

# A New Stable Boundary Layer Parameterization for Weather and Forecasting Models: a Heat Flux Budget Approach

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## Research Article

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1 **A New Stable Boundary Layer Parameterization**  
2 **for Weather and Forecasting Models: a Heat Flux**  
3 **Budget Approach**

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10 **Abstract** The present study introduces a new boundary layer parameteriza-  
11 tion for weather and forecasting models. It is implemented here as a boundary  
12 layer module in Weather Research and Forecasting (WRF) model. The main  
13 novelty in the new scheme is that it includes prognostic equations for the heat  
14 flux and temperature variance, being the first WRF boundary layer scheme  
15 with that feature. This is specially aimed at improving the representation of  
16 nocturnal stable boundary layer and of its turbulence regimes, weakly and  
17 very stable. The effort is supported by previous studies that found that the  
18 two regimes and the transitions between them are better represented by sim-  
19 plified numerical schemes that represent the interactions between the surface  
20 and the air adjacent to it when the heat flux and temperature variance are  
21 solved prognostically. The results show that the two regimes are adequately  
22 simulated by the new scheme. Such an evaluation is presented in terms of the  
23 relationship between the turbulence velocity scale and mean wind speed, of  
24 the dependence of the potential temperature gradient near the surface and the  
25 mean wind speed, and by the relationship between flux and gradient Richard-  
26 son numbers. In the new scheme, the relationship between thermal structure  
27 and the mean and turbulent flows arises naturally from the heat flux prognostic  
28 equation, not being arbitrarily imposed by an empirical stability function.

29 **Keywords** Eddy diffusivity · Stable boundary layer · Turbulence  
30 parametrizations · Turbulence regime · Weather Research and Forecast-  
31 ing model

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## 1 Introduction

Through mechanical drag and thermal effects, the Earth’s surface directly affects the flow in the atmospheric boundary layer (ABL), making it turbulent (Stull 1988; Cuxart et al. 2006). As a consequence, forecasting the mean state of the ABL demands an adequate representation of the effects of turbulence, which is usually described in turbulent models through the statistical moments of the fluctuations of the variables that describe the atmospheric flow. Furthermore, as the majority of human activities take place in the ABL, such flow is often affected by anthropogenic action. It is clear that a good representation of the ABL and its interaction with the surface is essential for all atmospheric applications.

The adequate modelling of the ABL and of its importance as a lower boundary to upper atmospheric flow has been pursued by the micrometeorological community for a long time (Taylor 1915; Blackadar 1962; Taylor and Delage 1971; Mellor and Yamada 1974; Wyngaard 1975; Louis 1979, among others). Naturally, the quality of such model representations is limited by the same difficulties that affect the bulk of knowledge of the ABL at any given historic context. Currently, a major challenge for this understanding regards the stable boundary layer (SBL), its turbulence regimes and how they relate to the mean flow and to quantities that are external to the SBL (van Hooijdonk et al. 2015; Vignon et al. 2017; van der Linden et al. 2017; Holdsworth and Monahan 2019). It has been studied by numerous authors in the last decades (Mahrt 1998; Sun et al. 2012, 2016; Acevedo et al. 2014; Mahrt 2014, among others), having been classified both observationally and in modelling in two distinct regimes: weakly stable and very stable. According to Mahrt (2014) the “*fundamental features of the very stable boundary layer still remain a mystery*”. In this context, the simulation of the very SBL and the representation of transitions from and to such a regime by planetary boundary layer parameterizations is a major micrometeorological challenge that limits the quality of the representation of the mean state of the atmosphere near the surface in both numerical weather prediction (NWP) and climate models today {(van de Wiel et al. 2017; Battisti et al. 2017; Baas et al. 2019; Holdsworth and Monahan 2019; Lapo et al. 2019; Maroneze et al. 2019, 2021; Lorenz et al. 2022; Kähnert et al. 2022). At the same time, the weakly SBL is well described by Monin-Obukhov similarity theory and, for this reason, is comparatively well simulated and represented by numerical planetary boundary layer (PBL) parametrization schemes (Mahrt 2014).

The Weather Research and Forecasting (WRF) model (Skamarock et al. 2008) is largely employed to investigate hydrology, renewable energy, regional climate, weather prediction, and other phenomena (Powers et al. 2017), both for scientific research and for operational numerical prediction. Maroneze et al. (2021) have analyzed the quality of the representation of turbulence in very stable conditions by the different WRF PBL parameterizations that solve the turbulence kinetic energy, either diagnostically or prognostically. They found that the variety of existing parameterizations lead to a large diversity in rep-

78 representations between turbulent and mean quantities, with different simulated  
 79 thresholds between the two regimes and consequent very different predictions  
 80 of mean quantities in specific situations. None of the boundary-layer param-  
 81 eterizations available in WRF solves a prognostic equation for the heat flux,  
 82 although there are indications in the literature that this variable has a decisive  
 83 control in the the SBL turbulence regime (van de Wiel et al. 2012; Maroneze  
 84 et al. 2019). Maroneze et al. (2019) have shown that the adequate simulation  
 85 of the transition between the two regimes is much better simulated when a  
 86 prognostic for the heat flux is solved by a prognostic equation, rather than  
 87 represented by a parameterization.

88 In NWP models, the influence of flow stability to the ABL flow is usually  
 89 prescribed as an empirical and arbitrary stability function (Costa et al. 2020).  
 90 Here, it is hypothesized that the model flow dependence on stratification, and,  
 91 more generally, the relationship between mean and turbulence quantities (such  
 92 as the flux and gradient Richardson numbers) arises naturally in the model  
 93 when the heat flux is prognostically solved, instead of being prescribed by a  
 94 stability function. Therefore, the main goal of the present study is to intro-  
 95 duce a new stable boundary layer parameterization for weather and forecasting  
 96 models that includes prognostic equations for the heat flux and temperature  
 97 variance. Such a scheme is implemented in the WRF model version 3.9, de-  
 98 scribed in section 2. The proposed parameterization is calibrated and validated  
 99 using GEWEX Atmospheric Boundary Layer Study (GABLS) I (Kosović and  
 100 Curry 2000) as a control case, in section 3.2. The ability of the new parame-  
 101 terization to represent the two contrasting SBL regimes is evaluated in section  
 102 4, through the relationships it simulates between the turbulence velocity scale  
 103 and the mean wind speed, between the potential temperature gradient near the  
 104 ground surface and the mean wind speed, and between the flux and gradient  
 105 Richardson numbers.

## 106 2 Description of the Planetary-Boundary-Layer Parameterization

107 Turbulence affects the mean state of the planetary boundary layer through the  
 108 divergence of turbulent fluxes. Such a divergence in the vertical direction is  
 109 often the dominant process affecting mean quantities near the surface. When  
 110 its role is sufficiently larger than that of other influences, the prognostic equa-  
 111 tions for mean horizontal velocity components and potential temperature may  
 112 be regarded as

$$\left[ \frac{\partial \bar{u}}{\partial t} \right]_{PBL} = -\frac{\partial \overline{w'w'}}{\partial z}, \quad (1)$$

$$\left[ \frac{\partial \bar{v}}{\partial t} \right]_{PBL} = -\frac{\partial \overline{w'v'}}{\partial z}, \quad (2)$$

$$\left[ \frac{\partial \bar{\theta}}{\partial t} \right]_{PBL} = -\frac{\partial \overline{w'\theta'}}{\partial z}. \quad (3)$$

113 In equations (1-3)  $\bar{u}$ ,  $\bar{v}$  and  $\bar{\theta}$  are respectively the mean velocity components  
 114 in zonal and meridional directions, and the mean potential temperature. The  
 115 vertical turbulent momentum fluxes are given by  $\overline{u'w'}$  and  $\overline{v'w'}$ , while  $\overline{w'\theta'}$  is  
 116 the turbulent heat flux.

