiments performed on a selected natural olivine-web-
147 sterite, whose origin by deep melt-rock reaction has been
148 well supported on the basis of microstructural and geo-
149 chemical observations (Borghini et al. 2013, 2016). The
150 major aims of our work are to derive the extent and com-
151 position of melts produced by moderate to high degrees of
152 melting of a secondary-type pyroxenite at 1 and 1.5 GPa
153 and discuss the contribution of such melts to the composi-
154 tion of oceanic basalts.

155 Experimental and analytical techniques

156 Choice and preparation of starting material

157 The investigated bulk composition is an olivine websterite
158 (sample GV10 of Borghini et al. 2016), representing a well-
159 preserved cm-width pyroxenite layer from the External
160 Liguride (EL) ophiolitic mantle sequences (Northern Apen-
161 nines, Italy). Although the EL peridotites are considered

162 exposed subcontinental lithospheric mantle sectors (e.g.,
163 Rampone et al. 1995), their MORB-like isotopic composi-
164 tion and chemical features suggest that they were formed
165 by lithospheric accretion of ordinary asthenospheric man-
166 tle (Rampone et al. 1995). Therefore, they represent good
167 proxies of an MORB-type fertile mantle source that is
168 modified by deep melt infiltration and melt–rock reaction
169 processes before it undergoes partial melting and basaltic
170 melt extraction (Borghini et al. 2013, 2016), as envisaged
171 by recent MORB petrogenesis models (e.g., Lambart et al.
172 2012).

173 Olivine-websterite GV10 belongs to a group of pyroxen-
174 ites having a geochemical signature inherited by a pri-
175 mary assemblage dominated by garnet and pyroxenes
176 (mostly clinopyroxene and minor orthopyroxene). Origin-
177 ally, they represented clinopyroxene-rich garnet pyrox-
178 enites with orthopyroxene-rich borders against the host
179 peridotite, indicative of an origin by silica-rich melt infil-
180 tration and crystallization within the peridotite at pres-
181 sure >1.5 GPa (Borghini et al. 2016). Microstructural and

182 geochemical observations on pyroxenite GV10 suggest
183 that it was originated by deep melt-rock reaction, likely
184 involving peridotite and a pyroxenite (or eclogite)-derived
185 melt, thus supporting its secondary origin (Borghini et al.
186 2016). Moreover, the emplacement of these pyroxenite lay-
187 ers introduced short-length scale isotopic heterogeneity
188 in the host peridotite, producing veined mantle domains
189 with enriched Nd-isotopic signature that could represent a
190 source of E-MORBs (Borghini et al. 2013). The selected
191 sample GV10 thus constitutes an excellent analog of a sec-
192 ondary pyroxenite component in a fertile MORB mantle
193 source.

194 Olivine-websterite GV10 has high X_{Mg} [molar Mg/
195 (Mg + Fe^{tot})] (0.82–0.88) coupled with high CaO, moderate
196 Al₂O₃ and low Na₂O contents (Fig. 1). Similar bulk composi-
197 tions have been documented in Group C pyroxenites from
198 Ronda (Spain; Garrido and Bodinier 1999) (Fig. 1). They
199 have been inferred to result from interaction between peri-
200 dotite and melts derived from garnet pyroxenites and have
201 been recently interpreted as secondary pyroxenite (Lambart

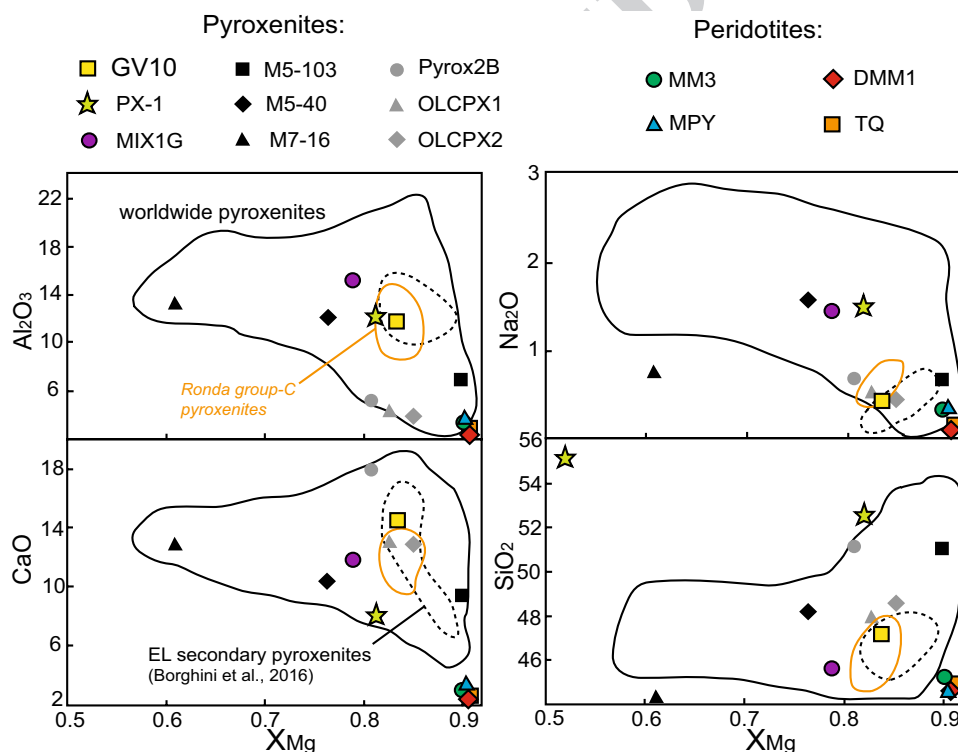


Fig. 1 X_{Mg} value vs. Al₂O₃, CaO, Na₂O and SiO₂ (wt%) diagrams showing the composition of the starting material used in this study (GV10), together with the composition of nominally “dry” pyroxenite and some peridotite bulks investigated in previous experimental studies at 1 and 1.5 GPa (see also Table 1). They are the silica-deficient pyroxenites M5-103, M5-40 and M7-16 studied by Lambart et al. (2009a); clinopyroxenites OLCPX1, OLCPX2 and Pyrox2B by Kogiso and Hirschmann (2001); garnet pyroxenite MIX1G (Hirschmann et al. 2003); lherzolite MM3 (Baker and Stolper 1994;

Falloon et al. 2008); depleted lherzolite DMM1 (Wasylenki et al. 2003); MORB-pyrolite, MPY, and Tinaquillo lherzolite, TQ (Robinson et al. 1998). Purple star is the composition of the model secondary (silica-excess) pyroxenites experimentally investigated at 3.5 GPa by Sobolev et al. (2007). Also reported are the compositional fields of worldwide pyroxenites (black line) and External Liguride (EL) pyroxenites (dotted line) from Borghini et al. (2016), and group-C pyroxenites from Ronda (orange line) from Garrido and Bodinier (1999)

et al. 2012). Compared to bulk compositions of nominally “dry” pyroxenites previously investigated at 1 and 1.5 GPa, GV10 has higher CaO and lower Na₂O contents with respect to pyroxenites investigated by Lambart et al. (2009a) (Table 1; Fig. 1), and higher Al₂O₃ and slightly lower SiO₂ than clinopyroxenites studied by Kogiso and Hirschmann (2001) at 1.0 GPa (Table 1; Fig. 1). In spite of similar X_{Mg}, GV10 significantly differs in composition from model secondary pyroxenite PX-1 studied by Sobolev et al. (2007), for much higher CaO and lower SiO₂ and Na₂O contents (Fig. 1). Such difference in bulk rock composition suggests that pyroxenite GV10 did not simply originate by eclogite melt–peridotite reaction (as for model secondary pyroxenite PX-1, Sobolev et al. 2005), but more presumably resulted by interaction between a melt produced by an eclogite–peridotite mixed source (e.g., Mallik and Dasgupta 2012) and a peridotite, as argued by Borghini et al. (2016).

Pyroxenite GV10 has been crushed in an agate mortar, and pulverized to 2–4 μm using an agate-micronizing mill. Two pellets obtained pressing about 0.5 g of rock powder each have been wrapped in platinum foils, previously preconditioned for Fe, and mounted at the end of a shaft in a vertical furnace. A first Fe preconditioning of platinum foil was performed at 1500 °C for about 12 h at FMQ fO₂ conditions, by covering it with GV10 rock powder. Glass GV10 was obtained by heating the pellets up to 1500 °C at FMQ fO₂ conditions through a night and quenched in dry ice. Electron microprobe analysis on the GV10 glass

indicated its homogeneous composition and the lack of any residual mineral phase. Fe loss was lower than 4% with respect to the starting rock powder composition (Table 1).

Experimental and analytical techniques

Experiments were conducted at 1.0 and 1.5 GPa, and temperatures from 1180 to 1400 °C, at the Laboratorio di Petrologia Sperimentale, Dipartimento di Scienze della Terra, University of Milano. Experiments at 1 GPa were carried out in a single-stage piston cylinder and experiments at 1.5 GPa in end-loaded piston cylinder using MgO-Pyrex-Salt assemblies. Run lasted from 50 to 240 h (Table 2). As we did not perform time series experiments, run durations of this work have been chosen following previous experiments on pyroxenite melting (i.e. Kogiso and Hirschmann 2001; Pertermann and Hirschmann 2003a, b). Accordingly, we performed super-solidus experiments with run durations ranging from 50 to 155 h. Subsolidus experiments lasted up to 240 h, in order to approach equilibrium of the run products and favor the development of textures with grain size suitable for microprobe analyses (Table 2).

Approximately, 20 mg of starting material was loaded into a graphite inner capsule (outer diameter 2.8 mm), and then welded into an outer Pt capsule (outer diameter 3.0 mm, length about 7–8 mm). Graphite is used to isolate the sample from the Pt capsule and avoid Fe-loss (e.g.,

Table 1 Starting compositions experimentally investigated at 1 and 1.5 GPa

	GV10	GV10	Lambart et al. (2009)			Kogiso and Hirschmann (2001)			B&S (1994)
	rock an.(*)	glass	M5-103	M5-40	M7-16	Pyrox2B	OLCPX1	OLCPX2	MM3
SiO ₂	47.54	47.57 (12)	51.64	48.53	43.58	51.56	47.92	48.73	45.50
TiO ₂	0.45	0.45 (6)	0.09	0.52	0.75	0.54	0.47	0.39	0.11
Al ₂ O ₃	10.27	10.39 (6)	7.17	12.37	13.73	5.58	4.49	4.13	3.98
Cr ₂ O ₃	0.16	0.15 (4)	0.60	0.12	0.07	0.00	0.00	0.00	0.68
FeO	7.23	6.99 (11)	4.97	9.02	14.51	7.01	9.26	7.93	7.18
MnO	0.17	0.18 (6)	0.10	0.20	0.30	0.10	0.13	0.12	0.13
MgO	19.15	19.23 (14)	24.57	16.64	12.52	16.19	23.38	24.42	38.30
CaO	14.43	14.49 (9)	10.13	10.89	13.77	18.18	13.72	13.71	3.57
Na ₂ O	0.50	0.48 (2)	0.71	1.65	0.75	0.84	0.63	0.57	0.31
K ₂ O	0.10	0.08 (2)	0.02	0.06	0.03	0.00	0.00	0.00	0.23
Total	100.00	100.00	100.00	100.00	100.00	100.00	100.00	100.00	100.00
X _{Mg}	0.82	0.83	0.90	0.77	0.61	0.81	0.82	0.85	0.09
CIPW ^b	0/27/4/37/31	0/27/5/37/30	2/32/0/28/36	0/28/8/22/41	0/18/31/27/22	0/12/5/63/19	0/37/0/47/15	0/35/4/47/14	0/63/16/7/13

All compositions are normalized to a sum of 100%. In the second column, in parentheses we report the errors (1σ) given in terms of least unit Cited; e.g., 47.57 (12) and 0.45 (6) represent 47.57 ± 0.12 and 0.45 ± 0.06, respectively. X_{Mg} = Mg/(Mg + Fe^{tot})

B&S (1994) refers to Baker and Stolper (1994)

^a Major element analyses by lithium metaborate/tetraborate fusion ICP on powder samples GV10 (Borghini et al. 2016)

^b Normative composition in terms of Nepheline/Olivine/Hypersthene/Diopside/Feldspars

Table 2 Summary of experimental conditions and run products

Run	T (°C)	Time (h)	Phase assemblages and modes (wt%)									
			Olivine	cpx	opx	Spinel	Garnet	Glass	R2 ^a	Fe loss (%)	Kd1	Kd2
Experiments at 1 GPa												
GV10-83-4	1180	240	5.2 (29)	74.4 (27)	16.1 (51)	4.5 (13)	0	0	0.1123	–	–	–
GV10-83-1	1200	120	6.6 (26)	78.9 (24)	7.3 (43)	5.4 (8)	0	1.8 (29)	0.1045	0.52	0.22	0.33
GV10-83-2	1250	96	10.5 (17)	72.2 (64)	0	4.5 (9)	0	12.6 (56)	0.1326	1.12	0.31	0.33
GV10-83-14	1280	83	12.1 (25)	45.4 (28)	0	2.2 (12)	0	40.2 (31)	0.0998	0.68	0.33	0.32
GV10-83-6	1300	95	13.2 (11)	29.7 (44)	0	0	0	56.7 (46)	0.1604	0.62	0.32	0.31
GV10-83-8	1330	146	8.6 (57)	0	0	0	0	91.3 (45)	0.1142	1.23	0.32	0.31
Experiments at 1.5 GPa												
GV10-83-16	1230	124	6.3 (21)	78 (41)	6.7 (53)	3.1 (19)	5.2 (46)	0	0.3421	–	–	–
GV10-83-9	1250	155	10.1 (35)	76.9 (36)	1.5 (38)	4.5 (28)	0	7.1 (38)	0.2046	0.05	0.31	0.34
GV10-83-10	1280	72	10.3 (35)	73.4 (58)	0	3.6 (11)	0	12.6 (53)	0.1899	0.24	0.31	0.33
GV10-83-13	1300	70	10.5 (32)	67.4 (55)	0	2.4 (13)	0	19.4 (44)	0.4765	1.89	0.30	0.33
GV10-83-18	1330	76	10.5 (17)	56.7 (33)	0	1 (9)	0	31.1 (37)	0.5624	2.33	0.30	0.32
GV10-83-15	1350	66	10.9 (21)	44.7 (57)	0	1.3 (11)	0	42.3 (62)	0.5471	0.82	0.33	0.32
GV10-83-21	1370	52	9.9 (11)	37.3 (41)	0	0	0	52.1 (35)	0.6333	0.56	0.33	0.32
GV10-83-22	1400	50	11.6 (8)	0	0	0	0	88.4 (22)	0.1944	0.77	0.34	0.32