117 According to K theory, the vertical turbulent momentum fluxes are parametrized  
 118 as (Mellor and Yamada 1974, 1982; Therry and Lacarrere 1983; Bougeault and  
 119 Lacarrere 1989; Janjić 1994; Nakanishi and Niino 2009; Bretherton and Park  
 120 2009, among others)

$$\overline{u'w'} = -K_m \frac{\partial \bar{u}}{\partial z}, \quad (4)$$

121 and

$$\overline{v'w'} = -K_m \frac{\partial \bar{v}}{\partial z}, \quad (5)$$

122 where  $K_m$  is the vertical turbulent diffusion coefficients for momentum. In  
 123 turbulence closures that solve the TKE either prognostically or diagnostically,  
 124 it is common to estimate  $K_m$  as a function of TKE, the mixing length  $l_m$ , and  
 125 a stability function  $f_m$  (Mellor and Yamada 1974, 1982; Nakanishi and Niino  
 126 2009; Bretherton and Park 2009; Costa et al. 2020, among others)

$$K_m = \sqrt{\bar{e}} l_m f_m. \quad (6)$$

127 For a horizontally homogeneous atmosphere, the prognostic equation for  
 128 TKE ( $\bar{e}$ ) is

$$\frac{\partial \bar{e}}{\partial t} = -\overline{u'w'} \frac{\partial \bar{u}}{\partial z} - \overline{v'w'} \frac{\partial \bar{v}}{\partial z} + \frac{g}{\Theta} \overline{w'\theta'} - \frac{\partial}{\partial z} \left[ \overline{w'e} + \frac{\overline{p'w'}}{\rho_0} \right] - \epsilon_e, \quad (7)$$

129 where  $p$  is pressure,  $\rho_0$  is a reference density and  $\Theta$  is a reference temperature.  
 130 In the right hand side (r.h.s) of Eq. 7, the first and second terms represent the  
 131 turbulence shear production (SP); the third term is the turbulence buoyant de-  
 132 struction (production) under stable (unstable) conditions (BD/P); the fourth  
 133 term is the vertical transport of TKE both by turbulence and by pressure  
 134 fluctuations (TR); and the fifth term is TKE viscous dissipation (DIS).

135 The TKE viscous dissipation ( $\epsilon_e$ ) can be parameterized as a function of  
 136 TKE and a characteristic dissipation length  $l_\epsilon$ :

$$\epsilon_e = c_1 \frac{\bar{e}^{3/2}}{l_\epsilon}, \quad (8)$$

137 where  $c_1$  is a numerical constant. According to Cuxart et al. (2006), values  
 138 from 0.08 to 0.7 have been used for  $c_1$ .

139 Following Dwyntkerke (1988), the TKE transport term is represented as:

$$-\frac{\partial}{\partial z} \left[ \overline{w'e} + \frac{\overline{p'w'}}{\rho_0} \right] = \frac{\partial}{\partial z} \left[ K_e \frac{\partial \bar{e}}{\partial z} \right], \quad (9)$$

140 here  $K_e = \alpha_e K_M$ .

141 The prognostic equation for the heat flux ( $\overline{w'\theta'}$ ) is:

$$\frac{\partial \overline{w'\theta'}}{\partial t} = -\overline{w'^2} \frac{\partial \overline{\theta}}{\partial z} + \frac{g}{\Theta} \overline{\theta'^2} - \frac{\partial \overline{w'w'\theta'}}{\partial z} + \frac{1}{\rho_0} \overline{\theta'} \frac{\partial p'}{\partial z}, \quad (10)$$

142 where the first term on the r.h.s represents the thermal gradient production of  
 143 heat flux in a SBL (TGP). The second term is heat flux buoyant destruction  
 144 (production) under stable (unstable) condition, and the third term represents  
 145 the transport of heat flux by turbulence, while the fourth term either one or  
 146 the other both the transport by pressure fluctuations and return-to-isotropy.

147 Following Therry and Lacarrere (1983) the last term of Eq. 10 is param-  
 148 eterized according to the idea of a pressure relaxation, as the sum of two  
 149 contributions. The first is proportional to the heat flux ( $P1$ ) itself and the  
 150 second is proportional to the temperature variance ( $P2$ )

$$\frac{1}{\rho_0} \overline{\theta'} \frac{\partial p'}{\partial z} = P1 + P2, \quad (11)$$

151 where

$$P1 = -c_2 \frac{\overline{e}^{1/2}}{l_\epsilon} \overline{w'\theta'} \quad \text{and} \quad P2 = -c_3 \frac{g}{\Theta} \overline{\theta'^2}, \quad (12)$$

152 and  $c_2$  and  $c_3$  are numerical constants.

153 For temperature variance, the prognostic equation is:

$$\frac{\partial \overline{\theta'^2}}{\partial t} = -2\overline{w'\theta'} \frac{\partial \overline{\theta}}{\partial z} - \frac{\partial \overline{w'\theta'^2}}{\partial z} - \epsilon_\theta, \quad (13)$$

154 where the first term in the r.h.s is the heat flux production of temperature  
 155 variance (Pr), the second is its turbulent transport (TR) and the third term  
 156 is its molecular dissipation (DIS).

157 The dissipation term of temperature variance is parametrized as

$$\epsilon_\theta = c_4 \frac{\sqrt{\overline{e}}}{l_\epsilon} \overline{\theta'^2}, \quad (14)$$

158 where  $c_4$  is numerical constants.

159 In analogy with Eq. 9, the transport term in Eqs. 10 and 13 are parametrized  
 160 as

$$-\frac{\partial \overline{w'w'\theta'}}{\partial z} = \frac{\partial}{\partial z} \left[ K_e \frac{\partial \overline{w'\theta'}}{\partial z} \right] \quad \text{and} \quad -\frac{\partial \overline{w'\theta'^2}}{\partial z} = \frac{\partial}{\partial z} \left[ K_e \frac{\partial \overline{\theta'^2}}{\partial z} \right].$$

161 To close the above set of equations, the numerical constants, the mixing  
 162 length  $l_m$  and the characteristic dissipation length  $l_\epsilon$  must be specified. In the  
 163 literature, the two mixing lengths are generally taken as equal to each other  
 164 (Therry and Lacarrere 1983; Weng and Taylor 2003; Costa et al. 2020, among  
 165 other) and many different formulations and numerical constants have already  
 166 been proposed.

167 WRF uses an Arakawa C grid, where the center of the grid cell (represented  
 168 by symbol  $\times$  in fig. 1) are referred as “mass points” or full levels, while the face  
 169 grid points staggered at one-half grid length from the mass points are referred  
 170 as half levels (represented by symbol  $\circ$  in fig. 1). Here, the prognostic equa-  
 171 tions for any turbulent variables are calculated at the full levels, in contrast  
 172 with other parameterizations present in WRF. Therefore, here the turbulent  
 173 flux divergences can be estimated directly through centered finite differences,  
 174 not being necessary the use of spatially averaged values (Fig. 1).

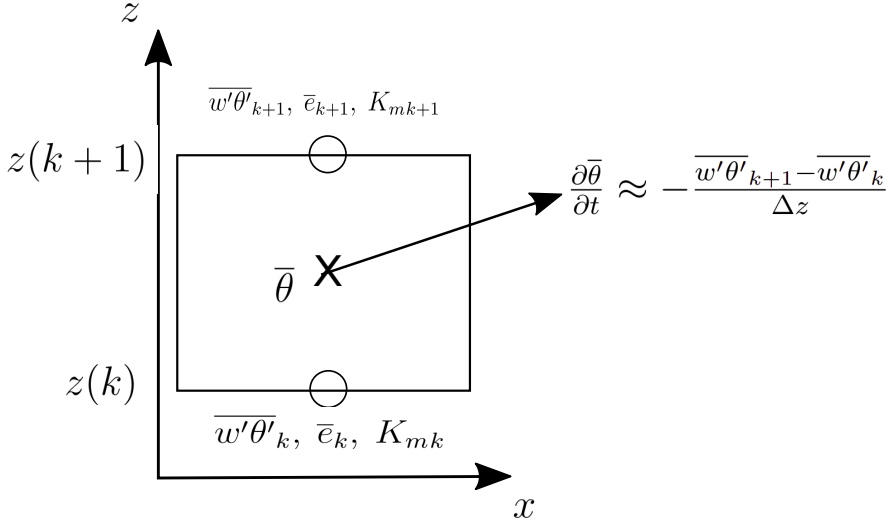
175 Equations (7),(10), and (13) are solved through an implicit time-integration  
 176 method. They can be generically discretized as

$$\underbrace{\frac{\Psi_k^{n+1} - \Psi_k^n}{\Delta t}}_{\text{tendency}} = \underbrace{\frac{\partial}{\partial z} \left[ K_m^n \frac{\partial \Psi^{n+1}}{\partial z} \right]_k}_{\text{Transport}} + \underbrace{P_k^n}_{\text{Production}} - \underbrace{F_k^n \Psi_k^{n+1}}_{\text{Destruction}}, \quad (15)$$

177 where  $\Psi$  is the turbulent variable,  $n$  denotes the time index and  $k$  denotes the  
 178 vertical full level index. Equation (15) is integrated in time by the relationship

$$a \Psi_{k-1}^{n+1} + b \Psi_k^{n+1} + c \Psi_{k+1}^{n+1} = d \quad (16)$$

179 where  $a$ ,  $b$ , and  $c$  are the elements of the matrix that solves the implicit system  
 for  $\Psi_{k-1}^{n+1}$ .



**Fig. 1** The schematic representation of vertical adapted Arakawa C grid for the present PBL parametrization.

### 3 Model Validation

#### 3.1 Control Case and model discretization

PBL schemes are typically calibrated through a control case. One of the most widely used reference cases for such a purpose is the first Global Energy and Water cycle EXperiment (GEWEX) Atmospheric Boundary Layer Study (GABLS I). GABLS I provides a model intercomparison for SCM (Cuxart et al. 2006) and LES (Beare et al. 2006) for a weakly stable boundary layer, with prescribed temperature at the surface (Kosović and Curry 2000). Over the years, GABLS I has been used to validate schemes designed for different purposes, such as the development of new parametrizations (Bretherton and Park 2009; Cheng et al. 2020), SBL regimes regime transitions studies (Costa et al. 2020).

The boundary and initial conditions used in all simulations described in this section are the same as those in GABLS I (Kosović and Curry 2000; Cuxart et al. 2006):

- Constant geostrophic wind components ( $u_G = 8 \text{ m s}^{-1}$  and  $v_G = 0 \text{ m s}^{-1}$ ) at the domain top, along the entire simulation;
- Constant surface cooling rate ( $0.25 \text{ K h}^{-1}$ ) along the simulation;
- The initial profiles of wind components, temperature and turbulence kinetic energy are:  $\bar{u}(z, t = 0) = u_G$ ;  $\bar{v}(z, t = 0) = v_G$ ;  $\bar{\theta}(z < 100 \text{ m}, t = 0) = 263.5 \text{ K}$ , while a constant lapse rate of  $0.1 \text{ K m}^{-1}$  is considered at heights  $z > 100 \text{ m}$ ;  $\bar{e} = 0.4(1 - z/250)^3$ .

WRF single column mode (WRF-SCM), with 170 levels between the surface ( $z = 0$ ) and the domain top ( $z = 6 \text{ km}$ ), is used in the PBL scheme validation. The first atmospheric level is fixed at  $z = 1.5 \text{ m}$ , and the grid spacing increases steadily from  $1.6 \text{ m}$  near the surface, to  $198 \text{ m}$  near the domain top.

#### 3.2 Model calibration

Typically, PBL schemes implemented in weather and climate models have a large number of tuning parameters. The number of numerical constants is proportional to the number of parametrizations necessary to close the system of equations. According to Audouin et al. (2021), the model calibration consists in adjusting each tuning parameter, by taking into account both the model performance and physical restrictions. It can be a difficult task because of the large number of degrees of freedom which demand a high computational cost (Audouin et al. 2021).

##### 3.2.1 Stable Boundary Layer Mixing Length Formulation

An adequate mixing length formulation is crucial to properly describe the boundary layer flow (Weng and Taylor 2003), having a very important role



in any PBL scheme. In the literature, many different formulations have been proposed. Blackadar (1962) suggested

$$\frac{1}{l_{Blackadar}} = \frac{1}{\kappa z} + \frac{1}{\lambda_0}, \quad (17)$$

where  $l_{Blackadar}$  is a mixing length,  $\kappa$  is von Karman constant, and  $\lambda_0$  is a reference length scale. Table 1 show different formulations for  $\lambda_0$  proposed in the literature.