Kd is the exchange coefficient of Fe²⁺ and Mg between olivine and liquid measured (1) and calculated (2) following Toplis (2005)

cpx clinopyroxene, *opx* orthopyroxene

^a Sum of squared residuals in mass balance calculations. Numbers in parentheses are propagated errors from Monte Carlo simulations, given in terms of the least unit cited; e.g., 5.2 (29) and 74.4 (27) represent 5.2 ± 2.9 and 74.4 ± 2.7 , respectively

255 Kinzler 1997; Walter 1998). Furthermore, the graphite–Pt
 256 assembly keeps the oxygen fugacity below the graphite–
 257 C–O vapor buffer (e.g., Ulmer and Luth 1991; Médard
 258 et al. 2008). We added a layer of vitreous carbon spheres
 259 (80–125 μm diameter) at the top of the capsule to enhance
 260 the storage and segregation of melt (e.g., Pickering–Witter
 261 and Johnson 2000; Wasylenki et al. 2003; Médard et al.
 262 2006). Carbon spheres may also absorb water if present in
 263 the charge (Robinson et al. 1998). The mass of the vitreous
 264 spheres layer never exceeded 15% of the starting glass. To
 265 maintain anhydrous conditions, the platinum–graphite cap-
 266 sule loaded with the starting material was dried overnight
 267 in an oven at 250 °C before being rapidly welded shut. The
 268 thermocouple tip was separated from the platinum capsule
 269 by a 0.5-mm thick hard corundum disc. Assemblies were
 270 kept in oven at about 200 °C for several hours before run-
 271 ning the experiments. Temperature was measured by K-type
 272 and S-type thermocouples and is considered to be accurate
 273 to ±5 °C. An initial pressure of 0.25 GPa was applied, then
 274 the sample was first heated to 400 °C for 10 min in order
 275 to soften the Pyrex; pressure was raised to the experimental
 276 value before reaching the desired temperature. Runs were
 277 terminated by turning off the power. Capsules were enclosed
 278 in epoxy, sectioned lengthwise, polished and carbon-coated.

279 Run products were inspected by back-scattered electron
 280 images (BSE) and microanalyses were performed using

a JEOL JXA 8200 Superprobe equipped with five WDS-
 wavelength-dispersive spectrometers and one energy disper-
 sive spectrometer (EDS) at the Dipartimento di Scienze della
 Terra, University of Milano. Both images and X-ray element
 maps were extremely useful in textural examination of the
 experimental charges. Analyses on mineral phases were per-
 formed using 1 μm beam size and beam conditions of 15 kV
 and 5 nA. Counting time was 30 s for peak and 10 s for back-
 ground. Whenever possible, in order to prevent alkali loss, we
 analyzed glass using a beam size of 5 μm, or 2–3 μm in thin
 interstitial films in experiments with very low melt fractions.

Results

Approach to equilibrium

The approach to equilibrium in experiments of this study is
 assessed by a number of observations, first of all the growth
 of compositionally homogeneous chemically unzoned min-
 erals, as revealed by X-ray element maps (Supplementary
 Figure 1). A solid proof of approach to equilibrium is the
 systematic and consistent variations in mineral chemistry
 at different pressure and temperature conditions, as well
 as coherent element partitioning (see below). Phase abun-
 dances have been calculated by mass balance between the

303 composition of coexisting phases and the bulk composition,
 304 using a weighted least squares method (Table 2). Residual
 305 sums of squares ranging from 0.09 and 0.63 are perfectly
 306 comparable to other experimental studies (i.e. Kogiso and
 307 Hirschmann 2001; Pertermann and Hirschmann 2003a, b;
 308 Lambart et al. 2009a). Moreover, results of mass balance
 309 calculations indicate that Fe loss, possibly due to fractures
 310 in graphite inner capsule to the Pt outer capsule, was negli-
 311 gible in these experiments (Table 2).

312 As a further test, we applied the method proposed by
 313 Toplis (2005). The values of distribution coefficient K_D
 314 between olivine and liquid measured in our experiments
 315 differ from those computed using thermodynamic calcula-
 316 tion (Toplis 2005) within the range of ± 0.02 (Table 2),
 317 thus supporting the attainment of equilibrium. As only
 318 exception, run GV10-83-1 at very low degree of melting
 319 ($F = 1.8\%$, see below) yielded an olivine/liquid $K_{D_{Fe-Mg}}$
 320 measured value much lower than the calculated one. On
 321 this experiment, we did not adopt specific technique to
 322 account for the effect of quench crystal growth on the very
 323 small spots of melt, as applied in previous dedicated stud-
 324 ies (i.e. Hirose and Kushiro 1993; Wasylenki et al. 2003;
 325 Laporte et al. 2004; Dasgupta and Hirschmann 2007; Fal-
 326 loon et al. 2008). Accordingly, the composition of glasses
 327 produced at very low melting degrees ($F < 5\%$) in our
 328 experiments must be considered with caution.

329 **Phase assemblages and textures**

330 Run products are summarized in Table 2, together with
 331 experimental details.

332 Crystal size ranges from 1 to 10 μm in subsolidus experi-
 333 ments and increases with increasing temperature, around
 334 5–15 μm in experiments with low melt fractions (<20
 335 wt%), and up to 80 μm in near-liquidus runs.

336 At 1 GPa and 1180 $^{\circ}\text{C}$, the stable mineral assemblage is
 337 made of prismatic grains of clinopyroxene, and orthopyrox-
 338 ene (5–10 μm in size), minor rounded olivine and tiny spin-
 339 el (mostly 2–5 μm in size). At 1200 $^{\circ}\text{C}$, few tiny interstitial
 340 patches of glass coexist with clinopyroxene, orthopyroxene,
 341 olivine and spinel. At 1250 $^{\circ}\text{C}$, orthopyroxene is absent; resid-
 342 ual clinopyroxene and olivine reach 15 μm in size (Fig. 2a),
 343 and interstitial glass blebs (mostly confined at the top of the
 344 capsule) are larger as compared to lower temperature experi-
 345 ment. At 1280 $^{\circ}\text{C}$, glass is still in equilibrium with clinopy-
 346 roxene, olivine and spinel and it occurs as large pools (about
 347 120–150 μm) at the rim of the capsule. At 1300 $^{\circ}\text{C}$, spinel is
 348 absent and a homogeneous glass coexists with coarse residual
 349 clinopyroxene and olivine (10–30 μm). At 1330 $^{\circ}\text{C}$, olivine
 350 is the only residual phase and it consists of coarse polygonal
 351 grains confined at the bottom portion of the capsule (Fig. 2b).

352 In experiments at 1.5 GPa, a subsolidus fine-grained
 353 (>10 μm) mineral assemblage made of clinopyroxene,

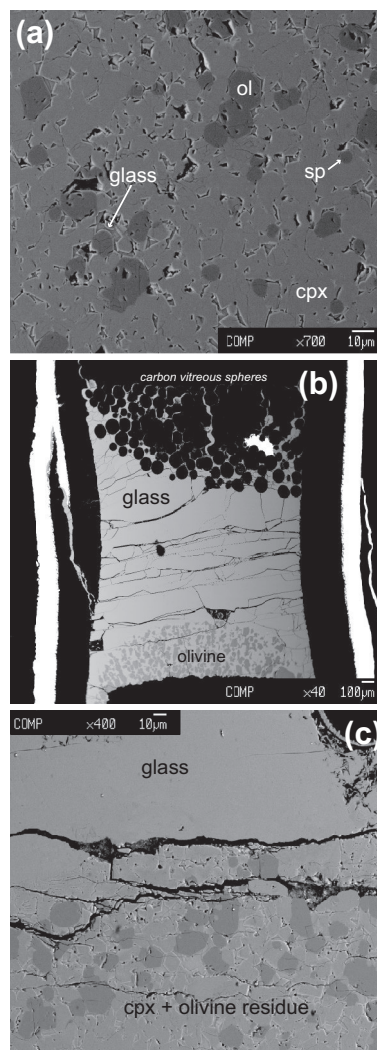


Fig. 2 Representative textures in melting experiments on pyroxenite GV10. **a** Back-scattered electron (BSE) image of melting experiment GV10-83-2 (1 GPa; 1250 $^{\circ}\text{C}$): glass is mostly confined at the top of the capsule and occurs as interstitial blebs between residual clinopyroxene and olivine (around 15 μm in size) and smaller spinel (up to 5 μm). **b** BSE image of run GV10-83-8 (1 GPa; 1330 $^{\circ}\text{C}$): residual olivine as coarse polygonal grains floating in the glass at the bottom of the capsule. **c** BSE image of experiment GV10-83-21 (1.5 GPa; 1370 $^{\circ}\text{C}$): coarse clinopyroxene and olivine coexist with glass either contained in the trap layer or as a pool towards the top of capsule

354 olivine, orthopyroxene, spinel and garnet is stable at
 355 1230 $^{\circ}\text{C}$. Garnet is the first phase to disappear at 1250 $^{\circ}\text{C}$,
 356 and orthopyroxene is no more present at 1280 $^{\circ}\text{C}$. In runs
 357 at temperatures from 1280 to 1350 $^{\circ}\text{C}$, glass proportion
 358 increases progressively with temperature and the residual
 359 phases are clinopyroxene, olivine and spinel (Table 2).
 360 Grain size is almost constant with clinopyroxene and oli-
 361 vine ranging 10–30 μm and spinel always smaller than
 362 5 μm . At 1370 $^{\circ}\text{C}$, spinel is entirely consumed and coarse
 363 clinopyroxene and olivine coexist with glass; in such

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364 experiments with high degrees of melting, glass is either
365 contained in the trap layer or as a pool towards the top of
366 the capsule (Fig. 2c). At 1400 °C, rounded large olivines
367 float in a homogeneous glass, thus indicating that olivine is
368 the liquidus phase also at 1.5 GPa.

369 Liquid compositions as a function of pressure 370 and temperature

AQ3 371 The glass compositions are reported in Table 3 and in
372 Fig. 3, where major and minor element contents are plot-
373 ted against melt fraction. Melt compositions are basaltic
374 and vary significantly as a function of melt fraction (F) and
375 only slightly with pressure.

376 At increasing melt fraction, MgO and CaO abundances
377 progressively increase and Al₂O₃, Na₂O and TiO₂ contents
378 decrease at both 1 and 1.5 GPa (Fig. 3). Na₂O and TiO₂
379 contents are lower than about 2 wt% in all experiments, and
380 they decrease at increasing melt fraction at both pressures
381 (Fig. 3). K₂O behaves as highly incompatible element,
382 reaching about 0.4 wt% in experiments with the lowest
383 melt fractions at 1 and 1.5 GPa, and drastically decreasing
384 at increasing melting degree (Table 3).

385 The highest SiO₂ contents are observed in the experi-
386 ments with the lowest melt fraction ($F < 8\%$); SiO₂ abun-
387 dance significantly decreases as long as orthopyroxene is a
388 melting phase and it remains almost constant once orthopy-
389 roxene is exhausted, i.e. at $F = 2\text{--}12\%$ at 1 GPa and
390 $F > 7\%$ at 1.5 GPa. The FeO compositional trend is rather
391 flat at 1 GPa, ranging from 7.1 to 8.3 wt%; at 1.5 GPa it
392 has a convex shape, increasing up to 9.7 wt% with orthopy-
393 roxene consumption (corresponding to 20% melt fraction),
394 and slightly decreasing after orthopyroxene disappearance
395 at higher melt fraction (Fig. 3). Cr₂O₃ melt contents are
396 very low as long as spinel is present as residual mineral,
397 and increases at high melt fraction ($F > 60\%$) after spinel
398 exhaustion, at both 1 and 1.5 GPa (Fig. 3). The X_{Mg} value
399 [$X_{\text{Mg}} = \text{Mg}/(\text{Mg} + \text{Fe}_{\text{tot}})$] of melts increases with the melt
400 fraction, from 0.53 to 0.79 in runs at 1 GPa, and from 0.63
401 to 0.80 at 1.5 GPa (Table 3). It is positively correlated with
402 X_{Mg} value of olivine and clinopyroxene (Fig. 4), supporting
403 approach to equilibrium.

404 Differences in melt composition as a function of pres-
405 sure are generally modest, except for higher SiO₂ and CaO
406 abundances in experiments at 1 GPa with respect to those
407 at 1.5 GPa (Supplementary Figure 2).

408 In Fig. 5, the compositions of melts produced by GV10
409 pyroxenite are projected from diopside (Di) in the pseudo-
410 ternary diagram forsterite (Fo)–Ca–Tschermak's pyrox-
411 ene (CaTs)–quartz (Qz), together with the compositions
412 of partial melts obtained in previous pyroxenite melting
413 experiments at 1–1.5 GPa (Kogiso and Hirschmann 2001;
414 Lambart et al. 2009). Lambart et al. (2009) demonstrated

415 that the majority of melts produced by partial melting of
416 silica-deficient pyroxenites plot on the right side of the
417 CaTs–En join (Fig. 5), thus attesting the absence of a ther-
418 mal divide at $P < 2$ GPa. Partial melts from GV10 pyrox-
419 enite, which is silica-deficient too, also plot on the right
420 side of the CaTs–En join (Fig. 5), further supporting that
421 the thermal divide is not relevant at such low pressures
422 (1–1.5 GPa).

423 Mineral compositions as a function of pressure 424 and temperature

425 Average compositions of minerals from each experimental
426 run are listed in Table 3.

427 X_{Mg} value of olivine is 0.83 in the subsolidus runs at
428 1 and 1.5 GPa, and it increases with increasing degree of
429 melting up to 0.92 (Table 3; Fig. 4). CaO content in olivine
430 is rather high, spanning 0.3–0.5 wt% at both 1 and 1.5 GPa
431 (Table 3), and it is comparable to previous experimen-
432 tal studies (e.g., Libourel 1999; Kogiso and Hirschmann
433 2001).