The mixing length tends to be smaller under stable thermal stratification (André et al. 1978), and can be given by the harmonic average between  $l_{Blackadar}$  and a buoyant length scale  $L_B$  (Sukoriansky et al. 2005; Nakanishi and Niino 2009, among other)

$$\frac{1}{l_m} = \frac{1}{l_{Blackadar}} + \frac{1}{L_B}, \quad (18)$$

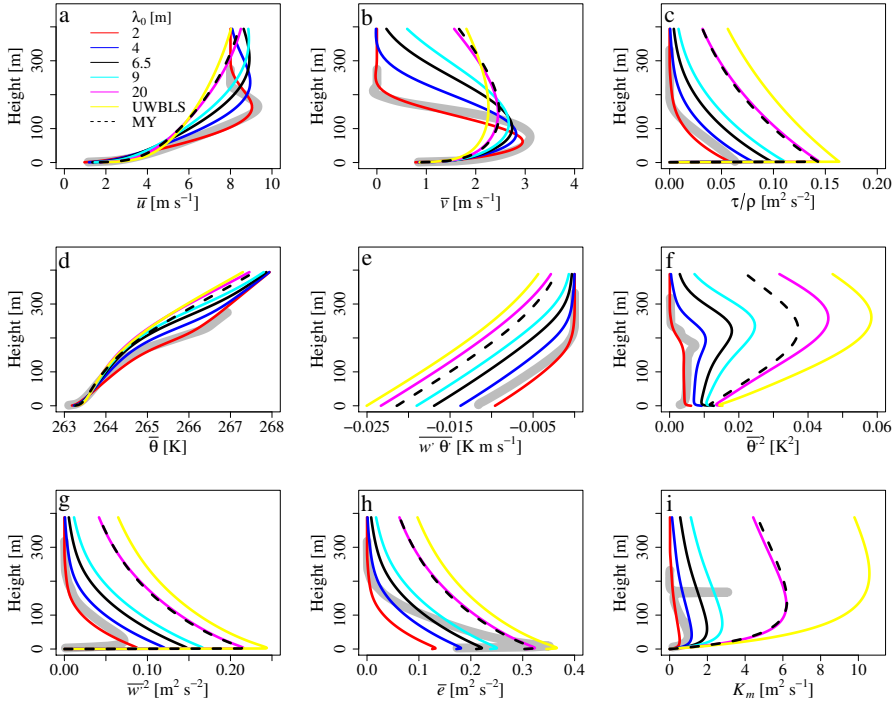
where

$$L_B = C_N \frac{\sqrt{2} \bar{\epsilon}}{N}, \quad (19)$$

where  $N$  is the Brunt–Väisälä frequency and  $C_N$  is a numerical constant whose used values range from 0.2 to 1 (André et al. 1978; Baas et al. 2008; Nakanishi and Niino 2009).

The role of the reference length scale can be evaluated by considering  $L_B \rightarrow \infty$  in Eq. (18) and varying  $\lambda_0$ . Fig. 2 shows that the mean vertical profiles of the different atmospheric quantities vary largely with  $\lambda_0$ . For example, the height and width of the near surface maxima of the wind-velocity profile increases linearly with  $\lambda_0$  (Fig. 2a,b). The absolute values of momentum flux ( $\tau/\rho = \sqrt{\overline{u'w'^2} + \overline{v'w'^2}}$ ), potential temperature ( $\bar{\theta}$ ), heat flux ( $\overline{w'\theta'}$ ), temperature variance ( $\overline{\theta'^2}$ ), vertical wind velocity variance component ( $\overline{w'^2}$ ), and TKE ( $\bar{\epsilon}$ ) and the momentum eddy viscosity ( $K_m$ ) generally increase at all heights as  $\lambda_0$  increases. When compared to the GABLS1 observations, the simulated TKE is underestimated for most of the values of  $\lambda_0$  considered. In general, the comparison indicates that when  $\lambda_0 = 2$  m, the model outputs are closest to the GABLS1 reference case (Fig. 2 c-i). However, the intensity of the wind-speed maximum near the surface (low-level-jet nose, Klein et al. 2016) is underestimated. For larger values of  $\lambda_0$ , the temperature variance increases with height until it reaches a maximum near the SBL top (Fig. 2f). When  $\lambda_0$  is evaluated by formulations such as those proposed by Mellor and Yamada (1974) (MY, black dashed lines in Fig. 2) and Bretherton and Park (2009) (UWBLS, yellow lines in Fig. 2) the mean vertical profiles of the atmospheric quantities approach their values when large values of  $\lambda_0$  are considered.

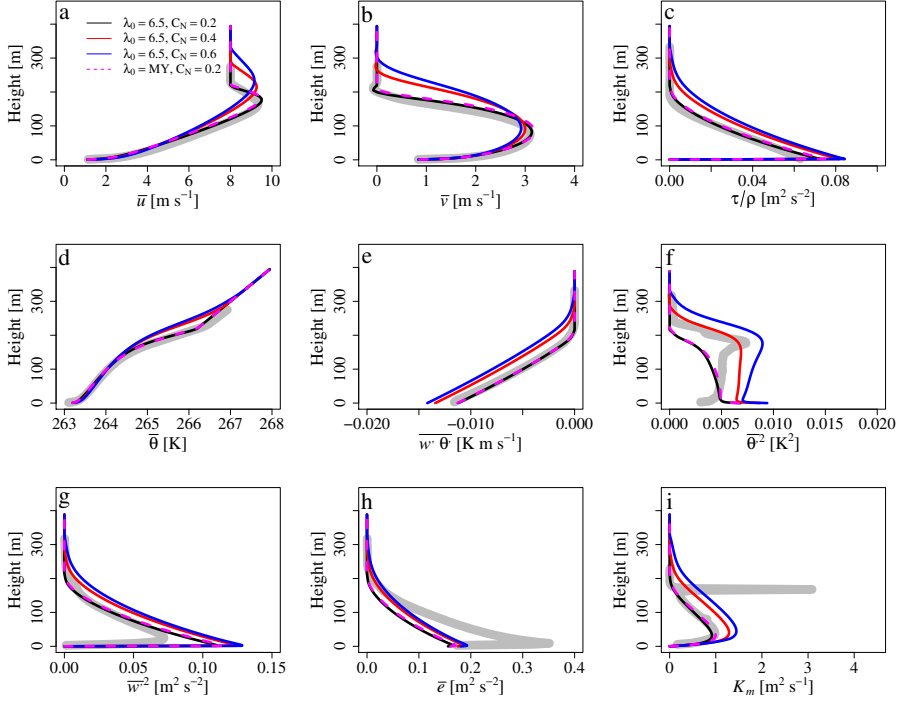
Finite values of the buoyant length scale, in Eq. (18), make the wind-speed maximum near the surface sharper and more intense. Moreover, it reduces the turbulence intensity promoting a shallower SBL (not shown). The value chosen for  $C_N$  is important because the buoyant length scale plays a similar



**Fig. 2** Vertical profiles, mean values taken over last two hours from the simulation of zonal and meridional velocity components  $\bar{u}$  (a) and  $\bar{v}$  (b), the absolute values of momentum flux  $\tau/\rho$  (c), potential temperature  $\theta$  (d), kinematic heat flux  $w'\theta'$  (e), temperature variance  $\theta'^2$  (f), vertical wind velocity variance component  $\overline{w'^2}$  (g), TKE (h), and the momentum eddy viscosity  $K_m$  (i) for different reference length scales. The solid thick gray line is the GABLS I case.

256 role as a stability function does in the mixing length formulation under stable  
 257 conditions (Fig. 3). When smaller values of  $C_N$  are considered, the profiles  
 258 of the wind speed components present a sharper and more intense maximum  
 259 than when  $C_N$  is larger (Figs. 3a,b). On the other hand, in those cases the  
 260 momentum eddy diffusivity is smaller than in the cases with large  $C_N$  (Fig.  
 261 3i). This occurs because the buoyant length scale will decrease, for small  $C_N$ ,  
 262 and then, it will reduce the mixing length intensity and the diffusion coeffi-  
 263 cients, consequently. When large values of  $C_N$  are assumed the momentum  
 264 flux, the heat flux, the temperature variance and TKE become more intense  
 265 (Fig. 3c,e,f,h), determining a more turbulent and deeper SBL. Even though the  
 266 turbulence intensity increases with  $C_N$ , the temperature vertical profile shows  
 267 that the SBL will become colder near the the boundary layer top because the  
 268 enhanced turbulent transport allows the cold air reaching higher levels (Fig.  
 269 3d). In almost all comparisons between the profiles presented in Fig. 3, one  
 270 can verify that the model profiles are closer to the control case when the pa-  
 271 rameters  $\lambda_0 = 6.5$  m and  $C_N = 0.2$  are used. Therefore, in order to keep the

272 parametrization as simple and as accurate as possible, such optimized values  
 273 of  $\lambda_0$  and  $C_N$  are used in the following analysis.



**Fig. 3** Same of Fig. 2 considering a finite values of the buoyant length scale and different values of  $C_N$ .

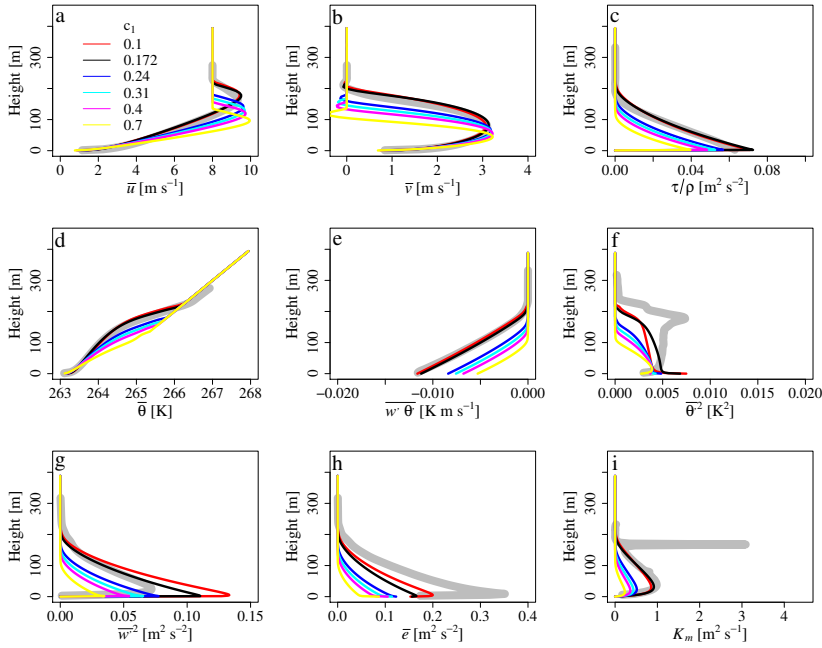
**Table 1** Reference length scale

Reference	$\lambda_0$ [m]
Blackadar (1979)	$0.00027 \frac{U_G}{f}$
Bretherton and Park (2009)	$0.085 PBLH$
Mellor and Yamada (1974)	$0.1 \frac{\int_0^\infty  z  q dz}{\int_0^\infty q dz}$
Thery and Lacarrere (1983)	50
QNSE	$0.0063 \frac{u_*}{f}$

### 274 3.2.2 Second order constants dependence

275 As mentioned in section 2, the value of  $c_1$  in the parameterization of the viscous  
 276 dissipation (Eq. 8), varies by almost an order of magnitude among existing

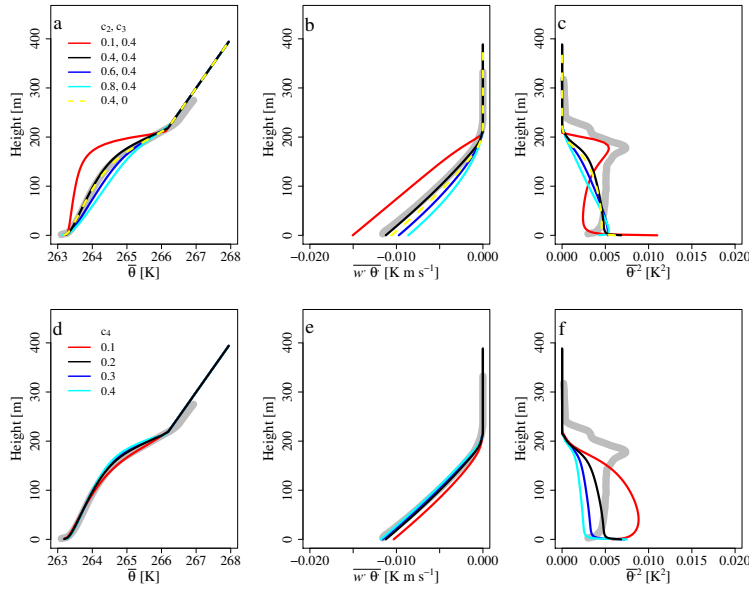
277 PBL schemes. Fig. 4 shows how the variation of  $c_1$  impacts the model. All  
 278 profiles show that the SBL height decreases appreciably as  $c_1$  increases (Fig.  
 279 4). When larger values of  $c_1$  are considered, the magnitudes of momentum flux,  
 280 heat flux, temperature variance, TKE, and momentum eddy diffusivity become  
 281 smaller and a shallower SBL is simulated (Fig. 4c,e,f,h,i). Increasing  $c_1$  leads  
 282 to larger TKE dissipation, reducing the turbulent quantities and causing the  
 283 SBL to be more stratified. Moreover, the turbulent flux divergence also varies,  
 284 affecting the vertical potential temperature profile, which is nearly linear in  
 285 the SBL (Fig. 4d) for large  $c_1$  values, while it is curved, which a smaller  
 286 stratification at lower levels when  $c_1$  is smaller.



**Fig. 4** Same of Fig. 2 for different values of  $c_1$  in the parametrization of the viscous dissipation.

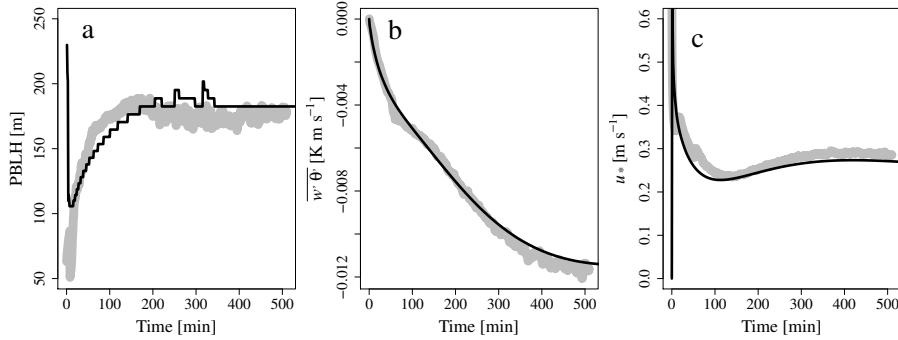
287 While  $c_1$  values impact the SBL depth, the choice of parameters  $c_2$  and  $c_3$   
 288 (Eq. 12) does not have the same effect. On the other hand, varying  $c_2$  affects  
 289 the vertical profiles of temperature (Fig. 5a), heat flux (Fig. 5b) and of the  
 290 mean temperature variance (Fig. 5c). The constants  $c_2$  and  $c_3$  are, respec-  
 291 tively, control parameters in the parameterization of the transport by pressure  
 292 fluctuations and in the return-to-isotropy terms in the heat flux budget (Eq.  
 293 10). In particular,  $c_2$  is a coefficient in the parameterization that mimics a heat  
 294 flux dissipation rate ( $P_1$ ). Therefore, larger values of  $c_2$  cause a decrease on  
 295 the simulated heat flux magnitude (Fig. 5b), consequently reducing variance  
 296 as well (Fig. 5c). Furthermore, smaller values of  $c_2$  imply larger absolute heat,