434 *Clinopyroxene* at subsolidus conditions has X_{Mg} values
435 equal to 0.84, and rather high Na content (recalculated on
436 the basis of 6 oxygens and all Fe = Fe²⁺; Na = 0.046 and
437 0.052 a.p.f.u. at 1 and 1.5 GPa, respectively), suggesting an
438 increase of the jadeitic component with pressure (Table 3).
439 At increasing melt fraction, residual clinopyroxene shows
440 progressively higher X_{Mg} values, ranging from 0.84 to
441 0.9 (Fig. 4), and systematically lower jadeite (from about
442 5 to 1%) and TiO₂ concentrations, at both 1 and 1.5 GPa
443 (Table 3). The Al₂O₃ content does not vary significantly
444 along the melting interval, but it tends to be higher in
445 residual clinopyroxene of melting experiments at 1.5 GPa
446 (Table 3). The CaO abundance tends to increase with melt
447 fraction in experiments at 1 GPa, whereas it remains rather
448 constant at 1.5 GPa (Table 3). The Cr₂O₃ content slightly
449 increases at higher melt fraction (Table 3).

450 *Spinel*s are Al-rich and have X_{Mg} (0.68–0.77 at 1 GPa,
451 0.66–0.82 at 1.5 GPa) and X_{Cr} (0.021–0.036 at 1 GPa,
452 0.019–0.038 at 1.5 GPa) that increase with melt fraction.
453 Their TiO₂ content stays low at both 1 and 1.5 GPa (TiO₂
454 ≤ 0.26 wt%) without any correlation with pressure or melt
455 fraction (Table 3).

456 *Orthopyroxenes* have X_{Mg} ranging 0.84–0.85, and Al₂O₃
457 contents around 7–8 wt% in subsolidus runs, as well as
458 in near-solidus experiments. The CaO contents are rather
459 high, with values from 1.9 to 2.1 wt% at 1 GPa and from
460 1.8 to 2.1 wt% at 1.5 GPa (Table 3).

461 Average composition of garnet in subsolidus experi-
462 ments at 1.5 GPa and 1230 °C consists of 0.62 pyrope, 0.20
463 almandine, 0.14 grossular molar fractions, and it shows
464 very low Cr₂O₃, TiO₂, Na₂O contents (Table 3).

Table 3 Average compositions (wt%) of liquids and residual phases

Run	Phase	SiO ₂		TiO ₂		Al ₂ O ₃		Cr ₂ O ₃		FeO		MgO	
Exp. at 1 GPa													
GV10-83-4	ol (10)	39.75	(51)	0.03	(2)	0.13	(8)	0.04	(3)	15.60	(26)	43.95	(55)
<i>F</i> = 0	cpx (15)	49.98	(89)	0.34	(23)	8.79	(124)	0.12	(4)	5.47	(28)	15.91	(41)
	opx (9)	51.55	(77)	0.21	(5)	8.02	(54)	0.14	(3)	9.47	(32)	28.16	(25)
	sp (10)	0.34	(13)	0.17	(5)	63.34	(85)	2.11	(17)	15.05	(43)	17.78	(55)
GV10-83-1	gl (11)	50.92	(39)	1.98	(14)	19.66	(45)	0.02	(1)	7.99	(32)	5.03	(18)
<i>F</i> = 0.02	ol (18)	40.39	(28)	0.03	(2)	0.22	(21)	0.03	(2)	15.51	(26)	44.15	(46)
	cpx (17)	50.81	(55)	0.49	(6)	7.99	(47)	0.14	(3)	5.62	(19)	16.89	(28)
	opx (16)	53.01	(90)	0.22	(4)	7.19	(110)	0.13	(3)	9.27	(19)	28.39	(39)
	sp (4)	0.25	(9)	0.16	(5)	63.42	(7)	2.24	(14)	15.36	(39)	18.21	(41)
GV10-83-2	gl (12)	50.53	(37)	1.29	(19)	16.37	(30)	0.05	(2)	8.03	(23)	8.90	(20)
<i>F</i> = 0.13	ol (17)	40.69	(20)	0.02	(3)	0.11	(6)	0.04	(3)	13.05	(57)	46.21	(43)
	cpx (18)	51.19	(31)	0.42	(6)	7.83	(45)	0.17	(3)	5.24	(30)	17.47	(55)
	sp (10)	0.33	(14)	0.21	(5)	64.87	(147)	2.84	(37)	12.67	(113)	19.14	(56)
GV10-83-14	gl (8)	48.36	(14)	0.76	(5)	14.50	(16)	0.06	(2)	8.31	(10)	12.43	(8)
<i>F</i> = 0.40	ol (8)	40.72	(16)	0.05	(2)	0.14	(8)	0.05	(1)	10.63	(7)	48.01	(7)
	cpx (12)	50.85	(34)	0.31	(3)	7.26	(29)	0.17	(2)	4.26	(9)	17.91	(14)
	sp (5)	0.36	(7)	0.21	(3)	63.63	(40)	3.56	(30)	11.06	(19)	20.43	(38)
GV10-83-6	gl (13)	47.42	(59)	0.71	(5)	13.99	(76)	0.07	(2)	7.87	(11)	13.38	(57)
<i>F</i> = 0.57	ol (12)	40.90	(18)	0.02	(2)	0.10	(4)	0.06	(4)	9.37	(20)	49.13	(21)
	cpx (16)	50.96	(80)	0.26	(7)	7.30	(117)	0.36	(8)	3.61	(15)	17.47	(60)
GV10-83-8	gl (15)	47.96	(66)	0.48	(8)	11.54	(58)	0.18	(3)	7.11	(37)	15.23	(33)
<i>F</i> = 0.91	ol (11)	41.47	(14)	0.03	(1)	0.06	(3)	0.09	(4)	7.51	(25)	50.67	(19)
Exp. at 1.5 GPa													
GV10-83-16	ol (6)	39.48	(13)	0.02	(2)	0.10	(4)	0.05	(3)	15.60	(23)	43.90	(32)
<i>F</i> = 0	cpx (13)	50.48	(32)	0.50	(8)	8.43	(42)	0.12	(4)	5.65	(16)	17.11	(21)
	opx (9)	52.33	(69)	0.28	(17)	7.19	(91)	0.08	(6)	9.23	(14)	28.82	(40)
	sp (4)	0.27	(13)	0.16	(2)	63.20	(107)	1.96	(53)	15.72	(66)	16.97	(37)
	gnt (9)	43.24	(25)	0.23	(4)	21.53	(108)	0.18	(4)	10.52	(53)	18.27	(95)
GV10-83-9	gl (6)	51.42	(28)	1.95	(3)	19.20	(56)	0.01	(2)	7.70	(15)	7.78	(12)
<i>F</i> = 0.07	ol (11)	40.19	(34)	0.02	(2)	0.06	(3)	0.03	(3)	14.50	(14)	45.13	(22)
	cpx (11)	50.88	(26)	0.51	(3)	8.02	(23)	0.14	(4)	5.55	(19)	16.81	(26)
	opx (8)	52.49	(90)	0.22	(4)	7.80	(87)	0.14	(2)	9.02	(12)	28.73	(44)
	sp (6)	0.32	(18)	0.20	(6)	64.05	(164)	2.32	(43)	13.59	(180)	18.88	(59)
GV10-83-10	gl (10)	49.27	(34)	0.92	(21)	16.33	(46)	0.04	(2)	9.38	(21)	10.45	(47)
<i>F</i> = 0.13	ol (8)	40.69	(19)	0.02	(3)	0.11	(4)	0.04	(3)	12.92	(27)	46.39	(36)
	cpx (12)	50.34	(62)	0.42	(10)	8.49	(81)	0.17	(3)	5.13	(28)	17.24	(51)
	sp (5)	0.24	(11)	0.26	(4)	65.48	(39)	2.95	(13)	11.62	(30)	19.05	(13)
GV10-83-13	gl (15)	47.76	(36)	0.71	(8)	16.00	(18)	0.05	(3)	9.70	(24)	11.49	(42)
<i>F</i> = 0.19	ol (12)	40.22	(27)	0.02	(3)	0.11	(5)	0.04	(4)	11.99	(31)	46.92	(50)
	cpx (18)	50.19	(56)	0.36	(7)	8.64	(76)	0.23	(5)	4.44	(41)	17.78	(42)
	sp (4)	0.34	(19)	0.19	(6)	64.28	(139)	3.14	(57)	12.03	(100)	19.56	(60)
GV10-83-18	gl (10)	46.94	(20)	0.66	(5)	15.40	(19)	0.05	(3)	9.18	(13)	12.22	(14)
<i>F</i> = 0.31	ol (13)	40.60	(27)	0.02	(1)	0.10	(5)	0.05	(3)	10.92	(9)	48.07	(57)
	cpx (14)	50.48	(37)	0.34	(5)	8.80	(56)	0.24	(5)	4.10	(30)	17.62	(20)
	sp (4)	0.24	(8)	0.17	(2)	65.09	(107)	3.75	(24)	8.84	(24)	21.97	(21)
GV10-83-15	gl (9)	46.55	(88)	0.64	(2)	14.85	(55)	0.06	(2)	8.71	(48)	12.95	(7)
<i>F</i> = 0.42	ol (17)	40.47	(27)	0.04	(2)	0.07	(4)	0.04	(3)	10.73	(15)	48.39	(43)

Table 3 continued

Run	Phase	SiO ₂		TiO ₂		Al ₂ O ₃		Cr ₂ O ₃		FeO		MgO	
	cpx (17)	51.01	(66)	0.23	(3)	8.27	(69)	0.23	(7)	3.86	(11)	18.50	(33)
	sp (5)	0.28	(10)	0.15	(4)	66.50	(52)	3.91	(66)	8.89	(19)	22.06	(11)
GV10-83-21	gl (9)	46.87	(26)	0.63	(15)	13.97	(16)	0.05	(2)	8.49	(33)	13.54	(43)
<i>F</i> = 0.52	ol (17)	40.60	(26)	0.02	(2)	0.08	(3)	0.05	(2)	10.09	(76)	48.39	(53)
	cpx (17)	50.47	(51)	0.22	(6)	8.39	(50)	0.33	(9)	3.85	(19)	19.11	(19)
GV10-83-22	gl (9)	47.76	(32)	0.58	(7)	11.93	(38)	0.11	(3)	6.71	(4)	15.23	(38)
<i>F</i> = 0.88	ol (17)	41.70	(22)	0.01	(2)	0.10	(2)	0.07	(4)	7.40	(14)	49.53	(29)
Run	Phase	CaO		Na ₂ O		K ₂ O		Total	X _{Mg}		CIPW ^a		
Exp. at 1 GPa													
GV10-83-4	ol (10)	0.41	(4)	0.01	(1)	–		99.91	(71)	0.83	(0)		
<i>F</i> = 0	cpx (15)	19.00	(56)	0.66	(6)	0.09	(0)	100.37	(51)	0.84	(1)		
	opx (9)	1.93	(68)	0.04	(3)	–		99.52	(79)	0.84	(0)		
	sp (10)	0.24	(5)	0.01	(2)	–		99.02	(32)	0.68	(1)		
GV10-83-1	gl (11)	10.47	(57)	2.17	(15)	0.44	(12)	98.67	(46)	0.53	(2)	4/20/8/0/64	
<i>F</i> = 0.02	ol (18)	0.48	(13)	0.01	(1)	–		100.81	(48)	0.83	(0)		
	cpx (17)	17.89	(31)	0.34	(4)	0.01	(0)	100.19	(42)	0.84	(0)		
	opx (16)	2.11	(13)	0.05	(3)	–		100.37	(37)	0.85	(0)		
	sp (4)	0.22	(3)	0.01	(1)	–		99.84	(59)	0.68	(1)		
GV10-83-2	gl (12)	12.38	(57)	1.41	(11)	0.27	(7)	99.23	(70)	0.66	(1)	1/26/20/0/51	
<i>F</i> = 0.13	ol (17)	0.44	(9)	0.01	(1)	–		100.57	(47)	0.86	(1)		
	cpx (18)	17.67	(77)	0.29	(4)	0.02	(1)	100.30	(36)	0.86	(1)		
	sp (10)	0.26	(7)	0.01	(1)	0.01	(0)	100.34	(86)	0.74	(2)		
GV10-83-14	gl (8)	13.84	(9)	0.84	(8)	0.15	(4)	99.25	(37)	0.73	(0)	0/16/27/12/44	
<i>F</i> = 0.40	ol (8)	0.27	(8)	0.01	(1)	–		99.88	(26)	0.88	(0)		
	cpx (12)	19.24	(20)	0.22	(3)	0.02	(0)	100.24	(56)	0.88	(0)		
	sp (5)	0.27	(4)	0.00	(0)	–		99.52	(42)	0.77	(1)		
GV10-83-6	gl (13)	14.49	(68)	0.66	(7)	0.12	(8)	98.71	(75)	0.75	(1)	0/13/30/14/41	
<i>F</i> = 0.57	ol (12)	0.45	(3)	0.01	(1)	–		100.04	(33)	0.90	(0)		
	cpx (16)	19.69	(38)	0.18	(2)	0.01	(0)	99.85	(40)	0.90	(1)		
GV10-83-8	gl (15)	15.77	(28)	0.46	(6)	0.08	(1)	98.79	(29)	0.79	(1)	0/9/40/17/34	
<i>F</i> = 0.91	ol (11)	0.51	(2)	0.00	(0)	–		100.34	(26)	0.92	(0)		
Exp. at 1.5 GPa													
GV10-83-16	ol (6)	0.38	(3)	0.01	(1)	–		99.53	(35)	0.83	(0)		
<i>F</i> = 0	cpx (13)	17.74	(32)	0.75	(3)	0.11	(0)	100.89	(22)	0.84	(0)		
	opx (9)	1.82	(9)	0.05	(2)	–		99.80	(57)	0.85	(0)		
	sp (4)	0.28	(12)	0.01	(1)	–		98.56	(150)	0.66	(1)		
	gnt (9)	6.89	(87)	0.04	(3)	0.01	(0)	100.90	(64)	0.76	(1)		
GV10-83-9	gl (6)	8.21	(22)	1.99	(10)	0.38	(6)	98.71	(55)	0.64	(1)	4/31/0/0/61	
<i>F</i> = 0.07	ol (11)	0.41	(5)	0.01	(1)	–		100.34	(22)	0.84	(1)		
	cpx (11)	18.00	(29)	0.42	(4)	0.03	(1)	100.36	(36)	0.84	(0)		
	opx (8)	2.12	(14)	0.05	(3)	–		100.57	(36)	0.85	(0)		
	sp (6)	0.26	(8)	0.01	(1)	0.01	(0)	99.62	(35)	0.71	(3)		
GV10-83-10	gl (10)	10.41	(52)	1.55	(24)	0.26	(8)	98.61	(55)	0.66	(1)	0/27/12/7/52	
<i>F</i> = 0.13	ol (8)	0.47	(14)	0.01	(1)	–		100.65	(28)	0.86	(0)		
	cpx(12)	17.75	(83)	0.37	(7)	0.04	(1)	99.95	(46)	0.86	(1)		
	sp (5)	0.28	(10)	0.00	(0)	0.01	(0)	99.90	(49)	0.75	(1)		
GV10-83-13	gl (15)	11.82	(29)	1.26	(22)	0.21	(2)	99.10	(35)	0.68	(2)	0/16/17/16/50	
<i>F</i> = 0.19	ol (12)	0.40	(13)	0.01	(1)	–		99.70	(44)	0.87	(0)		