297 allowing larger turbulent heat transport at levels near the SBL top, which  
 298 becomes colder (read line in Fig. 5a), although not affecting the SBL thick-  
 299 ness. In such case, the stability of the entire SBL is reduced, and it becomes a  
 300 near-neutral layer from the ground until near the top, where a large thermal  
 301 inversion is present. As the temperature variance production is proportional  
 302 to both the heat flux and the thermal gradient, in this case the temperature  
 303 variance is small for most of the SBL, reaching a local maximum near the max-  
 304 imum of the thermal gradient, at the SBL top (red line in Fig. 5c). Fig. 5 also  
 305 shows that the SBL height is not affected by the choice of free coefficient in  
 306 the temperature variance dissipation rate ( $c_4$ ). A 4-fold variation in  $c_4$  causes  
 307 the surface heat flux to vary by less than 11 % (Fig. 5e).



**Fig. 5** Vertical profiles, mean values taken over last two hours from the simulation, for potential temperature  $\bar{\theta}$  (a, d), kinematic heat flux  $\overline{w'\theta'}$  (b, e), temperature variance  $\overline{\theta'^2}$  (c, f) for different values of  $c_2$ ,  $c_3$  and  $c_4$ . The solid thick gray line is the GABLS I case.

308 Based on these results, the free parameters of the model have been chosen  
 309 to optimize the model performance, when the outputs are compared with the  
 310 values from GABLS I case (Table 2). Such a comparison is shown in Fig. 6  
 311 for PBL height (Fig. 6a), surface sensible heat flux (Fig. 6b) and the surface  
 312 friction velocity (Fig. 6c). Following Cuxart et al. (2006), in this analysis the  
 313 boundary-layer height is the level where the sensible heat flux decays to 5%  
 314 of its surface value.



**Fig. 6** PBL height (a), heat flux (b), and friction velocity (c) temporal evolution. The solid thick gray points is the GABLS I case.

**Table 2** Free Parameter

$c_1$	0.172
$c_2$	0.4
$c_3$	0.4
$c_4$	0.2
$c_N$	0.2
$\alpha_e$	3

#### 315 4 SBL Regimes

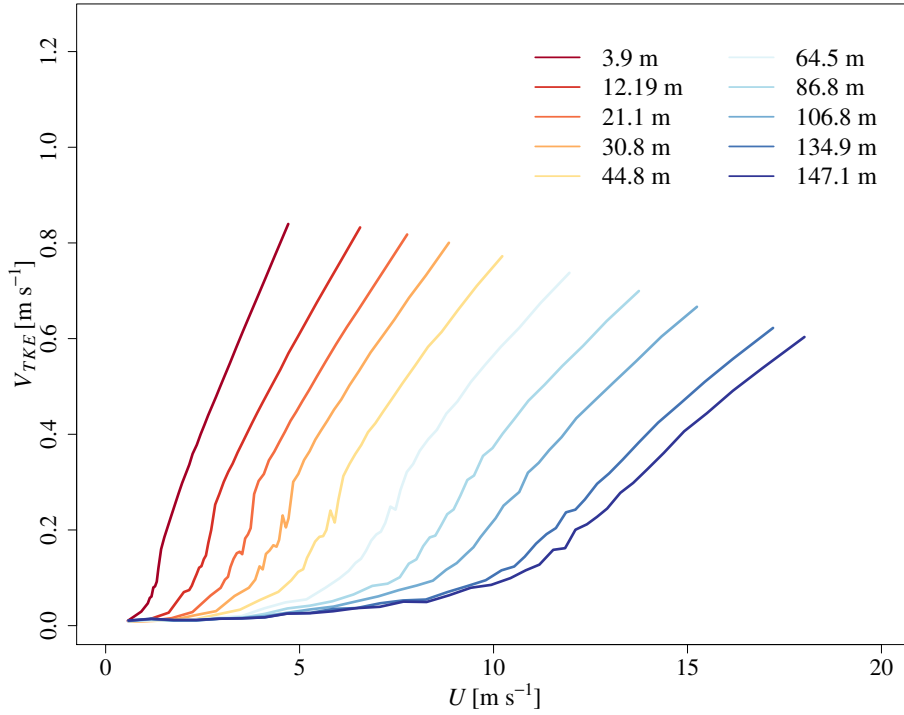
316 The methodology employed to evaluate whether the newly proposed PBL  
 317 scheme correctly reproduces the physical properties of both SBL turbulence  
 318 regimes is similar to that employed by Maroneze et al. (2021). The geostrophic  
 319 zonal velocity component  $u_G$  is assumed to be temporally and vertically con-  
 320 stant for each simulation. Values of  $u_G$  considered varied from  $0.5 \text{ m s}^{-1}$  to  
 321  $24 \text{ m s}^{-1}$ , with  $0.5 \text{ m s}^{-1}$  steps between  $0.5$  and  $14 \text{ m s}^{-1}$ , and with  $1 \text{ m}$   
 322  $\text{s}^{-1}$  steps between  $14$  and  $24 \text{ m s}^{-1}$ . In this section the skin temperature is  
 323 estimated through the Unified Noah Land Surface Model (Mukul Tewari et al.  
 324 2004), while the RRTM Longwave Scheme (Mlawer et al. 1997) and the Dud-  
 325 hia Shortwave Scheme (Dudhia 1989) are adopted for radiative processes. The  
 326 Purdue Lin scheme (Chen and Sun 2002) has been used to represent the mi-  
 327 crophysical process. For all simulations one dryland cropland and pasture land  
 328 vegetation with vegetation fraction of  $0.5$  are considered. The roughness length  
 329 for this surface type is  $0.15 \text{ m}$ .

330 The SBL turbulence regime transition can be determined from both lo-  
 331 cal and bulk variables. Initially, both SBL regime classifications and regime  
 332 transitions have been marked in terms of stability parameters (Mahrt 1998).  
 333 However, the use of a single stability parameter such as the Obukhov length,

334 or the Richardson number (either in its flux or gradient form), is ineffective  
 335 at distinguishing the SBL regimes universally (Monahan et al. 2015). On the  
 336 other hand, since the mean wind speed has a crucial control on the SBL regime,  
 337 it has become common practice to assess it through the relationship among  
 338 turbulent quantities, such as the turbulence velocity scale ( $V_{TKE} = \sqrt{\bar{e}}$ ), and  
 339 the mean wind speed (Sun et al. 2012; van de Wiel et al. 2012; Acevedo et al.  
 340 2016, 2019, among others).