Table 3 continued

Run	Phase	CaO		Na ₂ O		K ₂ O		Total	X _{Mg}	CIPW ^a
GV10-83-18 <i>F</i> = 0.31	cpx (18)	17.67	(69)	0.26	(2)	0.03	(1)	99.61 (50)	0.88 (1)	0/11/23/18/47
	sp (4)	0.35	(6)	0.01	(1)	–		99.89 (39)	0.74 (2)	
	gl (10)	13.16	(22)	0.94	(2)	0.18	(9)	98.73 (17)	0.70 (1)	
	ol (13)	0.36	(6)	0.01	(1)	–		100.12 (30)	0.88 (0)	
GV10-83-15 <i>F</i> = 0.42	cpx (14)	18.05	(55)	0.21	(3)	0.02	(0)	99.88 (49)	0.88 (1)	0/8/27/19/44
	sp (4)	0.16	(4)	0.00	(0)	–		100.21 (112)	0.82 (1)	
	gl (9)	14.16	(67)	0.69	(8)	0.12	(8)	98.72 (26)	0.72 (1)	
	ol (17)	0.26	(6)	0.01	(1)	–		99.99 (36)	0.88 (0)	
GV10-83-21 <i>F</i> = 0.52	cpx (17)	18.24	(44)	0.19	(4)	0.05	(1)	100.58 (46)	0.90 (0)	0/8/31/19/41
	sp (5)	0.15	(5)	0.00	(1)	0.01	(0)	101.96 (58)	0.82 (0)	
	gl (9)	14.76	(43)	0.58	(9)	0.11	(6)	99.02 (39)	0.74 (1)	
	ol (17)	0.46	(7)	0.01	(1)	–		99.69 (49)	0.89 (1)	
GV10-83-22 <i>F</i> = 0.88	cpx (17)	17.12	(28)	0.17	(3)	0.04	(3)	99.69 (39)	0.90 (1)	0/8/38/17/35
	gl (9)	15.65	(21)	0.46	(5)	0.12	(11)	98.55 (49)	0.80 (0)	
	ol (17)	0.49	(3)	0.03	(2)	–		99.33 (35)	0.92 (0)	

Numbers in parentheses correspond to 1σ standard deviations, given in terms of the least unit cited; e.g., 5.2 (29) and 74.4 (27) represent 5.2 ± 2.9 and 74.4 ± 2.7, respectively

gl glass, *ol* olivine, *cpx* clinopyroxene, *opx* orthopyroxene, *sp* spinel, *gnt* garnet. Along with the phase, we report the number of analyses in parentheses

^a CIPW norms: quartz/hypersthene/diopside/olivine/feldspars. “*F*” below the run label is the melt fraction calculated in each experiments

465 **Solidus temperature and melt productivity**

466 The solidus of pyroxenite GV10 is between 1180 and
467 1200 °C at 1 GPa, and between 1230 and 1250 °C at
468 1.5 GPa (Fig. 6). At pressure of 1 and 1.5 GPa, pyroxenite
469 GV10 starts melting at lower temperature, compared to all
470 fertile lherzolites experimentally investigated (Fig. 6).

471 Melt fraction vs. temperature trends at both 1 and
472 1.5 GPa are concave upward, suggesting that isobaric
473 melt productivity (dF/dT)_P increases with temperature
474 (Fig. 7), in agreement with previous experimental results
475 on pyroxenite or eclogite compositions (e.g., Pickering-
476 Witter and Johnson 2000; Kogiso and Hirschmann 2001;
477 Pertermann and Hirschmann 2003b; Lambart et al. 2009a),
478 and thermodynamic models of peridotite partial melting
479 (i.e. Hirschmann et al. 1999; Asimow et al. 2001). At both
480 pressures, the isobaric melt productivity at low melting
481 degrees (*F* < 12.6%) is about 0.2%/°C (Fig. 7). At higher
482 melting degrees (*F* > 12.6%), the mean isobaric productiv-
483 ity is 0.9%/°C at 1 GPa and 0.6%/°C at 1.5 GPa (Fig. 7).

484 **Discussion**

485 **Melting and phase relations**

486 Phase abundances and glass proportions have been cal-
487 culated by weighted least squares mass balance that

488 included nine oxides, SiO₂, TiO₂, Al₂O₃, Cr₂O₃, FeO,
489 MgO, CaO, Na₂O and K₂O. Detailed explanation of mass
490 balance calculations is reported in Borghini et al. (2010).

491 In the pressure and temperature range investigated
492 in this study, olivine is stable along the whole melt-
493 ing range with low modal abundances (<14%; Table 2;
494 Fig. 8). Clinopyroxene is the dominant phase in GV10
495 pyroxenite residual assemblage (Table 2; Fig. 8). Spinel
496 persists over than 40% of melt fraction at both pressures
497 (Table 2; Fig. 8). Orthopyroxene disappears from the res-
498 idue at *F* higher than 12%, and low modal abundance of
499 garnet is stable only in the subsolidus assemblage at 1.5
500 GPa (Table 2; Fig. 8). The sequence of mineral phases
501 disappearance at increasing melt fraction is, therefore,
502 orthopyroxene–spinel–clinopyroxene–olivine and garnet–
503 orthopyroxene–spinel–clinopyroxene–olivine, at 1 and
504 1.5 GPa, respectively.

505 To constrain the stoichiometry of reactions control-
506 ling the consumption of mineral phases along the melting
507 interval of GV10, we performed mass balance calcula-
508 tions (within the same chemical system used for deter-
509 mination of modal abundances, see above), using the
510 mineral and glass compositions at the beginning and the
511 end of each phase exhaustion step with increasing melt
512 fraction. Average melting reactions for pyroxenite GV10
513 are reported in Table 4, together with those for pyroxen-
514 ites and peridotites at 1 and 1.5 GPa available in the lit-
515 erature. Melting reactions in pyroxenite GV10 at 1.5 GPa

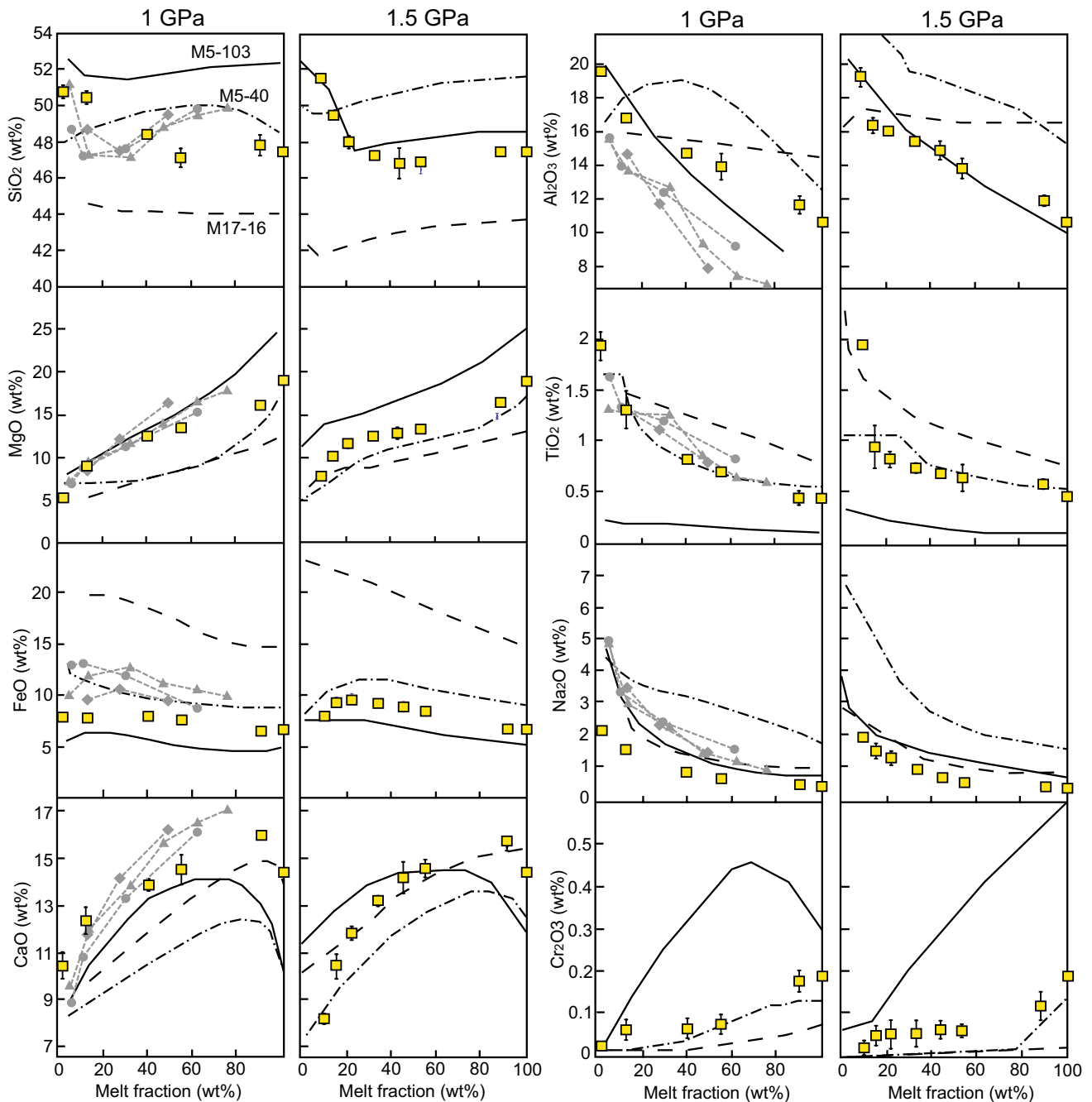


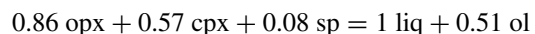
Fig. 3 Oxide concentrations in experimental melts plotted as a function of melt fractions (wt%). Data plotted at $F = 100\%$ correspond to the bulk-rock composition. Also reported for comparison are trends (black lines) extrapolated by melting experiments at 1 and 1.5 GPa on three silica-deficient pyroxenites by Lambart et al. (2009a). Grey

symbols linked by dotted lines are experimental results at 1 GPa on clinopyroxenite bulks (Kogiso and Hirschmann 2001). Comparison between the compositions of melts produced by GV10 at 1 and 1.5 GPa is illustrated in Supplementary Figure 2

516 are similar to those at 1 GPa, except for the first melting
 517 interval at 1.5 GPa in which garnet is rapidly consumed
 518 (reaction 14, Table 4).

519 Although the melting relation $opx + cpx + sp \rightarrow$
 520 $liq + ol$ is the same for pyroxenites and peridotite
 521 (Table 4), the contribution of orthopyroxene with respect to

clinopyroxene is higher when pyroxenite is considered, as
 suggested for example by reaction 1 at 1 GPa (Table 4):



This results in rapid consumption of orthopyroxene dur-
 ing melting, and is likely related to the high CaO content in

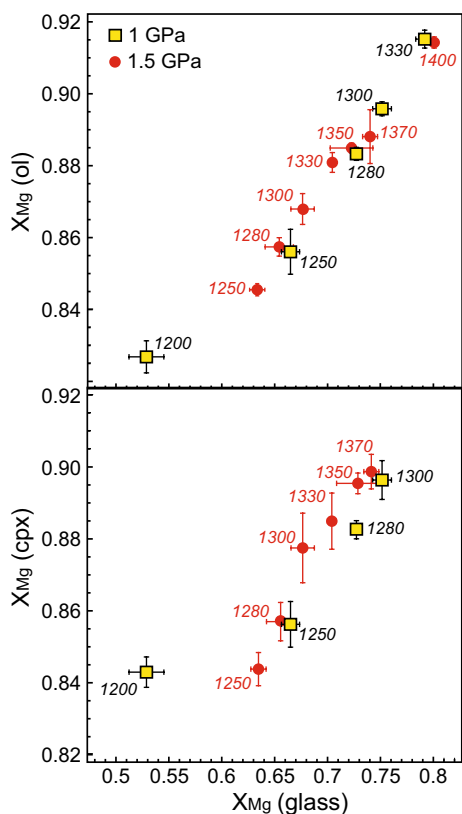


Fig. 4 Fe-Mg partitioning of glass vs. olivine and clinopyroxene. $X_{Mg} = Mg/(Mg + Fe^{tot})$

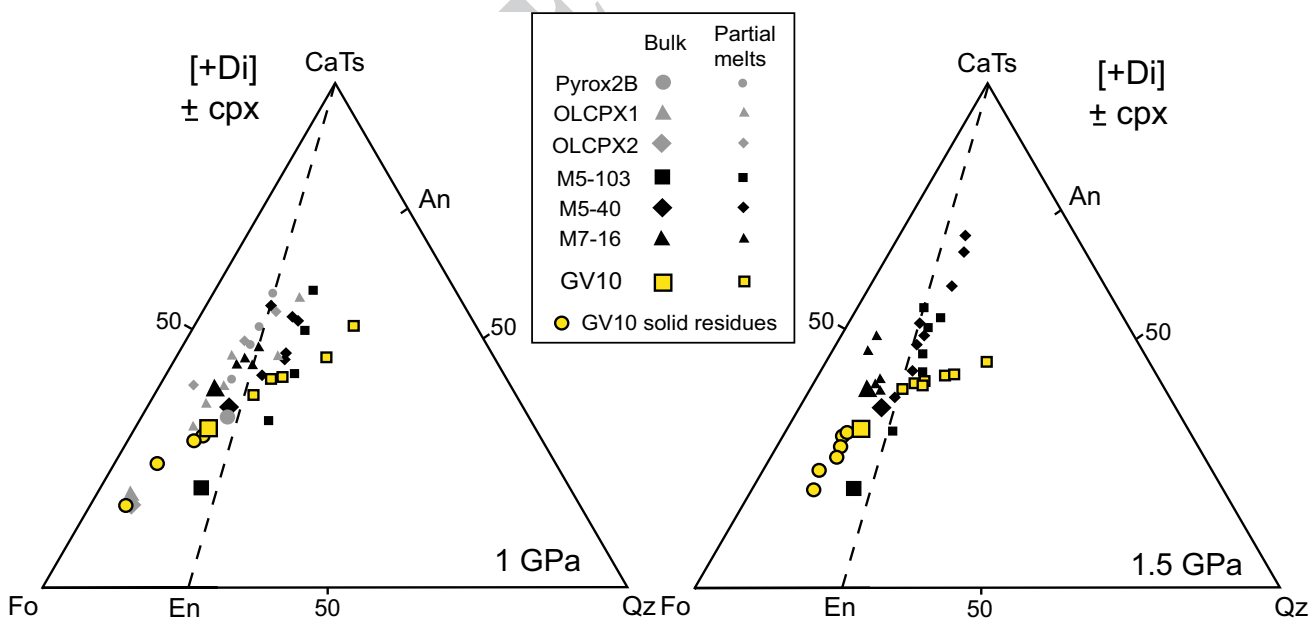
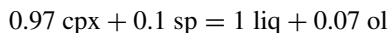


Fig. 5 Molar projections from diopside [Di] onto the pseudo-ternary diagram forsterite (Fo)—calcium tschermak (CaTs)—quartz (Qz) (O'Hara 1972) of melts produced by experiments on GV10 at 1 and 1.5 GPa. The *small yellow circles* represent the composition of solid residues calculated using the modes and composition of residual

the bulk rock (Fig. 1), favoring the stabilization of clinopyroxene over orthopyroxene in the residue. Similar melting relation was documented in other pyroxenites and clinopyroxene-rich lherzolites (e.g., FER-B, Pickering-Witter and Johnson 2000; M5-103, Lambart et al. 2009a).

At *F* higher than 12%, clinopyroxene is the main phase that contributes to melting, with minor contribution of spinel until its exhaustion (Table 4; Fig. 8), following the average reaction (e.g., at 1 GPa, reaction 2, Table 4):



The coefficients are comparable to those obtained in previous experimental works on anhydrous pyroxenites and clinopyroxene-rich peridotites (reactions 5 and 6, Table 4). Around *F* = 90%, clinopyroxene is also consumed and olivine is the liquidus phase (reactions 3 and 17, Table 4; Fig. 8).

As expected, along the whole melting interval, residual phases show progressively higher X_{Mg} values at increasing melt fraction (Table 3; Fig. 4). At low melting degrees, the high coefficient of orthopyroxene involved in the melting reaction (reactions 1, 14 and 15, Table 4) results in relatively high SiO_2 content in partial melts at both pressures (Fig. 3). After orthopyroxene exhaustion, melt composition is controlled by the composition of coexisting clinopyroxene and, to a lesser extent, spinel. CaO content of partial melts increases with melt fraction until the disappearance of clinopyroxene (Fig. 3).

minerals in experimental charges. The melt compositions reported in previous experimental studies on other pyroxenites are also reported (Lambart et al. 2009a; Kogiso and Hirschmann 2001). Symbols are as in Fig. 1. Larger symbols represent the starting materials

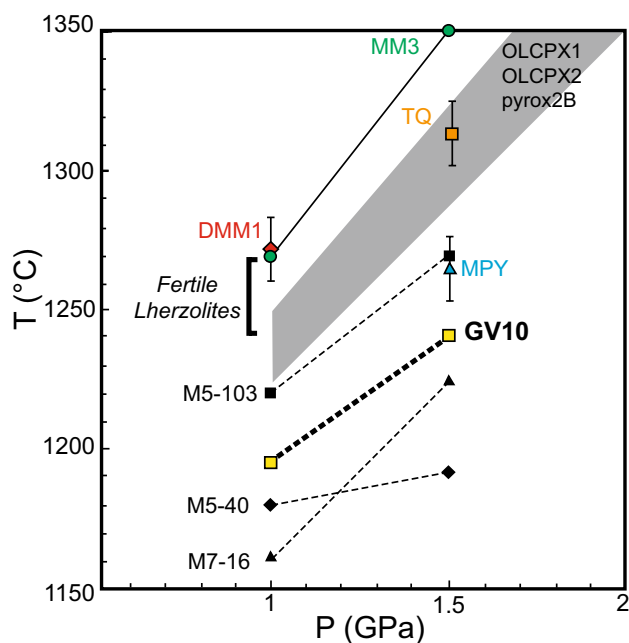


Fig. 6 Experimentally determined solidus temperature at 1 and 1.5 for pyroxenite GV10 compared to other pyroxenites and peridotites. Pyroxenites are: M5-103, M5-40 and M7-16 studied by Lambart et al. (2009a); OLCPX1, OLCPX2 and pyrox2B by Kogiso and Hirschmann (2001). Peridotites are: lherzolite MM3 (Falloon et al. 2008), fertile lherzolite MPY and depleted lherzolite TQ (Robinson et al. 1998), depleted lherzolite DMM1 (Wasylenki et al. 2003). Symbols are as in Fig. 1. Temperature interval named “Fertile peridotite” at 1.0 GPa includes experiments of variably fertile lherzolites from Pickering-Witter and Johnson (2000) and Schwab and Johnston (2001)

553 The most striking compositional feature observed in
 554 clinopyroxene is the decrease in jadeite component with
 555 increasing melt fraction. As a whole, the low Na_2O contents
 556 in melts (see Fig. 3) reflect relatively high partition coefficient
 557 of Na between residual clinopyroxene and partial melt
 558 ($D_{\text{Na}_2\text{O}}^{\text{cpx/liquid}}$). As shown in Fig. 9, $D_{\text{Na}_2\text{O}}^{\text{cpx/liquid}}$ calculated
 559 from experiments on pyroxenite GV10 are positively correlated
 560 with pressure and temperature. The increase of $D_{\text{Na}_2\text{O}}^{\text{cpx/liquid}}$
 561 with pressure observed in GV10 experiments is well correlated
 562 with the $D_{\text{Na}_2\text{O}}^{\text{cpx/liquid}}$ calculated from experiments at 2
 563 and 2.5 GPa on pyroxenites with comparable bulk composition
 564 (MIX1G, Hirschmann et al. 2003; M5-40, Lambart
 565 et al. 2013) (Fig. 9a). This overall correlation is in good
 566 agreement with available quantitative models for Na partitioning
 567 between clinopyroxene and silicate melt (Blundy
 568 et al. 1995; Villiger et al. 2007), as well as with the trend
 569 defined by fitting literature experimental data (Villiger et al.
 570 2007) (Fig. 9a).

571 On the other hand, the correlation of Na cpx/liquid partitioning
 572 with temperature has been poorly discussed in the literature. Figure
 573 9b reports a compilation of $D_{\text{Na}_2\text{O}}^{\text{cpx/liquid}}$, derived from experiments
 574 on fertile peridotites and

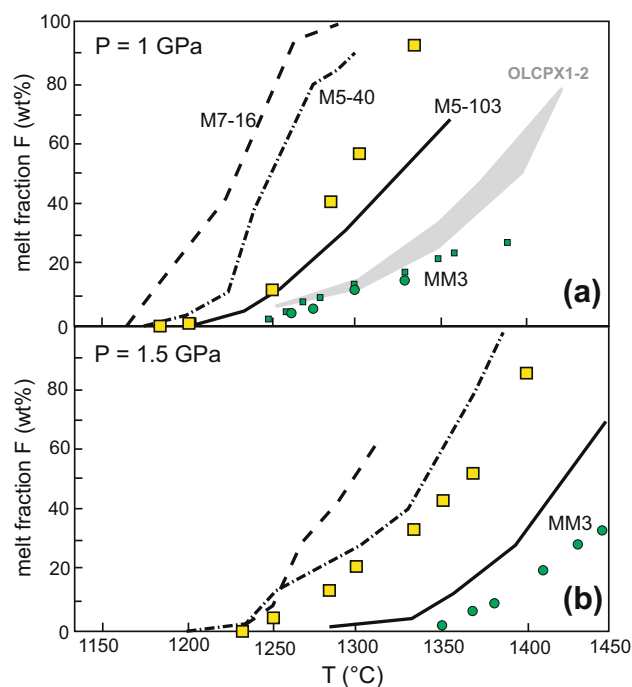


Fig. 7 Temperature-melt fraction curves determined at 1 GPa (a) and 1.5 GPa (b) for pyroxenite GV10 (yellow squares) compared to those extrapolated by previous experimental work on dry pyroxenite and peridotite: (1) pyroxenites M5-103, M5-40 and M7-16 from Lambart et al. (2009a); (2) clinopyroxenites Pyrox2B, OLCPX1 and OLCPX2 from Kogiso and Hirschmann (2001); (3) lherzolite MM3 from Baker et al. (1995) (green squares) and Falloon et al. (2008) (green circles)

575 pyroxenites at 0.7–2.5 GPa (see the figure caption for refer-
 576 ences). Overall, these data point to a positive correlation
 577 of Na cpx/liquid partitioning with temperature (Fig. 9b).
 578 Blundy et al. (1995) argued that, at 1 GPa, temperature
 579 dependence of $D_{\text{Na}_2\text{O}}^{\text{cpx/liquid}}$ is greatly subordinate to that of
 580 pressure. In fractional crystallization experiments, Villiger
 581 et al. (2007) observed that the Na distribution coefficients
 582 between clinopyroxene and liquid decrease at decreasing
 583 temperature, as observed in our experiments (Fig. 9b), but
 584 they concluded that a clear relationship between Na parti-
 585 tioning and temperature cannot be established. Our results
 586 corroborate the existence of such correlation, although a
 587 proper understanding of factors controlling such relation-
 588 ship would need a more focused study.

589 Similar to Na_2O , Al_2O_3 and TiO_2 behave as incompatible
 590 elements (Fig. 3), whereas Cr_2O_3 is compatible and mostly
 591 partitioned in residual spinel (Table 3; Fig. 3).

592 Comparison with previous studies

593 The solidus of GV10 at 1 and 1.5 GPa is located at lower tem-
 594 perature with respect to experimental results on variably fer-
 595 tile lherzolites (Pickering-Witter and Johnson 2000; Schwab
 596 and Johnston 2001) and clinopyroxenites investigated by

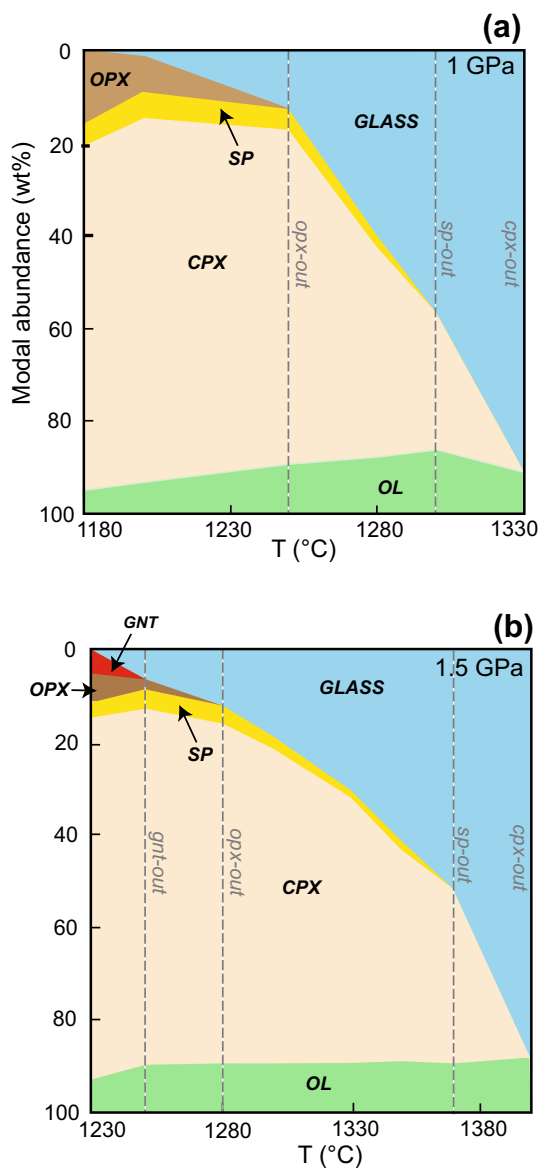


Fig. 8 Modal abundance calculated by weighted least squares mass balance in experiments at 1 GPa (a) and 1.5 GPa (b) as a function of temperature. The grey dotted lines represent the temperature of mineral phases disappearance. ol olivine, cpx clinopyroxene, opx orthopyroxene, sp spinel, gnt garnet

AQ4 Kogiso and Hirschmann (2001) (Fig. 6). Melting experiments at 1 GPa on these clinopyroxenites have clinopyroxene + olivine as residual minerals (Kogiso and Hirschmann 2001). In our experiments, the stability of residual spinel, in addition to clinopyroxene and olivine, likely resulted from higher Al₂O₃ content of pyroxenite GV10 (Table 1). Lower CaO and higher MgO contents in GV10 could have enhanced minor modal orthopyroxene (Tables 1, 2), never found in clinopyroxenites (Kogiso and Hirschmann 2001).