341 The study of Sun et al. (2012) established the relationship between  $V_{TKE}$   
 342 and  $\bar{U}$  as a criterion to determine the local SBL regime (Acevedo et al. 2016,  
 343 *incluir varios*). In general,  $V_{TKE}$  increases linearly with  $\bar{U}$  in both regimes,  
 344 but the rate of increase is notably larger in the weakly stable than in the  
 345 very stable regime. Therefore, the regime transition occurs at the value of  $\bar{U}$   
 346 for which the slope of such a relationship changes abruptly (Sun et al. 2012).  
 347 The scheme presently proposed simulates a very stable regime, characterized  
 348 by a subtle  $V_{TKE}$  increase as  $\bar{U}$  increases, in general agreement with the ob-  
 349 servations. It is important to notice that for very weak winds speeds,  $V_{TKE}$   
 350 sometimes assumes a constant value that is equal to minimum value imposed  
 351 to the scheme. Contrastingly, when the wind speed is larger, a weakly stable  
 352 regime is simulated, with  $V_{TKE}$  increasing rapid as  $\bar{U}$  increases (Fig. 7). The  
 353 change between the two slopes is abrupt, characterizing the value of  $\bar{U}$  for which  
 354 the SBL regime transition occurs (Fig. 7). Maroneze et al. (2021) compared  
 355 the relationship between  $V_{TKE}$  and  $\bar{U}$  for six different boundary layer schemes  
 356 that explicitly or implicitly solve TKE in WRF: Mellor-Yamada Nakanishi Ni-  
 357 ino 2.5 (MYNN2.5), Mellor-Yamada Nakanishi Niino 3.0 (MYNN3.0), Mellor  
 358 Yamada-Janjic (MYJ) Quasi-Normal Scale Elimination (QNSE), University of  
 359 Washington (UWBLS) and Bougeault-Lacarrère (BouLac). Among these, only  
 360 MYNN2.5 was able to reproduce both SBL regimes and the transition between  
 361 them as according to the observations as does the scheme presently proposed.  
 362 The others typically simulate poorly the very-stable regime, as  $V_{TKE}$  varies  
 363 little of nothing as  $\bar{U}$  increases.

364 The SBL regime transition is also clear when the temperature difference  
 365 near the ground is considered (Fig. 8). As shown by Vignon et al. (2017), the  
 366 near surface potential temperature difference reaches its larger values for weak  
 367 wind speeds, when turbulence is not strong enough to mix the whole layer,  
 368 so that the lower levels become decoupled from higher levels. As the mean  
 369 wind speed increases, the potential temperature difference starts to decrease  
 370 sharply characterizing the regime transition (Fig. 8). Moreover, for large wind  
 371 speeds the temperature difference tends to very small values, because the tur-  
 372 bulence is sufficient to keep the bottom and top of the layer coupled. Maroneze  
 373 et al. (2021) have shown that the relationship between the vertical tempera-  
 374 ture difference and the mean wind speed is not well solved by the many turbu-  
 375 lence schemes used in WRF. Besides, the solutions of the different schemes are  
 376 largely variable in terms of the thermal gradient they simulate in the limits of  
 377 small and large wind speed, or in terms of how abrupt is the variation of such  
 378 a gradient with wind speed at the regime transition. In the present scheme, a  
 379 temperature difference of 3.5 K between 1 and 30 m is simulated under low

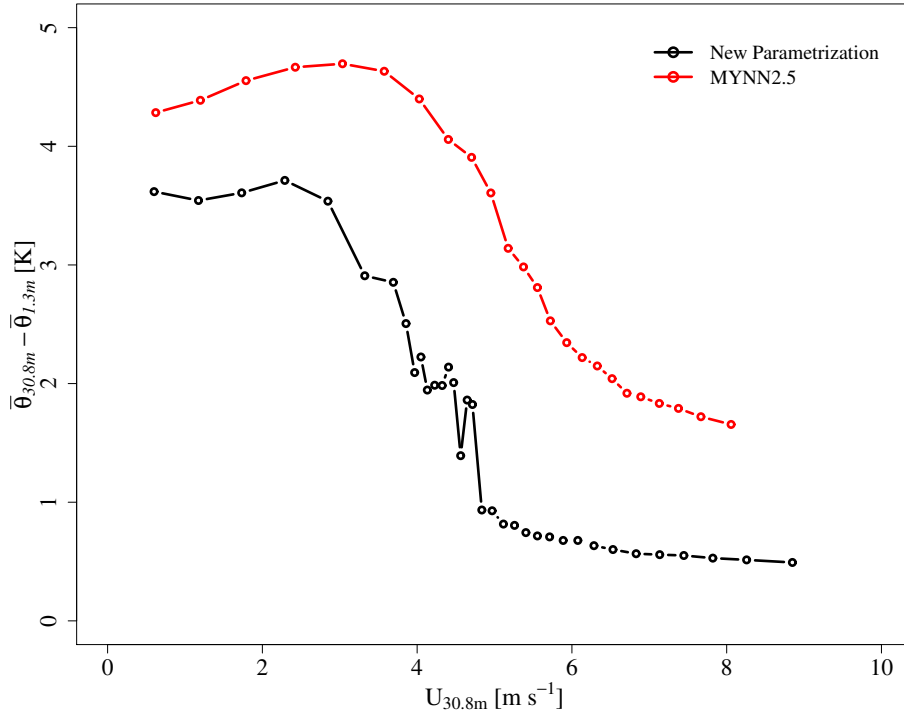


**Fig. 7** Average turbulence velocity scale  $V_{TKE}$  as a function of the mean wind speed  $U$ . The panel shows bin-averaged values taken from the all night of simulation, considering all runs.

380 winds, dropping below 1 K as the 30-m wind speed exceeds  $5 \text{ m s}^{-1}$ . In general,  
 381 the new formulation leads to smaller thermal gradients than MYNN2.5,  
 382 and a more abrupt transition. In Fig. 8, these results are compared to those  
 383 obtained with MYNN2.5, that was found to best reproduce the regime transi-  
 384 tion among those schemes compared by Maroneze et al. (2021). Particularly,  
 385 the reduced thermal gradients in the weakly stable regime are confirmed by  
 386 CASES-99 observations (Sun et al. 2012; Acevedo et al. 2021). The difference  
 387 is partially a consequence that the present scheme lacks a stability function,  
 388 so that the dependence on stratification arises naturally from the heat flux  
 389 prognostic equation.

390 The turbulent Prandtl number is defined as the ratio between the momen-  
 391 tum eddy diffusivity and the heat diffusivity ( $Pr_T = K_M/K_H$ ). Generally,  
 392 PBL schemes use a constant value for  $Pr_T$ , thus estimating the turbulent  
 393 heat diffusivity ( $K_H$ ). Such a use of a constant  $Pr_T$  implies that the ratio  
 394 between the flux ( $Ri_f$ ) and gradient ( $Ri_g$ ) Richardson numbers is also con-  
 395 stant, since  $Pr_T = Ri_g/Ri_f$ . However, both atmospheric data and laboratory  
 396 experiments show that only in the weakly stable regime  $Ri_f$  increases lin-  
 397 early with  $Ri_g$ , implying a constant  $Pr_T$  (Zilitinkevich et al. 2013), whose  
 398 value may be from 0.7 to 0.9 (Basu and Holtslag 2021). On the other hand,

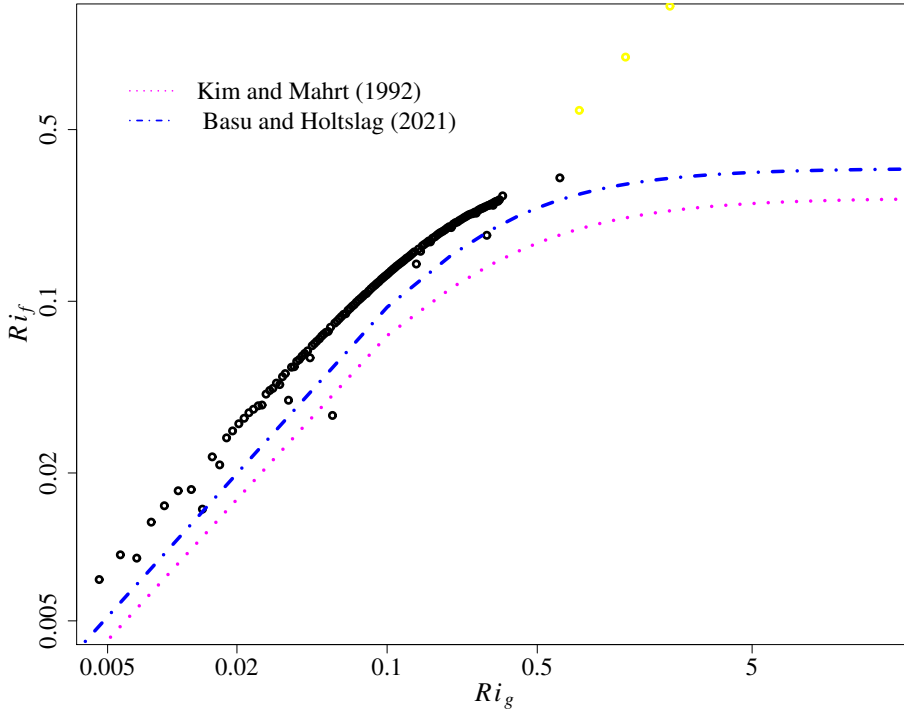




**Fig. 8** Average potential temperature difference between 30.8 and 1.3 m as a function of the 30.8-m wind speed for each PBL parametrization, according to the legend. The panel shows bin-averaged values taken from the all night of simulation, considering all runs.

399 observations show that when  $Ri_g$  exceeds a critical value ( $Ri_g \approx 0.2$ ),  $Ri_f$   
400 tends to a finite asymptotic value, characterizing the very stable regime (Zil-  
401 itinkevich et al. 2013; Bou-Zeid et al. 2018; Basu and Holtslag 2021, among  
402 other). Furthermore, under very stable stratification it is possible that wave  
403 phenomena increase the momentum diffusivity but not the heat diffusivity  
404 (Grachev et al. 2007). In such situation, generally  $Pr_T > 1$ . In a model that  
405 uses stability functions, the flux and gradient Richardson numbers are related  
406 as  $Ri_f = f_h/f_m Ri_g$ , where  $f_h$  is the stability function for heat. Therefore,  
407 the maximum flux Richardson number is limited by the imposed ratio  $f_h/f_m$ .  
408 As  $Pr_T$  is an important model tuning parameter (Sukoriansky et al. 2005) its  
409 values can affect the model performance.

410 In the presently proposed scheme, where a prognostic equation for heat  
411 flux is solved, the turbulent Prandtl number is not imposed and its value  
412 arises naturally from the solved ratio between  $Ri_f$  and  $Ri_g$  (Fig. 9). Fig. 9  
413 shows that it is able to reproduce the relationship between  $Ri_f$  and  $Ri_g$  for  
414 both SBL regimes. In the weakly stable regime,  $Pr_T$  is approximately constant  
415 for  $Ri_g < 0.2$ , while the turbulent Prandtl number increases linearly with  
416  $Ri_g$  if  $Ri_g > 0.2$ . A very similar pattern is presented by both observational



**Fig. 9** Average flux Richardson number as a function of the gradient Richardson number, showing bin-averaged values taken from the all night of simulation, considering all runs

417 and model studies (Kim and Mahrt 1992; Zilitinkevich et al. 2013; Basu and  
 418 Holtslag 2021) (Fig. 9). In very stable conditions,  $Ri_g > 0.5$ , the model results  
 419 are limited by the imposed minimum TKE value assumed equal  $0.00001 \text{ m}^2$   
 420  $\text{s}^{-2}$  (yellow circles Fig 9).

## 421 5 Conclusion

422 A new boundary layer parameterization, in which the turbulent heat flux and  
 423 the temperature variance are solved through a prognostic equation, has been  
 424 introduced. The new scheme has been validated through simulations using  
 425 the single-column Weather Research and Forecasting (WRF) model with the  
 426 GABLS I experiment as a case-control.

427 A more realistic simulation of the nocturnal turbulence regimes of very  
 428 stable and weakly stable boundary layers was the main motivation behind  
 429 the idea of introducing a scheme that prognostically solves the heat flux and  
 430 temperature variance. This goal has been achieved to a large degree, as shown  
 431 by the comparison between mean and turbulent quantities presented in Section  
 432 4.

433 The present parameterization does not assume any prescribed stability  
434 function, differently from others SBL schemes. Here, the stratification depen-  
435 dence on other characteristics of the mean and turbulent flows arises natu-  
436 rally from the use of a prognostic equation for the heat flux . This way, the  
437 present scheme is able to realistically simulate the large temperature gradient  
438 that often occurs when mean wind speeds are very small. On the other hand,  
439 for large wind speeds, the temperature gradient is nearly destroyed, so that  
440 the model simulates a near-neutral SBL. Besides, the new formulation leads  
441 to smaller thermal gradients and a more abrupt transition between the SBL  
442 regimes than occurs in the Mellor-Yamada Nakanishi Niino 2.5 scheme, which  
443 was found by Maroneze et al. (2021) to best simulate the differences between  
444 the two regimes among the WRF schemes that solve TKE prognostically. This  
445 is also in better agreement with CASES-99 observations. Along the same line,  
446 SBL schemes typically use a fixed and prescribed turbulent Prandtl number,  
447 which may cause a excessive turbulent heat diffusion under very stable condi-  
448 tions (Grachev et al. 2007; Zilitinkevich et al. 2013; Maroneze et al. 2021). In  
449 the present parametrization, the turbulent Prandtl number is not prescribed,  
450 being directly calculated from the ratio between  $Ri_f$  and  $Ri_g$ .

451 The future steps of this research include the specific development and val-  
452 idation of a parametrization for the convective boundary layer, the inclusion  
453 of a more sophisticated parametrization for  $\overline{w'^2}$  and the inclusion of humidity  
454 effects in the parameterization. Furthermore, three-dimensional model simula-  
455 tion of real cases, where advection and other external processes affect the PBL  
456 are necessary to validate the parameterization in a more realistic scenario, so  
457 that it can ultimately be implemented in operational models.

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