The solidus curve and the isobaric melt productivity (dF/dT)_p of pyroxenite GV10 are in agreement with the data

reported for pyroxenites previously investigated by Lambart et al. (2009a) (Fig. 7). In particular, pyroxenite GV10 displays solidus and melt productivity curves intermediate between M5-40 and M5-103 (Figs. 6, 7).

According to Kogiso et al. (2004a), alkali contents and X_{Mg} of the bulk composition mostly influence the solidus temperatures of natural anhydrous pyroxenites. On this basis, Lambart et al. (2016) have recently proposed an empirical parameterization to derive the melt fraction versus temperature from 5% melting up to clinopyroxene disappearance for pyroxenites as a function of pressure, temperature and bulk composition. Pyroxenite GV10 has X_{Mg} and (Na₂O + K₂O) lower than M5-103 (Table 1). The model successfully predicts the relative location of the (dF/dT)_p between GV10 and pyroxenites studied by Lambart et al. (2009a), but locates this curve at slightly higher temperature with respect to the curve extrapolated by the experiments.

The most important differences in the melting relations between GV10 and the pyroxenites M5-40 and M5-103 of Lambart et al. (2009a) are the presence of plagioclase and the abundance of spinel. Both M5-40 and M5-103 have plagioclase in the subsolidus assemblage at 1 GPa, 30 and 4%, respectively, whereas plagioclase is not stable in GV10 pyroxenite. This is due to the much lower Na₂O/CaO ratio of GV10, which tends to stabilize plagioclase at lower pressure (Green and Falloon 1998; Borghini et al. 2010). Spinel is present in very low modal abundance (<1%) or absent in the experiments on M5-40 and M5-103. At 1.5 GPa, spinel is absent in M5-103 but it is present at near-solidus conditions in equilibrium with plagioclase and garnet in the M5-40 (Lambart et al. 2009a).

The variational trends of MgO and FeO abundances in melts at both 1 and 1.5 GPa obtained for pyroxenite GV10 are similar to those of M5-40 and M5-103, with absolute contents mostly varying at intermediate values (Fig. 3), according to the bulk compositions (Table 1). TiO₂ and Cr₂O₃ compositional trends defined by GV10 melts follow those observed for M5-40 pyroxenite, except for the higher TiO₂ contents in GV10 at low *F*. Progressive Al₂O₃ content decrease with increasing *F* at both pressures is similar to what observed for M5-103, but with a lower slope, because spinel is a residual phase up to relatively higher *F* in GV10 pyroxenite. The high CaO contents of GV10 melts mimic those documented in melts produced by clinopyroxenites at 1 GPa (Kogiso and Hirschmann 2001), and by garnet-olivine clinopyroxenite M7-16 at 1.5 GPa (Lambart et al. 2009a) (Fig. 3). At both investigated pressure conditions, the variation of SiO₂ contents in GV10 melts defines a trajectory of rapid decrease at low *F* and a slight increase at higher *F*, similar to M5-103 but more accentuated (Fig. 3). This trend results by early exhaustion of orthopyroxene and

Table 4 Melting reactions (wt%) for pyroxenites and peridotites at 1 and 1.5 GPa

Reference	Lithotype	Bulk	Melting interval	Melting reactions	R-n ^a
1 GPa					
This study	pyrox	GV10	Solidus to opx-out	0.86 opx + 0.57 cpx + 0.08 sp = 1 liq + 0.51 ol	1
			opx-out to sp-out	0.97 cpx + 0.1 sp = 1 liq + 0.07 ol	2
			sp-out to cpx-out	1.04 cpx = 1 liq + 0.04 ol	3
Lambart et al. (2009)	pyrox	M5-103	Main reaction	0.49 opx + 0.61 cpx = 1 liq + 0.10 ol	4
	pyrox	M7-16	Main reaction	0.95 cpx + 0.1 sp = 1 liq + 0.05 ol	5
Pickering-Witter and Johnson (2000)	perid	FER-B	Solidus to cpx-out	0.99 cpx + 0.05 sp = 1 liq + 0.04 ol	6
	perid	FER-D	Solidus to cpx-out	0.14 opx + 0.75 cpx + 0.17 sp = 1 liq + 0.06 ol	7
	perid	FER-E	Solidus to cpx-out	0.27 opx + 0.78 cpx + 0.05 sp = 1 liq + 0.10 ol	8
Baker and Stolper (1994)	perid	MM3	Main reaction	0.38 opx + 0.71 cpx + 0.13 sp = 1 liq + 0.22 ol	9
Falloon et al. (2008)	perid	MM3	Solidus to plag-out	0.14 ol + 0.31 opx + 0.81 plag = 0.75 liq + 0.48 cpx + 0.004 sp	10
			Plag-out to cpx-out	0.275 opx + 0.887 cpx + 0.052 sp = 1.036 liq + 0.178 ol	11
Wasylenki et al. (2003)	perid	DMM1	Solidus to cpx-out	0.56 opx + 0.72 cpx + 0.04 sp = 1 liq + 0.34 ol	12
			Above cpx + sp-out	1.24 opx = 1 liq + 0.24 ol	13
1.5 GPa					
This study	pyrox	GV10	Solidus to gnt-out	0.59 gnt + 0.82 opx + 0.18 cpx = 1 liq + 0.59 ol	14
			gnt-out to opx-out	0.87 opx + 0.49 cpx + 0.1 sp = 1 liq + 0.46 ol	15
			opx-out to sp-out	0.97 cpx + 0.09 sp = 1 liq + 0.06	16
			sp-out to cpx-out	1.07 cpx = 1 liq + 0.07	17
Falloon et al. (2008)	perid	MM3	Solidus to cpx-out	0.275 opx + 0.887 cpx + 0.052 sp = 1.036 liq + 0.178 ol	18
Robinson et al. (1998)	perid	MPY	0–1%	0.61 opx + 0.58 cpx + 0.001 sp = 1 liq + 0.19 ol	19
	perid	TQ	Main reaction	0.33 opx + 0.83 cpx + 0.08 sp = 1 liq + 0.24 ol	20

pyrox pyroxenite, perid peridotite, opx orthopyroxene, cpx clinopyroxene, ol olivine, sp spinel, gnt garnet, plag plagioclase, liq liquidus

^a Reaction number (as discussed in the text)

661 late disappearance of spinel. Variation trends of Na₂O
662 contents, marked by progressive decrease at increasing
663 F, are similar to those documented for all pyroxenites at
664 1–1.5 GPa, but partial melts of GV10 display the lowest
665 Na₂O absolute concentrations (Fig. 3), consistent with
666 higher $D_{\text{Na}_2\text{O}}^{\text{cpx/liquid}}$ (Fig. 9). Differently from pyroxenites
667 investigated by Lambart et al. (2009), no partial melts
668 with nepheline-normative compositions are produced by
669 GV10 pyroxenite at 1–1.5 GPa. By contrasts, at low melt-
670 ing degrees GV10 melts have quartz-normative composi-
671 tions (Table 3).

672 Major element chemical signature 673 of pyroxenite-derived melts

674 Lambart et al. (2013) highlighted that the determination
675 of pyroxenite contribution in major element composition
676 of oceanic basalts may be difficult because, at identi-
677 cal pressure and temperature conditions, many pyroxen-
678 itic melts are quite similar to peridotitic melts. Although

679 many pyroxenites are expected to have lower X_{Mg} and
680 higher TiO₂ and alkali contents, these chemical features
681 are diluted in their partial melts by the higher degrees of
682 melting compared to peridotites (e.g., Médard et al. 2006;
683 Lambart et al. 2009a). However, some pyroxenites can
684 produce SiO₂-poor and/or Fe-rich melts that are able to
685 significantly affect the composition of melts produced
686 by a pyroxenite–peridotite mixed mantle source. Accord-
687 ingly, Lambart et al. (2013) identified a group of primitive
688 MORB that represents the best candidates to preserve this
689 pyroxenite signature, even after the mixing with perido-
690 titic melts (px-MORB; with MgO ≥ 9 wt%, SiO₂ < 48.8
691 wt% and FeO > 8.3). Further major element markers for
692 pyroxenite contribution, namely high CaO/Al₂O₃ ratios
693 and high TiO₂ contents, have been proposed for OIBs
694 (e.g., Prytulak and Elliot 2007; Jackson and Dasgupta
695 2008; Shorttle and MacLennan 2011). Moreover, pri-
696 mary magma calculations (PRIMELTS2.XLS, Herzberg
697 and Asimow 2008) suggest that low CaO–MgO trends
698 in some OIBs are not consistent with the composition of

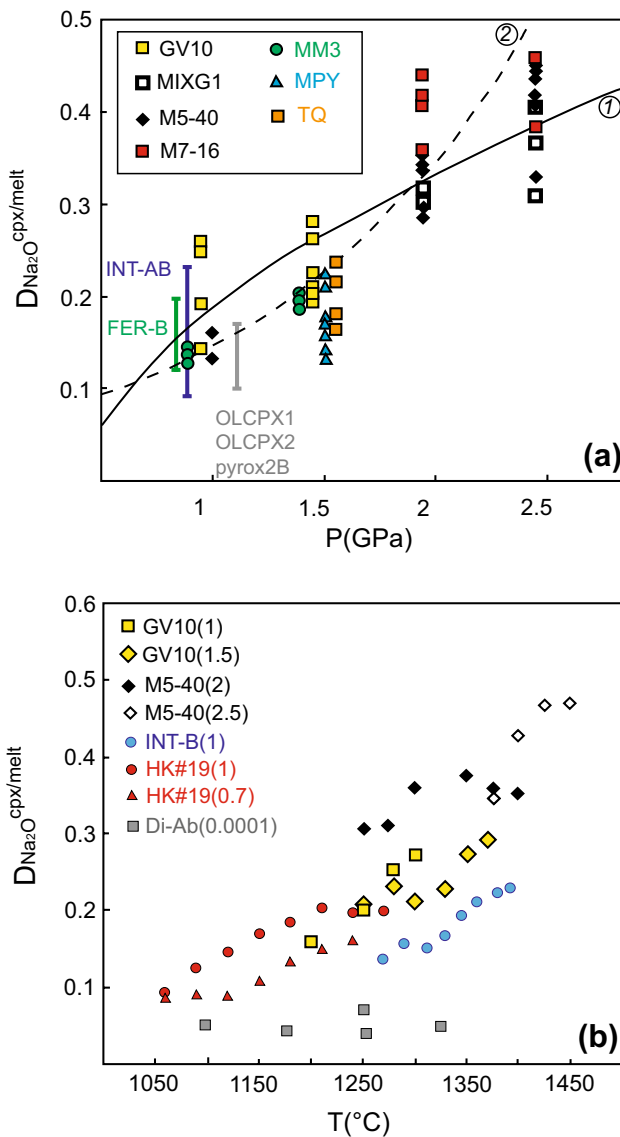


Fig. 9 Partitioning of Na₂O between clinopyroxene and glass ($D_{Na_2O}^{cpx/liquid}$) as a function of pressure (a) and temperature (b). Our data are compared with $D_{Na_2O}^{cpx/liquid}$ from experiments on lherzolites MM3 (Falloon et al. 2008), MPY and TQ (Robinson et al. 1998), fertile lherzolites FER-B (Pickering-Witter and Johnson 2000), INT-A and INT-B (Schwab and Johnston 2001), pyroxenites MIXG1 from Hirschmann et al. (2003), M5-40 and M7-16 from Lambart et al. (2009a), OLCPX1, OLCPX2 and pyrox2B from Kogiso and Hirschmann (2001). Data in the system diopside-albite are from Blundy et al. (1995) and from fractional and equilibrium crystallization experiments are by Villiger et al. (2007). Curves 1 and 2 in (a) refer to quantitative model proposed by Blundy et al. (1995) and fit of literature experimental data by Villiger et al. (2007), respectively. Number between parenthesis in (b) refer to experimental pressure (GPa)

699 melts from a pure peridotite and they can discriminate the
 700 contribution of melts derived from a low-Ca pyroxenite,
 701 as the modeled olivine-free Px-1 (Sobolev et al. 2005,
 702 2007).

To identify the major element signature of secondary
 703 pyroxenite GV10, we compare the major element
 704 composition of melts produced in our experiments with those
 705 of melts from experiments on various peridotite bulks,
 706 at 1.0 and 1.5 GPa (Figs. 10, 11, see figure captions for
 707 references). We only considered the melt compositions
 708 produced at the same temperature interval investigated
 709 in this study, i.e. 1200–1330 °C at 1 GPa and 1250–
 710 1400 °C at 1.5 GPa.

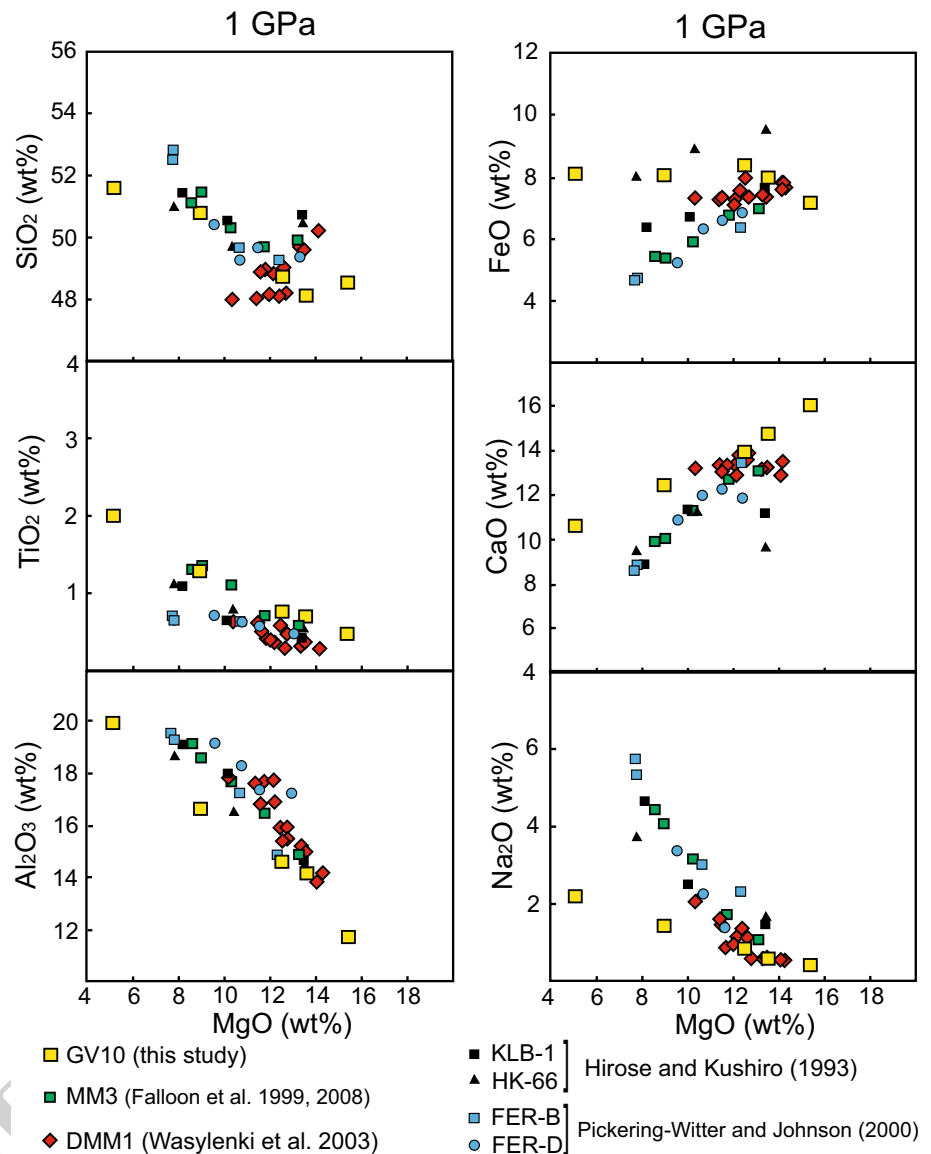
GV10 melts span within almost the same MgO range
 712 defined by experimental peridotitic melts (Figs. 10, 11),
 713 except for low MgO contents at very low melting degree
 714 ($F < 2\%$) and high MgO abundances in melts resulting
 715 from F of about 90% at 1 GPa (Fig. 10).

At 1 GPa, SiO₂, TiO₂, FeO and Al₂O₃ abundances of
 717 GV10 melts largely match the composition of many peridotite-derived melts. However, at low melting degrees
 718 GV10 melts display higher FeO and lower Al₂O₃ contents
 719 relative to melts from a normal fertile lherzolite
 720 (e.g., MM3, Fig. 10). Peculiar chemical features are the
 721 CaO and Na₂O vs MgO trends (Fig. 10). At similar MgO,
 722 pyroxenite GV10 tends to produce melts with higher CaO
 723 and lower Na₂O contents than peridotite-derived melts
 724 (Fig. 10). For high degrees of melting (i.e. the highest
 725 MgO abundances), GV10 melts display CaO contents
 726 much higher than all fertile lherzolites, but coupled to
 727 very low Na₂O contents. Comparably low Na₂O abundances
 728 are only observed in melts produced by moderate melting
 729 degrees of depleted lherzolites (e.g., DMM1,
 730 Wasylenki et al. 2003; Fig. 10).

At 1.5 GPa, GV10 melts resemble those of melts produced
 733 by very fertile lherzolites (e.g., lherzolite HK-66,
 734 Hirose and Kushiro 1993, or refertilized lherzolites KG1
 735 and KG2 from Kogiso et al. 1998), except for much lower
 736 Na₂O contents in the entire pyroxenite melting interval,
 737 and slightly lower SiO₂ coupled to higher CaO in high-
 738 MgO pyroxenitic melts (Fig. 11). These latter also show
 739 higher FeO and slightly lower Al₂O₃ and TiO₂ contents
 740 compared to melts from a normal fertile lherzolite (e.g.,
 741 MM3; Fig. 11).

In summary, as previously observed by Lambart et al.
 743 (2009a, 2013), for many oxides the compositions of partial
 744 melts from secondary pyroxenite GV10 are similar
 745 or very close to those of melts produced by depleted and
 746 variably fertile peridotites. The most striking geochemical
 747 signature of pyroxenite GV10 melts is their very low
 748 Na₂O contents. In addition, they display higher CaO and
 749 slightly lower Al₂O₃ concentrations, thus resulting in
 750 higher CaO/Al₂O₃ ratio than peridotitic melts. In the next
 751 section, we discuss to what extent the GV10 chemical
 752 signature can impact the compositions of melts produced
 753 by decompressional melting of a mixed pyroxenite–peridotite
 754 mantle source at MOR settings.

Fig. 10 Compositions (in wt%) of GV10 partial melts at 1.0 GPa from this study compared with those of experimental melts from variably fertile and depleted lherzolites within the temperature range 1200–1330 °C



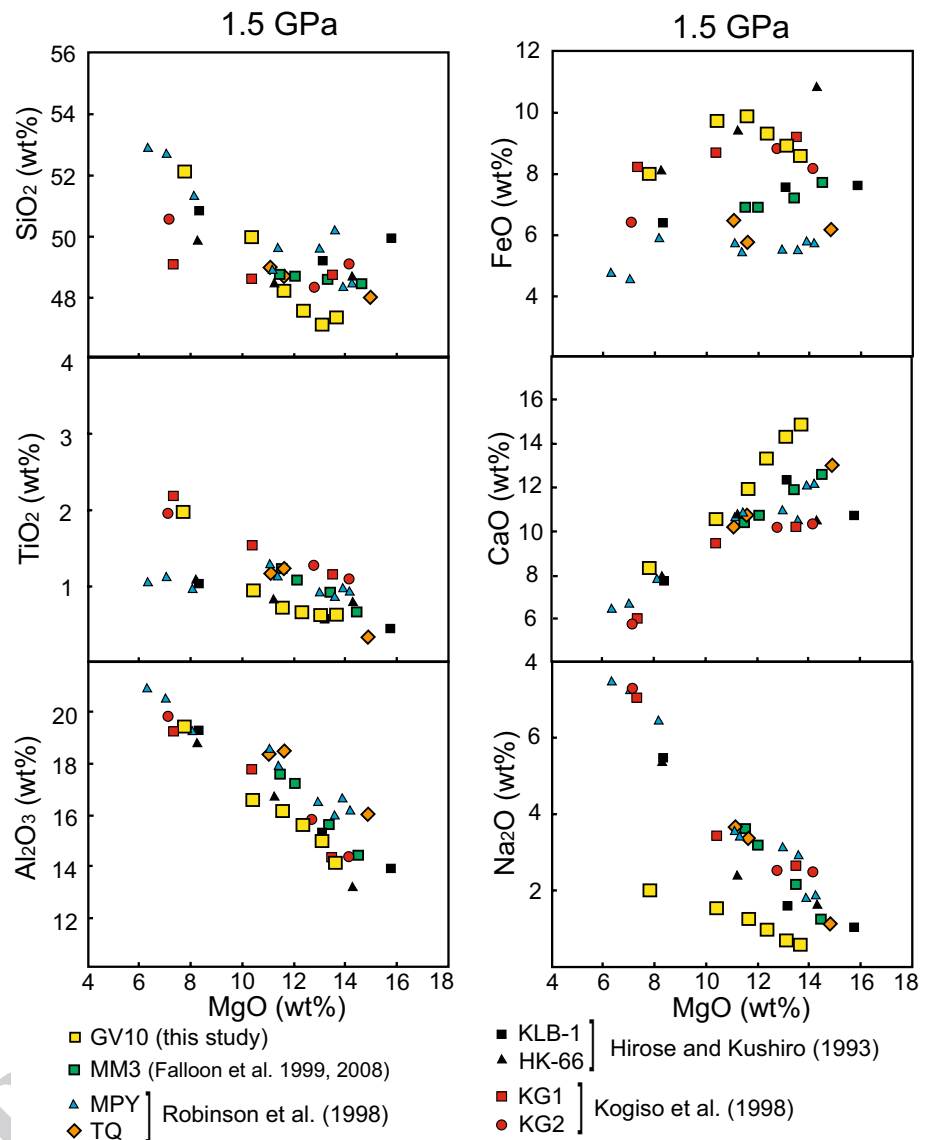
756 Secondary pyroxenites in heterogeneous upwelling 757 mantle

758 Experimental results on secondary pyroxenite GV10 can
759 be used to derive information on the composition of melts
760 produced by adiabatic upwelling and melting of heteroge-
761 neous mantle sources. In our modeling, we envisage that
762 mantle portions made by different proportions of pyrox-
763 enite GV10 and fertile lherzolite rise along mantle adi-
764 abats assuming a range of mantle potential temperatures
765 (T_p) consistent with mantle melting at MOR settings
766 ($T_p = 1260\text{--}1430$ °C; e.g., Langmuir et al. 1992; Shen and
767 Forsyth 1995; Green et al. 2001; Presnall et al. 2002; Her-
768 zberg et al. 2003). At this purpose, we applied the recent
769 parameterization of Lambart et al. (2016) that provides a
770 method to simulate the decompression melting of mixed

pyroxenite–peridotite mantle sources. Specifically, this
771 model predicts the melt productivity of the two melting
772 components at a given T_p of the rising mantle, as a func-
773 tion of bulk composition and mass fraction of pyroxenite
774 (Lambart et al. 2016), using a modified version of the ther-
775 modynamic treatment of adiabatic fractional melting by
776 Phipps Morgan (2001). 777

In the calculations, we assumed the parameteriza-
778 tion of the peridotite solidus temperature proposed by
779 Hirschmann (2000), which predicts slightly higher soli-
780 dus temperature than Katz's parameterization adopted by
781 Lambart et al. (2016), and is more consistent with experi-
782 mental results on fertile lherzolite (e.g., Falloon et al.
783 2008) (see Fig. 6). Using Lambart et al. (2016) modeling,
784 pyroxenite GV10 starts to melt at higher pressure than
785 peridotites, consistent with experimental results (Fig. 6). 786

Fig. 11 Compositions (in wt%) of GV10 partial melts at 1.5 GPa from this study compared with those of experimental melts from variably fertile and depleted lherzolites within the temperature range 1250–1400 °C



787 We considered two outermost mantle sources with
 788 pyroxenite mass fractions of 5 and 50%. Although the
 789 pyroxenite proportion into the mantle is commonly
 790 assumed to be around 2–10% (e.g., Hirschmann and
 791 Stolper 1996; Pertermann and Hirschmann 2003b), field
 792 evidence from orogenic and ophiolitic massifs indicates
 793 that the extent of pyroxenites in mantle sequences might
 794 be locally significantly higher than 10% (e.g., Gysi et al.
 795 2011). This is for example the case of some outcrops of the
 796 External Liguride mantle sequences, in which secondary
 797 pyroxenite abundance can reach a proportional ratio of 1:1
 798 to the associated peridotites (Borghini et al. 2013, 2016).

799 We first envisaged a mantle upwelling scenario along an
 800 adiabat with potential temperature (T_p) lower than 1310 °C.
 801 At such temperature and 1.5 GPa, pyroxenite GV10 gener-
 802 ates low to moderate melt amounts (up to 6.7 and 10.7%
 803 for pyroxenite mass fractions of 50 and 5%, respectively),

while lherzolite is at subsolidus or incipient melting condi-
 tions. Partial melts from low to moderate degrees of melt-
 ing of GV10 ($F < 12\%$) are Qz- or Hy-normative (Table 3).
 Therefore, they are expected to react with adjacent
 unmolten peridotite, producing orthopyroxene-rich reaction
 zone, as described for eclogite melts–peridotite interaction
 (Yaxley and Green 1998; Yaxley 2000; Sobolev et al. 2005,
 2007). As the ambient peridotite has zero or near-zero
 porosity and rather higher solidus temperature, these melts
 should freeze and crystallize at the pyroxenite–peridotite
 boundary. If pyroxenite-derived melts are able to percolate
 and impregnate the adjacent peridotite, they are expected to
 form orthopyroxene- and clinopyroxene-enriched perido-
 tites (e.g., Lambart et al. 2012). Similar peridotites, refer-
 tilized by interaction with pyroxenite-derived melts, have
 been documented in various natural case studies (e.g., Pear-
 son et al. 1993; Rivalenti et al. 1995; Varfalvy et al. 1996;

821 Zanetti et al. 1996; Mukasa and Shervais 1999; Bodinier
822 et al. 2004; Borghini et al. 2013).

823 Along the adiabatic upwelling of such heterogeneous
824 mantle, this metasomatic process is likely to be continu-
825 ous, as documented by experimental works on peridotite-
826 eclogite reaction (Yaxley and Green 1998; Rosenthal et al.
827 2014). Multiple events of reaction and freezing processes
828 introduce lithological zonation at the pyroxenite–perido-
829 tite boundary. In the case of reaction between silica-rich
830 partial melts and peridotite, the resulting orthopyroxene-
831 rich borders can act as permeability barrier that prevents
832 melt escaping from pyroxenite (e.g., Kogiso et al. 2004b;
833 Rosenthal et al. 2014). If infiltration of melts from pyrox-
834 enite into surrounding peridotite is inhibited, large batch
835 of pyroxenitic melts could be retained in equilibrium with
836 solid residues, isolated from further reaction with perido-
837 tite (Lambart et al. 2012). Mechanical instabilities caused
838 by melt accumulation can lead to rapid expulsion of these
839 melts into veins or channels (Kogiso et al. 2004b).

840 Extraction of melt fractions from pyroxenite GV10
841 would leave a wehrlitic residue made by a clinopyrox-
842 ene + olivine (\pm spinel) assemblage, introducing a range
843 of clinopyroxene-rich heterogeneities in the rising mantle.
844 Their bulk compositions, calculated using the composi-
845 tions of residual minerals in our experiments, are similar
846 to the bulk composition of clinopyroxenites investigated
847 by Kogiso and Hirschmann (2001) (Fig. 5). Such residual
848 components have Mg-rich minerals and low jadeite con-
849 tents in clinopyroxene, and their solidus temperature can be
850 significantly increased above that of the ambient peridotite
851 (e.g., Rosenthal et al. 2014; Lambart et al. 2016).

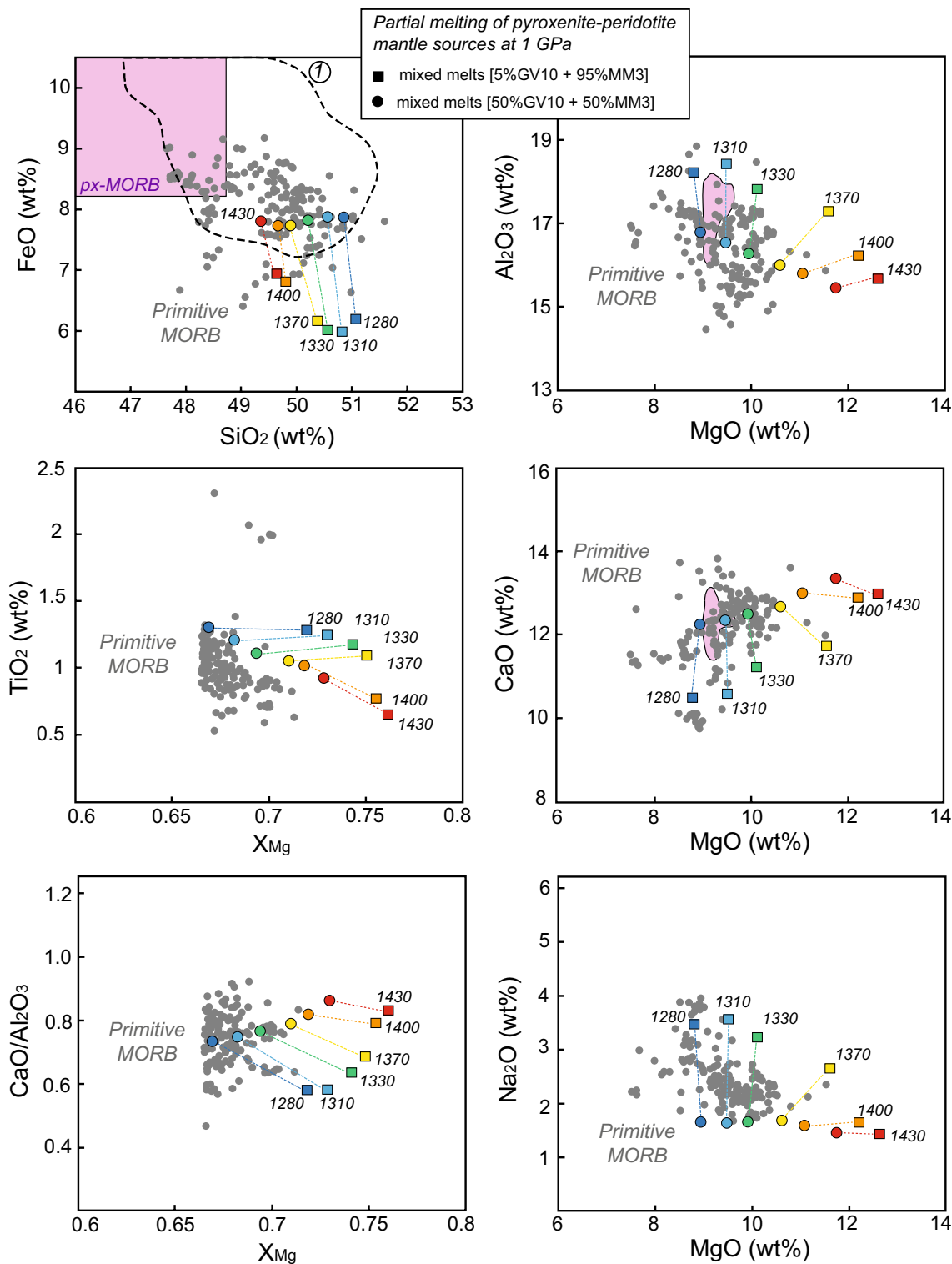
852 For mantle adiabats with potential temperature
853 (T_p) > 1310 °C, along the pressure interval 1–1.5 GPa,
854 pyroxenite and peridotite are melting simultaneously and
855 the melt productivity of the two components will depend
856 on their mass fractions in the source (e.g., Phipps Morgan
857 2001). To make some inferences on the contribution of sec-
858 ondary pyroxenite GV10 on major element chemistry of
859 basalts, we computed the compositions of melts produced
860 by mixed pyroxenite–peridotite sources at 1 GPa (Fig. 12).
861 We used the pyroxenite and peridotite “instantaneous” melt
862 fractions provided by the isentropic decompression melting
863 model of Lambart et al. (2016), assuming two pyroxenite/
864 peridotite mass fractions in the source as described above
865 (5 and 50% of GV10), and different T_p of the ascending
866 mantle (Suppl. Figure 3). A fundamental assumption of this
867 approach is that pyroxenite and peridotite behave as two
868 closed systems, i.e. in thermal equilibrium but chemically
869 isolated, until melt mixing and extraction, as assumed in
870 previous melting models (e.g., Lambart et al. 2009a; Short-
871 tle and MacLennan 2011; Sorbadere et al. 2013). This also
872 implies that partial melts remain in equilibrium with their
873 solid residues until extraction.

874 Major element compositions of melts used in the calcula-
875 tions were derived from experimental results on pyroxenite
876 GV10 (this study) and lherzolite MM3 (Baker and Stolper
877 1994; Falloon et al. 2008). This latter resembles the com-
878 position of External Liguride lherzolites hosting secondary
879 pyroxenite GV10, which were chemically and isotopically
880 identified as a potential MORB-type mantle source (Ram-
881 pone et al. 1995; Borghini et al. 2013). For each pyroxen-
882 ite–peridotite melt fraction pair, obtained at 1 GPa for dif-
883 ferent T_p using Lambart et al. (2016) modeling, we mixed
884 the composition of pyroxenite GV10 and lherzolite MM3
885 melts using the method proposed by Sorbadere et al. (2013)
886 (see details in caption of Fig. 12). In Fig. 12, the composi-
887 tions of mixed melts resulting from this model are plotted
888 together with the compositional field defined by the most
889 primitive MORBs. These latter were selected from a broad
890 database recently compiled by Gale et al. (2013), consider-
891 ing only glass compositions with $X_{Mg} \geq 0.67$, i.e. liquids
892 close to equilibrium with mantle olivine. This selection
893 reduced the original dataset of 12,527 compositions to only
894 184 glass compositions.

895 The dashed colored lines in Fig. 12 approximate the
896 range of composition of melts generated for different fixed
897 T_p by varying the mass fraction of secondary pyroxen-
898 ite in the source from 5 to 50%. The increase of second-
899 ary pyroxenite GV10 component in the source results in
900 mixed melts with higher FeO, CaO and lower Al_2O_3 , Na_2O
901 contents, whereas it only weakly influences the SiO_2 and
902 TiO_2 abundances (Fig. 12). SiO_2 and FeO contents plot far
903 from the composition of px-MORBs (Lambart et al. 2013),
904 and the CaO/ Al_2O_3 ratio of mixed melts still varies within
905 the range of MORB (CaO/ Al_2O_3 <0.9). MgO contents of
906 mixed melts remain rather constant for low T_p but it tends
907 to significantly decrease, at increasing pyroxenite mass
908 fraction, assuming $T_p > 1330$ °C (Fig. 12). As expected,
909 at fixed pyroxenite fraction in the source, MgO of mixed
910 melts is positively correlated with T_p , reflecting higher
911 melting degrees of both mantle components (peridotite and
912 secondary pyroxenite).

913 Melt compositions derived by the model match the vari-
914 ability defined by the most primitive MORB for most of
915 the major element oxides. Melt compositions best fitting
916 the compositional variability of MORBs are those calcula-
917 ted assuming T_p in the range 1280–1370 °C. The major
918 discrepancy is given by the higher X_{Mg} of mixed melts pro-
919 duced at higher T (>1370 °C), mostly from 5%-pyroxenite
920 source (Fig. 12), which more closely resemble the X_{Mg} val-
921 ues of peridotite-derived melts. As expected, this mismatch
922 tends to be higher at higher temperature (i.e. at higher melt-
923 ing degrees). However, the X_{Mg} of melts may be rapidly
924 homogenized to MORB values by early olivine fractiona-
925 tion (e.g., Langmuir et al. 1992), melt-rock reaction within
926 dunite channels (e.g., Kelemen et al. 1995; Asimow and

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927 Stolper 1999; Lambart et al. 2009b) and/or reactive crys-
 928 tallization (e.g., Collier and Kelemen 2010). Therefore, the
 929 results of our modeling do not allow discriminating the T_p
 930 of MORB generation.

931 To sum up, the results of this model support the postu-
 932 lation that secondary pyroxenite, as the one investigated in

this study, potentially represents a hidden component in the
 MORB mantle source (e.g., Shorttle and MacLennan 2011;
 Lambart et al. 2013; Rosenthal et al. 2014). A remarkable
 implication is that, if secondary pyroxenites inherited trace
 element and/or isotopic enriched signature during their for-
 mation by melt-rock reaction processes, their presence in

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Fig. 12 SiO₂ versus FeO; Al₂O₃, CaO and Na₂O versus MgO; CaO/Al₂O₃ and TiO₂ versus X_{Mg} of partial melts produced by decompression melting of two-lithologies mantle source at 1.0 GPa. Melt compositions resulting from the model are compared to the compositions of the most primitive MORB (with X_{Mg} ≥ 0.67; Gale et al. 2013). In SiO₂ versus FeO diagram, the field “1” (black dashed line) represent the compositional variability of MORB with MgO ≥ 9 wt%. Colored numbers refer to the potential Temperature (T_p) adopted in the isentropic decompression melting model (Lambart et al. 2016), to derive the “instantaneous” melt fractions produced simultaneously by secondary pyroxenite GV10 and a fertile lherzolite (see Supplementary Figure 3). Colored dashed lines represent the approximated trend of melt composition variability from mixed sources with 5 to 50% pyroxenite at a given T_p. For the fertile lherzolite we adopted the solidus temperature parameterization proposed by Hirschmann (2000) (see the text for further explanation). Calculations include two heterogeneous mantle sources made by 5% (secondary pyroxenite) GV10 + 95% (fertile lherzolite) MM3 and 50% GV10 + 50% MM3. We derive the mass fraction of pyroxenite-derived melt in aggregated liquids (LF_{py}) by the equation $LF_{py} = \Psi \times F_{py} / (\Psi \times F_{py} + (1 - \Psi) \times F_{per})$, where Ψ is the mass fraction of pyroxenite in the source, F_{py} and F_{per} are the melt fractions produced by pyroxenite and peridotite, respectively (Lambart et al. 2009a). The composition of aggregated “instantaneous” melts is provided by the equation: $X_{mix} = (X_{py} \times LF_{py}) + (X_{per} \times LF_{per})$, where X_{mix}, X_{py} and X_{per} are the weight percentages of a given element in the aggregated melt, pyroxenite melt and peridotite melt, respectively, and LF_{per} is the liquid fraction from the peridotite (1 - LF_{py}) (Sorbadere et al. 2013). Melt compositions at the given F have been derived from experiments on GV10 (this study) and MM3 (Falloon et al. 1999, 2008)

939 the mantle source could explain the decoupling between
940 “enriched” trace element/isotope and “normal” major ele-
941 ment compositions sometimes documented in oceanic
942 basalts (e.g., Salters and Dick 2002; le Roux et al. 2002;
943 Paulick et al. 2010; Zhang et al. 2012). In this context, man-
944 tle portions made by embedded secondary pyroxenite and
945 fertile lherzolite as documented in ultramafic massifs (e.g.,
946 Borghini et al. 2013), eventually delaminated and incorpo-
947 rated into the asthenosphere, could constitute the source of
948 trace element and isotopic enriched signature in MORBs.

949 Summary and concluding remarks

950 We have experimentally determined phase relations, melt-
951 ing conditions and melt and minerals compositions of an
952 anhydrous natural olivine-websterite (GV10), representing
953 a secondary-type pyroxenite, at 1 and 1.5 GPa and tem-
954 peratures from 1180 to 1400 °C. Residual mineral assem-
955 blages are dominated by clinopyroxene and minor olivine
956 (and spinel). Relatively high X_{Mg} and low alkali content
957 of our starting composition result in solidus temperature
958 only slightly lower than fertile lherzolite. At both pressure
959 conditions, higher modal orthopyroxene relative to clino-
960 pyroxene is involved in the melting reaction, thus resulting
961 in the early disappearance of orthopyroxene, and high SiO₂
962 contents in melts produced by low- to moderate-melting

degree (F < 13%). Although for many oxides the composi- 963
tions of partial melts from secondary pyroxenite GV10 are 964
quite similar to those of melts produced by variably fertile 965
lherzolites, they differ for higher CaO contents, resulting in 966
slightly higher CaO/Al₂O₃ ratios, and very low Na₂O abun- 967
dances. These latter reflect relatively high partition coef- 968
ficient of Na between residual clinopyroxene and partial 969
melt, together with the low initial bulk content. D_{Na₂O}^{cpx/liquid} 970
values obtained in this study, together with other experi- 971
mental, define a positive correlation with pressure and 972
temperature. 973

Profiting by the recent parameterization of Lambart 974
et al. (2016), we have simulated the partial melting of 975
mixed pyroxenite–peridotite sources in the context of adia- 976
batic mantle upwelling. At relatively low potential tempera- 977
tures (T_p < 1310 °C), low fractions of secondary pyroxenite 978
melts are expected to react and freeze within the subsolidus 979
adjacent peridotite, resulting in orthopyroxene-rich reac- 980
tion zones (or refertilized peridotite) and high-X_{Mg} clino- 981
pyroxene-rich residues. At higher T_p, secondary pyroxen- 982
ite and peridotite melt simultaneously. Melt compositions 983
computed by varying the pyroxenite mass fractions in the 984
source (5–50%) largely encompass the chemical variability 985
defined by most primitive MORBs (mostly in the T_p range 986
1280–1370 °C), supporting the idea that secondary pyrox- 987
enites potentially represent a hidden component in MOR 988
mantle sources. Therefore, fertile mantle peridotite veined 989
by secondary pyroxenite may act as a carrier of trace ele- 990
ment and isotopic enriched signature in MORBs. 991

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